



# Large Extra Dimensions at IceCube and DUNE

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# Neutrinos in the SM

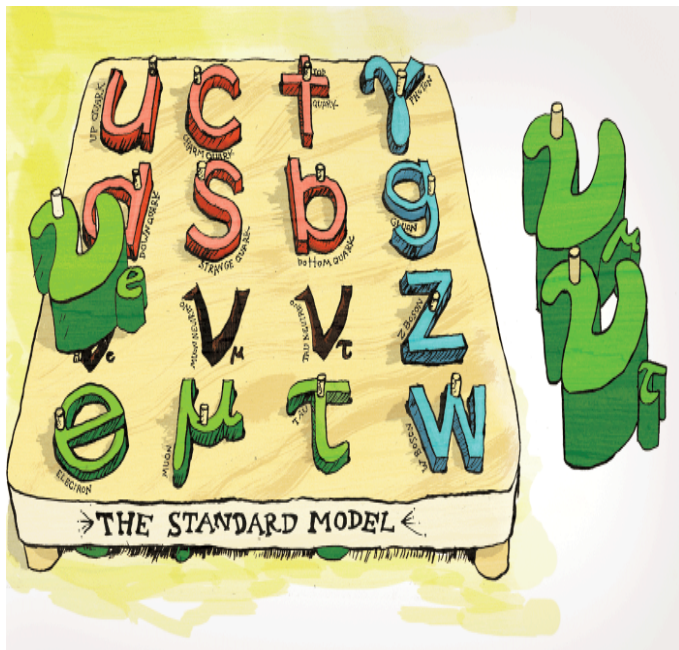
$$\begin{aligned} \mathcal{L} = & \sum_{\alpha=e,\mu,\tau} i\bar{\nu}_L^\alpha \not{\partial} \nu_L^\alpha \\ & - \frac{g}{2\sqrt{2}} \sum_{\alpha=e,\mu,\tau} \bar{\nu}_L^\alpha \gamma^\rho (1 - \gamma^5) l_\alpha W_\rho + h.c. \\ & - \frac{g}{4 \cos \theta_w} \sum_{\alpha=e,\mu,\tau} \bar{\nu}_L^\alpha \gamma^\rho (1 - \gamma^5) \nu_L^\alpha Z_\rho \\ & - \sum_{\alpha=e,\mu,\tau} \bar{l}_L^\alpha M^l l_R^\alpha + h.c. \end{aligned}$$

(neutrino Kinetic term)

(Charged Current Interaction)

(Neutral Current Interaction)

(leptonic mass term)



$$\mathbf{L}_e = \begin{pmatrix} \nu_{eL} \\ e_L \end{pmatrix}$$

$$\mathbf{L}_\mu = \begin{pmatrix} \nu_{\mu L} \\ \mu_L \end{pmatrix}$$

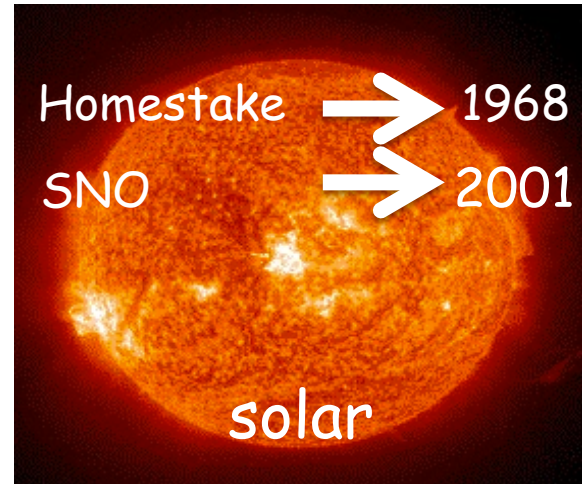
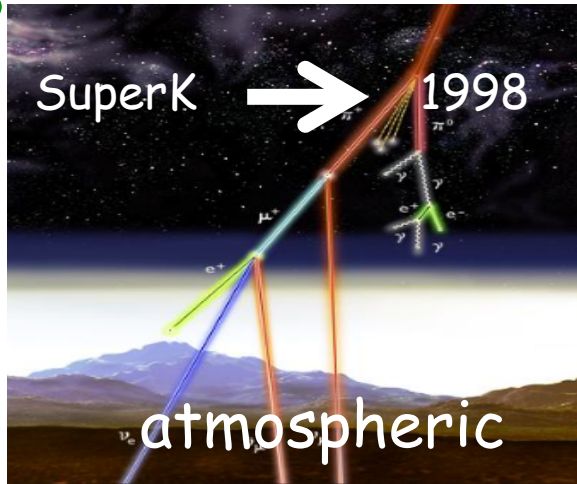
$$\mathbf{L}_\tau = \begin{pmatrix} \nu_{\tau L} \\ \tau_L \end{pmatrix}$$

No right handed neutrinos in the SM,  
hence no mass!!!

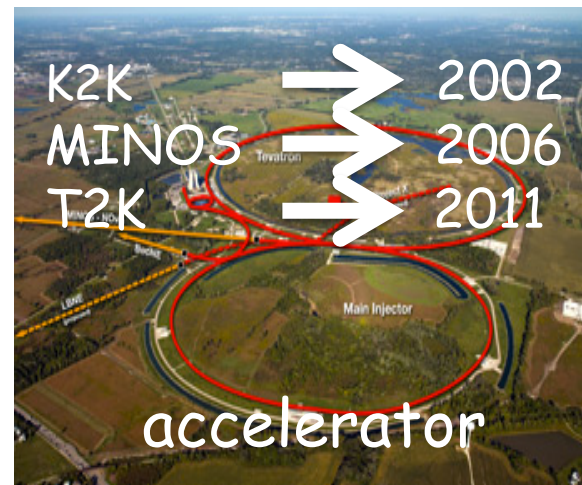
# Neutrinos have no mass in the SM!

However in nature....

Surprise!!!



