

# **The Emergence of Spacetime: Entanglement & Complexity**

**Robert Myers**  
**Lecture 5**

# **“It from Qubit”:** New Collision of Ideas