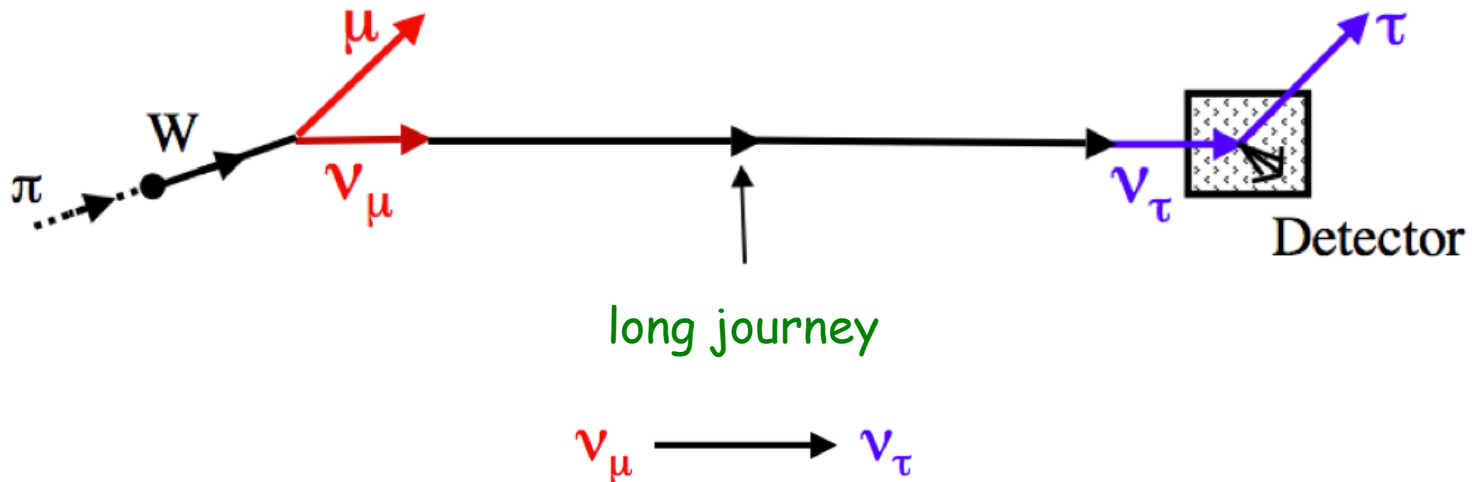
Neutrino oscillation needs masses and mixing!



# The oscillation is described by the "PMNS" matrix!

$$\mathcal{L}_{CC} = -\frac{g}{\sqrt{2}} \bar{\ell}_L^i \gamma^\mu \boxed{U_{PMNS}} \nu_j W_\mu^+ + h.c.$$

If neutrinos have mass, and leptons mix, we can have neutrino oscillation:



Neutrino experiments performed in the last 2 decades have proved that such flavor changes actually occur!

# Two Flavors

$$U = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix}$$

One mixing angle  
No CP phase

$$\Delta m^2 = \Delta m_{21}^2 = m_2^2 - m_1^2$$

One mass-squared  
difference

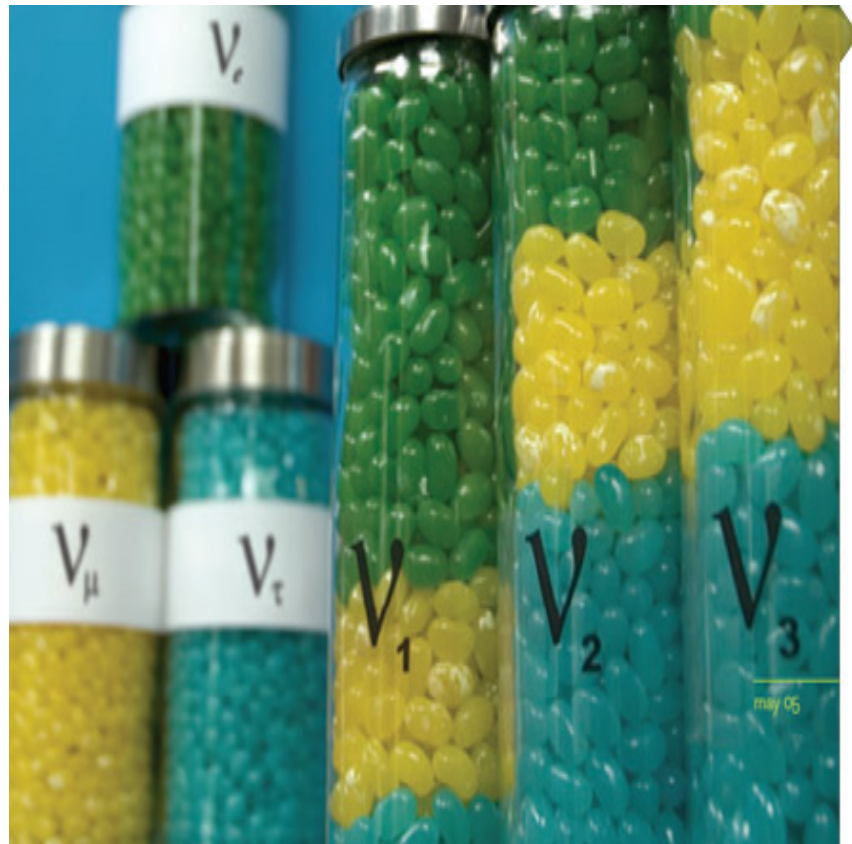
$$\begin{aligned} P(\nu_\alpha \rightarrow \nu_\beta) &= \frac{1}{2} \sin^2 2\theta \left[ 1 - \cos \left( \frac{\Delta m^2 L}{2E} \right) \right] \\ &= \sin^2 2\theta \sin^2 \left( \frac{\Delta m^2 L}{4E} \right) \end{aligned}$$

Transition  
probability

$$P(\nu_\alpha \rightarrow \nu_\alpha) = 1 - \sin^2 2\theta \sin^2 \left( \frac{\Delta m^2 L}{4E} \right)$$

Survival  
probability

# Standard picture of neutrinos

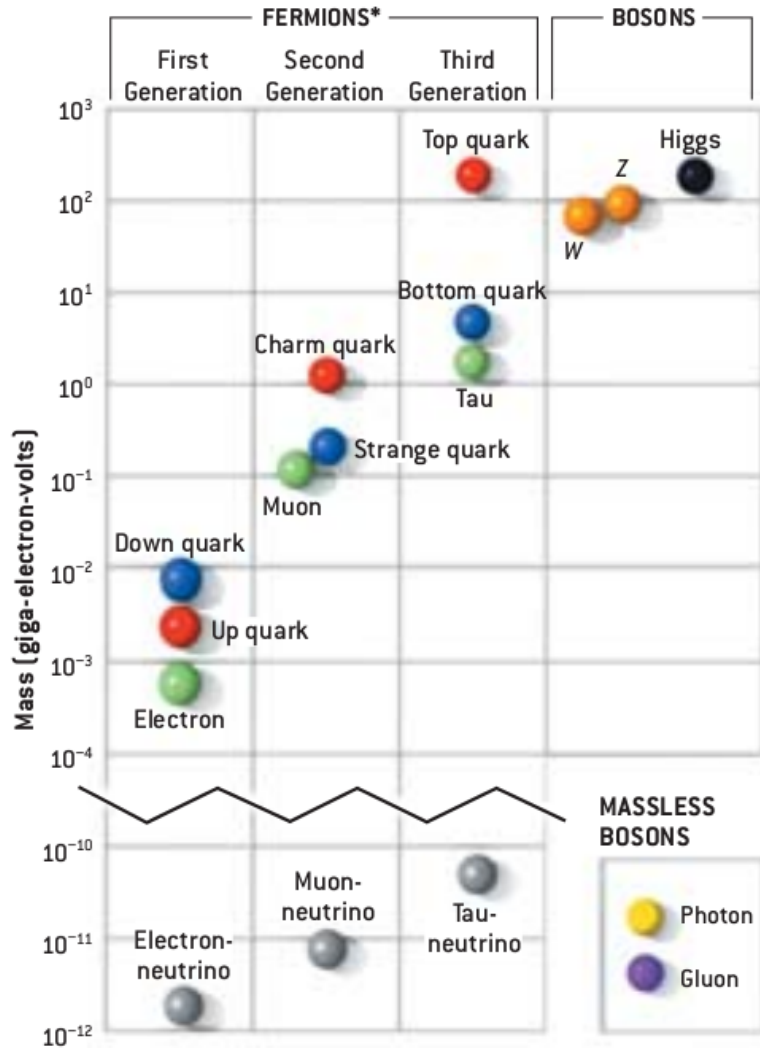


$$\nu_\alpha = \sum_{i=1}^3 U_{\alpha i} \nu_i$$

$$\alpha = e, \mu, \tau$$

$$U_{\text{PMNS}} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix}$$

# Smallness of Neutrino Masses



How can we explain it?!!

# LED: One solution to the hierarchy problem

[Arkani-Hamed, Dimopoulos and Dvali, 1998]

Idea: Graviton propagates in the extra dimension

$$4D \rightarrow (4 + n)D$$

$$\Lambda_{Pl} \rightarrow M \sim m_{ew}$$

✓ "n" compactified extra space dimensions with size R.

✓ Gravity in all "3+n" space dimensions

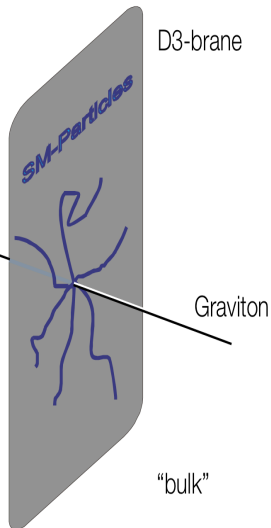
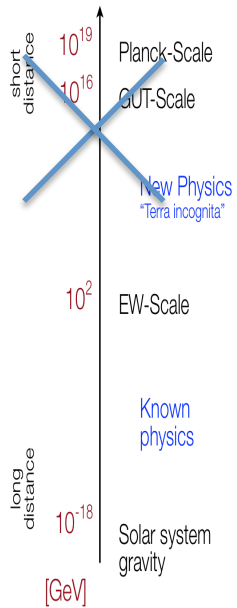
✓ SM interactions and all matter particles (including  $\nu_L$ ) have to be confined to our 4-dimensional world.

$$M_{Pl}^2 = V_n M_f^{2+n}$$

$$V_n = (2\pi)^n R^n$$

✓ At least 2 extra dimensions are needed:

$$n \geq 2$$





# Kaluza-Klein Theory

- Extra dimensions should be compact. Simplest case is periodic with radius  $R$ .

$$\phi(x^\mu, y) = \phi^0(x^\mu) + \sum_{n=1}^{\infty} \phi^{(n)}(x^\mu) e^{i \frac{n}{R} y} + h.c.$$

- Consider the 5-dimensional Klein-Gordon eq.:

$$\begin{aligned} \partial_M \partial^M \phi &= 0, \\ \partial_\mu \partial^\mu \phi - \partial_y \partial^y \phi &= 0, \\ (\partial_\mu \partial^\mu + (n/R)^2) \phi^{(n)} &= 0. \end{aligned}$$



- This is a tower of states with:  $m_n^2 = (n/R)^2$ .

# Formalism

If a SM singlet fermion (such as the sterile neutrino) is present, it can also propagate to more than 4-dimensions.

[Arkani-Hamed, Dimopoulos and Dvali, Russell 1998]

$$S = \int d^4x dy i\bar{\Psi}^\alpha \Gamma^A \partial_A \Psi_\alpha + \int d^4x (i\bar{\nu}_{\alpha L} \bar{\sigma}^\mu \partial_\mu \nu_L^\alpha + \lambda_{\alpha\beta} H \bar{\nu}_L^\alpha \psi_R^\beta(x, 0) + h.c.)$$

The 5-dimensional fermion

The SM neutrinos

After compactification:

$$\begin{aligned}\nu_R^{\alpha(0)} &\equiv \psi_R^{\alpha(0)}, \\ \nu_R^{\alpha(n)} &= \frac{\psi_R^{\alpha(n)} + \psi_R^{\alpha(n)}}{\sqrt{2}}, \quad n = 1, \dots, \infty, \\ \nu_L^{\alpha(n)} &= \frac{\psi_L^{\alpha(n)} + \psi_L^{\alpha(n)}}{\sqrt{2}}, \quad n = 1, \dots, \infty,\end{aligned}$$

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The

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Flavor neutrinos



$$i \frac{d}{dt} \begin{pmatrix} \nu_1^{(0)} \\ \nu_2^{(0)} \\ \nu_3^{(0)} \\ \nu_1^{(1)} \\ \nu_2^{(1)} \\ \nu_3^{(1)} \\ \nu_1^{(2)} \\ \nu_2^{(2)} \\ \nu_3^{(2)} \\ \vdots \\ \nu_1^{(N)} \\ \nu_2^{(N)} \\ \nu_3^{(N)} \end{pmatrix} = \frac{1}{2ER_{\text{ED}}^2} \begin{pmatrix} \eta_1 + V_{11} & V_{12} & V_{13} & \xi_1 & 0 & 0 & 2\xi_1 & 0 & 0 & \dots & N\xi_1 & 0 & 0 \\ V_{21} & \eta_2 + V_{22} & V_{23} & 0 & \xi_2 & 0 & 0 & 2\xi_2 & 0 & \dots & 0 & N\xi_2 & 0 \\ V_{31} & V_{32} & \eta_3 + V_{33} & 0 & 0 & \xi_3 & 0 & 0 & 2\xi_3 & \dots & 0 & 0 & N\xi_3 \\ \hline \xi_1 & 0 & 0 & 1 & 0 & 0 & 0 & 0 & 0 & \dots & 0 & 0 & 0 \\ 0 & \xi_2 & 0 & 0 & 1 & 0 & 0 & 0 & 0 & \dots & 0 & 0 & 0 \\ 0 & 0 & \xi_3 & 0 & 0 & 1 & 0 & 0 & 0 & \dots & 0 & 0 & 0 \\ \hline 2\xi_1 & 0 & 0 & 0 & 0 & 0 & 4 & 0 & 0 & \dots & 0 & 0 & 0 \\ 0 & 2\xi_2 & 0 & 0 & 0 & 0 & 0 & 4 & 0 & \dots & 0 & 0 & 0 \\ 0 & 0 & 2\xi_3 & 0 & 0 & 0 & 0 & 0 & 4 & \dots & 0 & 0 & 0 \\ \hline \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \ddots & \vdots & \vdots & \vdots \\ \hline N\xi_1 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & \dots & N^2 & 0 & 0 \\ 0 & N\xi_2 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & \dots & 0 & N^2 & 0 \\ 0 & 0 & N\xi_3 & 0 & 0 & 0 & 0 & 0 & 0 & \dots & 0 & 0 & N^2 \end{pmatrix} \begin{pmatrix} \nu_1^{(0)} \\ \nu_2^{(0)} \\ \nu_3^{(0)} \\ \nu_1^{(1)} \\ \nu_2^{(1)} \\ \nu_3^{(1)} \\ \nu_1^{(2)} \\ \nu_2^{(2)} \\ \nu_3^{(2)} \\ \vdots \\ \nu_1^{(N)} \\ \nu_2^{(N)} \\ \nu_3^{(N)} \end{pmatrix}$$

$$\eta_i = (N + 1/2) \xi_i^2, \quad V_{ij} = 2ER_{\text{ED}}^2 \sum_{\alpha=e,\mu,\tau} U_{\alpha i}^* U_{\alpha j} V_\alpha$$

$$\xi_i = \sqrt{2m_i^D} R_{\text{ED}}$$

$$i \frac{d}{dt} \begin{pmatrix} \nu_1^{(0)} \\ \nu_2^{(0)} \\ \nu_3^{(0)} \\ \nu_1^{(1)} \\ \nu_2^{(1)} \\ \nu_3^{(1)} \\ \nu_1^{(2)} \\ \nu_2^{(2)} \\ \nu_3^{(2)} \\ \vdots \\ \nu_1^{(N)} \\ \nu_2^{(N)} \\ \nu_3^{(N)} \end{pmatrix} = \frac{1}{2ER_{\text{ED}}^2} \begin{pmatrix} \eta_1 + V_{11} & V_{12} & V_{13} & \xi_1 & 0 & 0 & 2\xi_1 & 0 & 0 & \dots & N\xi_1 & 0 & 0 \\ V_{21} & \eta_2 + V_{22} & V_{23} & 0 & \xi_2 & 0 & 0 & 2\xi_2 & 0 & \dots & 0 & N\xi_2 & 0 \\ V_{31} & V_{32} & \eta_3 + V_{33} & 0 & 0 & \xi_3 & 0 & 0 & 2\xi_3 & \dots & 0 & 0 & N\xi_3 \\ \hline \xi_1 & 0 & 0 & 1 & 0 & 0 & 0 & 0 & 0 & \dots & 0 & 0 & 0 \\ 0 & \xi_2 & 0 & 0 & 1 & 0 & 0 & 0 & 0 & \dots & 0 & 0 & 0 \\ 0 & 0 & \xi_3 & 0 & 0 & 1 & 0 & 0 & 0 & \dots & 0 & 0 & 0 \\ \hline 2\xi_1 & 0 & 0 & 0 & 0 & 0 & 4 & 0 & 0 & \dots & 0 & 0 & 0 \\ 0 & 2\xi_2 & 0 & 0 & 0 & 0 & 0 & 4 & 0 & \dots & 0 & 0 & 0 \\ 0 & 0 & 2\xi_3 & 0 & 0 & 0 & 0 & 0 & 4 & \dots & 0 & 0 & 0 \\ \hline \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \ddots & \vdots & \vdots & \vdots \\ \hline N\xi_1 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & \dots & N^2 & 0 & 0 \\ 0 & N\xi_2 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & \dots & 0 & N^2 & 0 \\ 0 & 0 & N\xi_3 & 0 & 0 & 0 & 0 & 0 & 0 & \dots & 0 & 0 & N^2 \end{pmatrix} \begin{pmatrix} \nu_1^{(0)} \\ \nu_2^{(0)} \\ \nu_3^{(0)} \\ \nu_1^{(1)} \\ \nu_2^{(1)} \\ \nu_3^{(1)} \\ \nu_1^{(2)} \\ \nu_2^{(2)} \\ \nu_3^{(2)} \\ \vdots \\ \nu_1^{(N)} \\ \nu_2^{(N)} \\ \nu_3^{(N)} \end{pmatrix}$$

$$\eta_i = (N + 1/2) \xi_i^2, \quad V_{ij} = 2ER_{\text{ED}}^2 \sum_{\alpha=e,\mu,\tau} U_{\alpha i}^* U_{\alpha j} V_{\alpha}$$

$$\xi_i = \sqrt{2m_i^D} R_{\text{ED}}$$

The free parameters are "m<sub>i</sub>" and "R"

# The equivalence between the LED and (3+3n) sterile models

The relation between flavor and mass eigenstates:

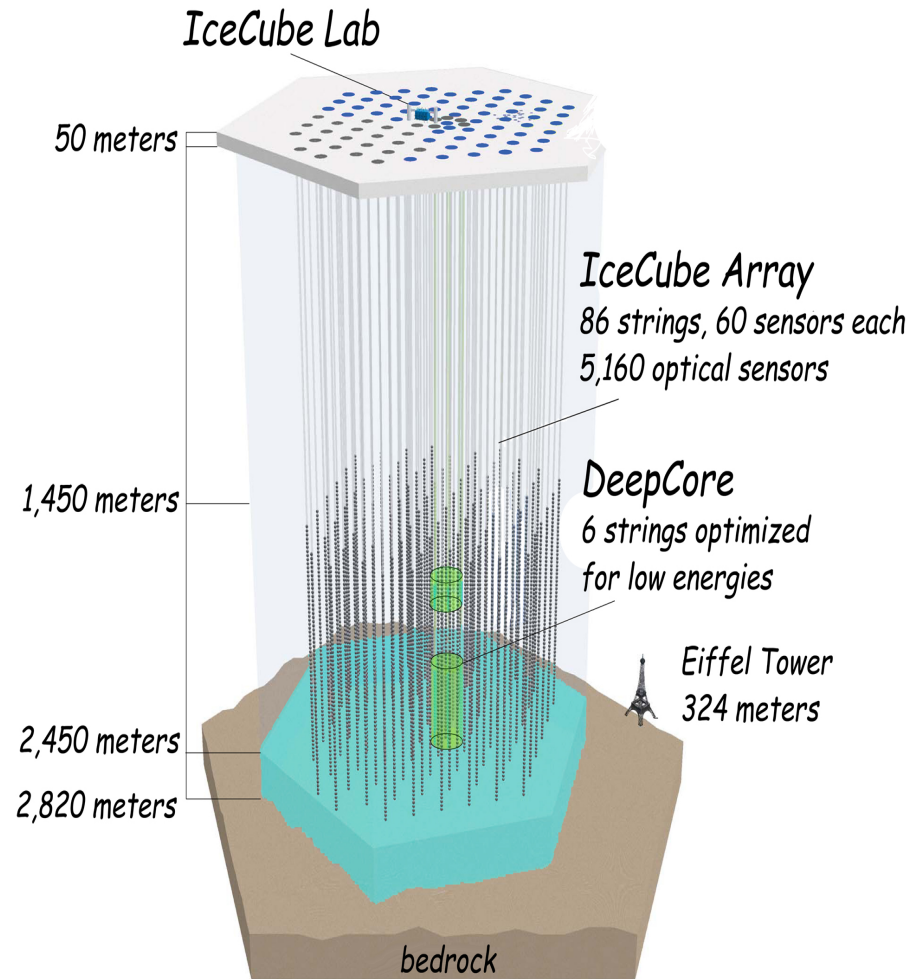
$$\nu_{\alpha L} = \sum_{i=1}^3 U_{\alpha i} \sum_{n=0}^N S_i^{0n} \nu_{iL}'^{(n)} = \sum_{i=1}^3 \sum_{n=0}^N W_{\alpha(i+3n)} \nu_{iL}'^{(n)}, \quad (\alpha = e, \mu, \tau)$$

In (3+3n) model:

$$W_{\alpha(i+3n)} = U_{\alpha i} S_i^{0n}, \quad (i = 1, 2, 3), \quad (\alpha = e, \mu, \tau), \quad (n = 0, 1, \dots, N)$$

$$P(\nu_{\alpha} \rightarrow \nu_{\beta}) = \delta_{\alpha\beta} - 4 \sum_{l>m} \Re[W_{\alpha l} W_{\beta l}^* W_{\alpha m}^* W_{\beta m}] \sin^2 \left( \frac{\Delta m_{lm}^2 L}{4E_{\nu}} \right) + 2 \sum_{l>m} \Im[W_{\alpha l} W_{\beta l}^* W_{\alpha m}^* W_{\beta m}] \sin \left( \frac{\Delta m_{lm}^2 L}{2E_{\nu}} \right),$$

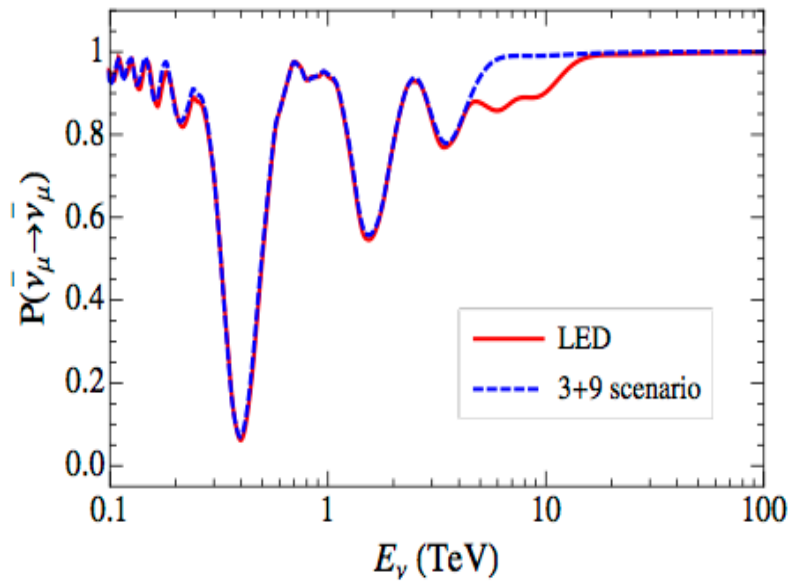
- ▶ 5160 PMTs
- ▶ 1 km<sup>3</sup> volume
- ▶ 86 strings
- ▶ 17 m PMT-PMT spacing per string
- ▶ 120 m string spacing
- ▶ Angular resolution  $\sim 1^\circ$
- ▶ Completed 2010



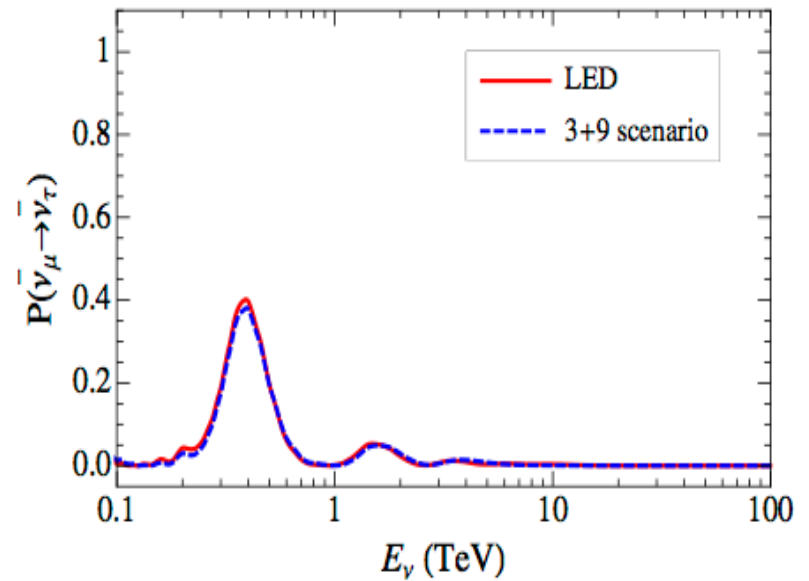
# The Oscillation Probabilities

- For 3 KK modes:

$$m_1^D = 0.01 \text{ eV and } R_{\text{ED}} = 5 \times 10^{-5} \text{ cm}$$

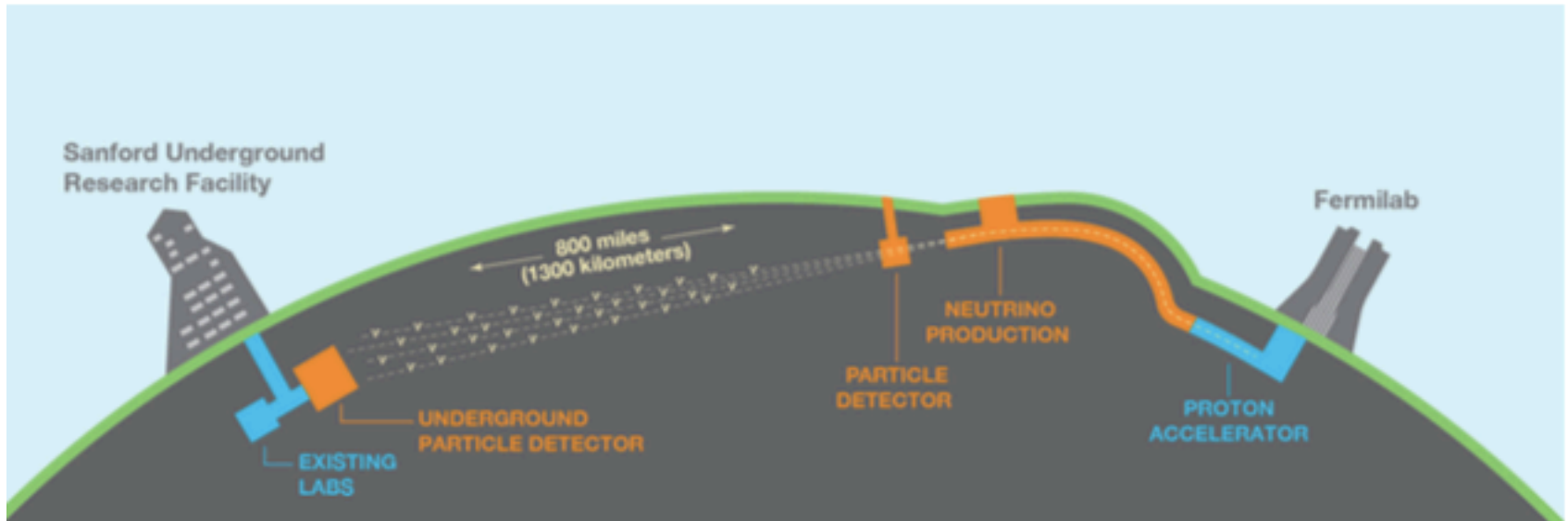


(a)  $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu)$



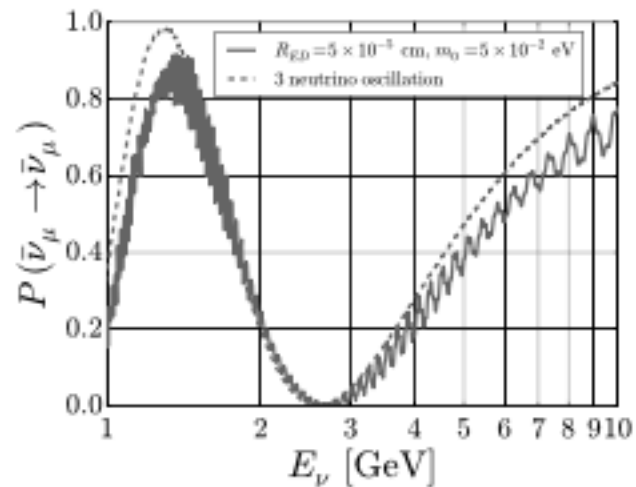
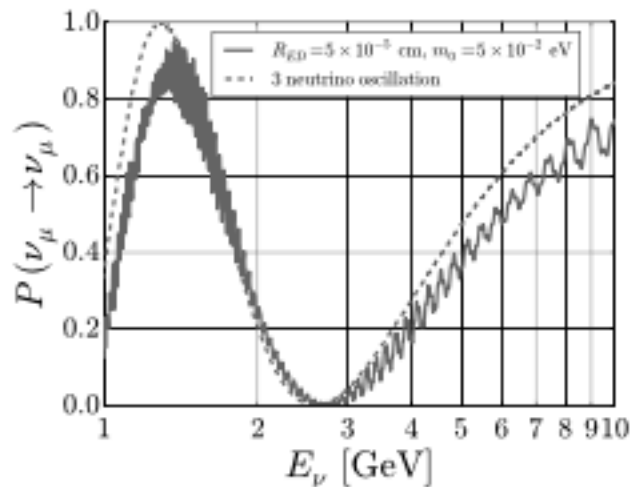
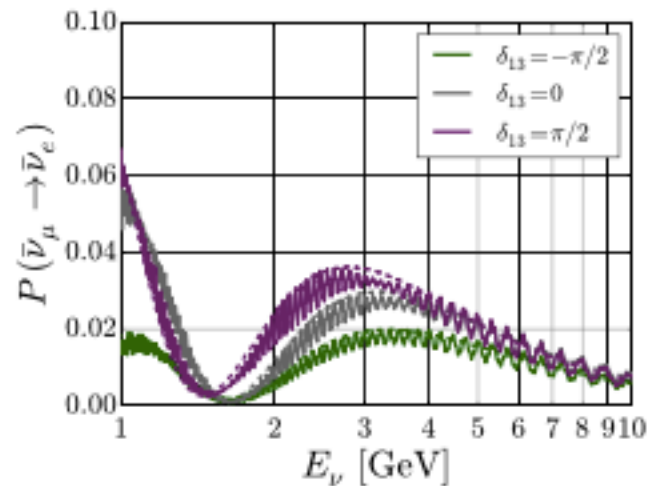
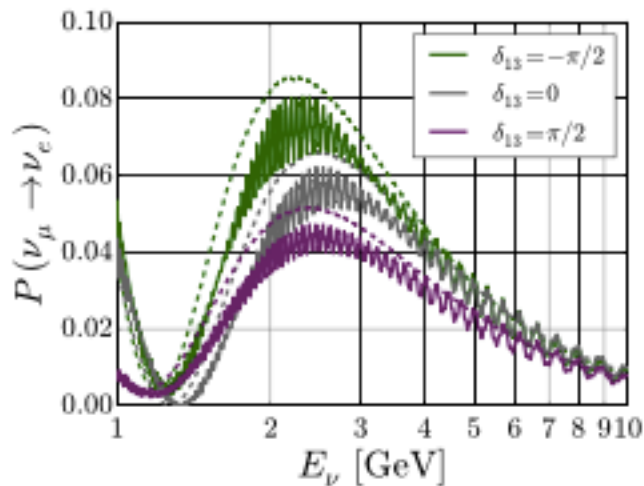
(b)  $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_\tau)$





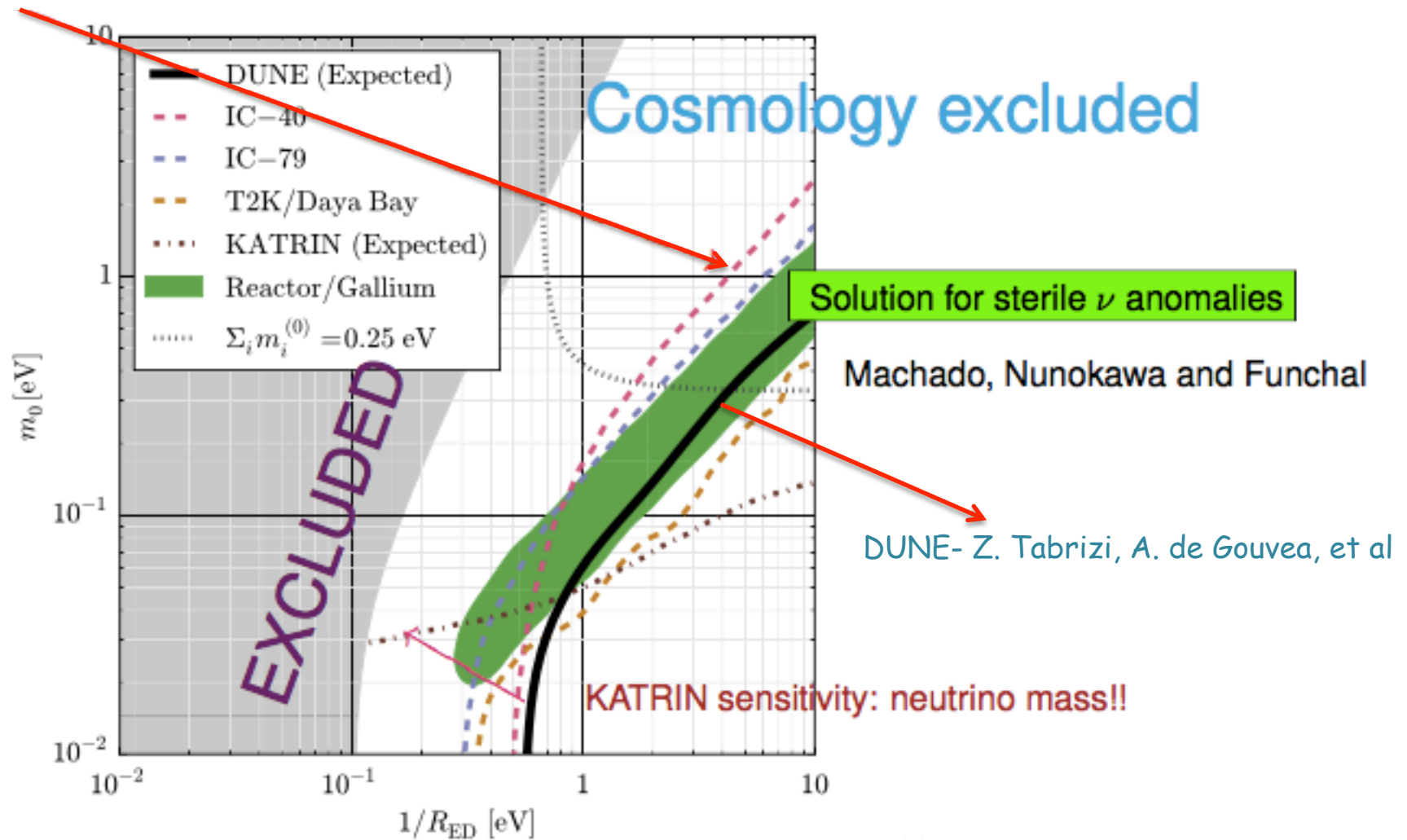
Measure  $\mu$  ( $e$ ) events: Test of  $P(\nu_\mu \rightarrow \nu_\mu)$  ( $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu)$ ) and  $P(\nu_\mu \rightarrow \nu_e)$  ( $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$ )

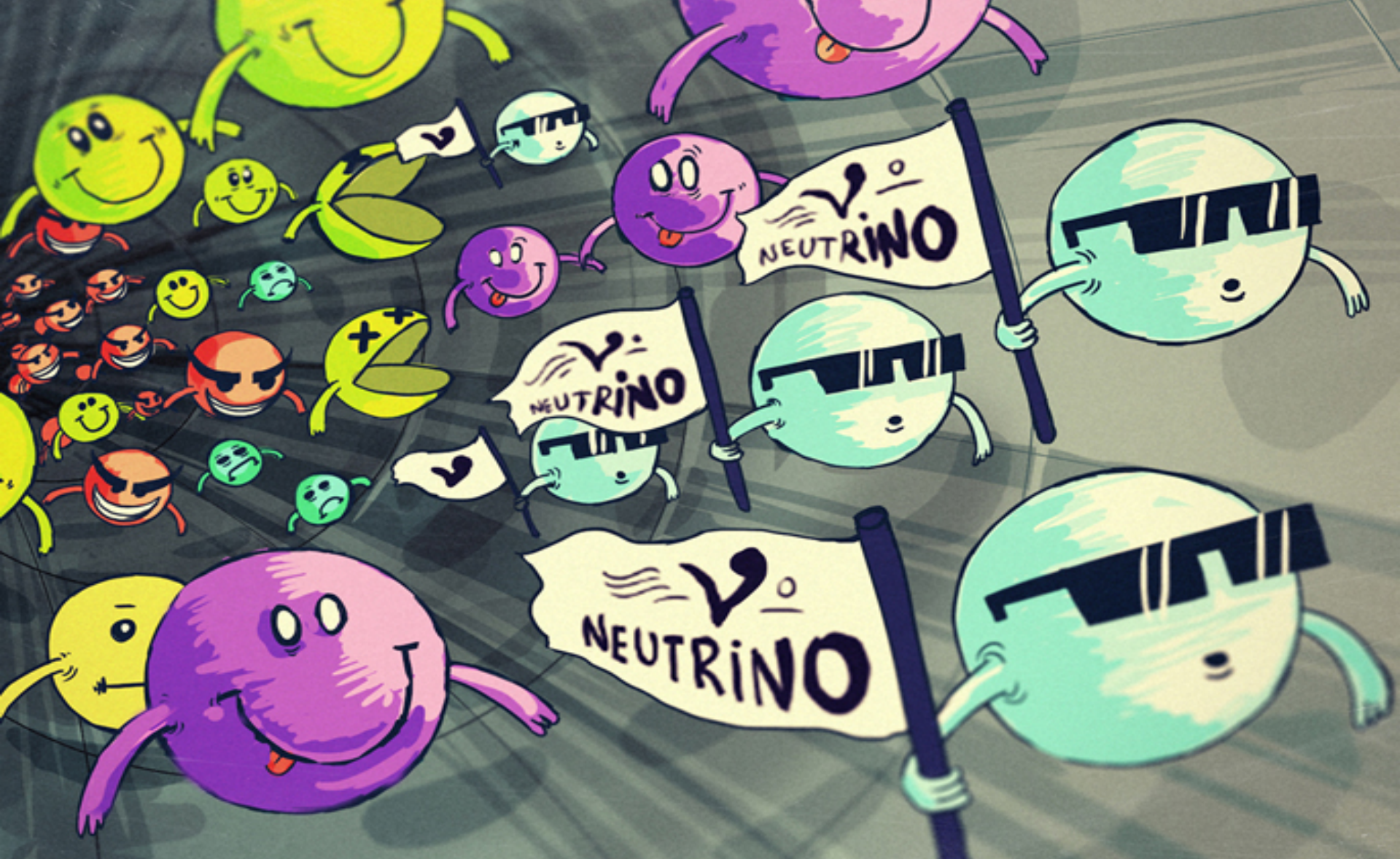
# The Oscillation Probabilities



# Constraining "LED"

IceCube- Z. Tabrizi, A. Esmaili, O.L.G. Peres





Thanks for your attention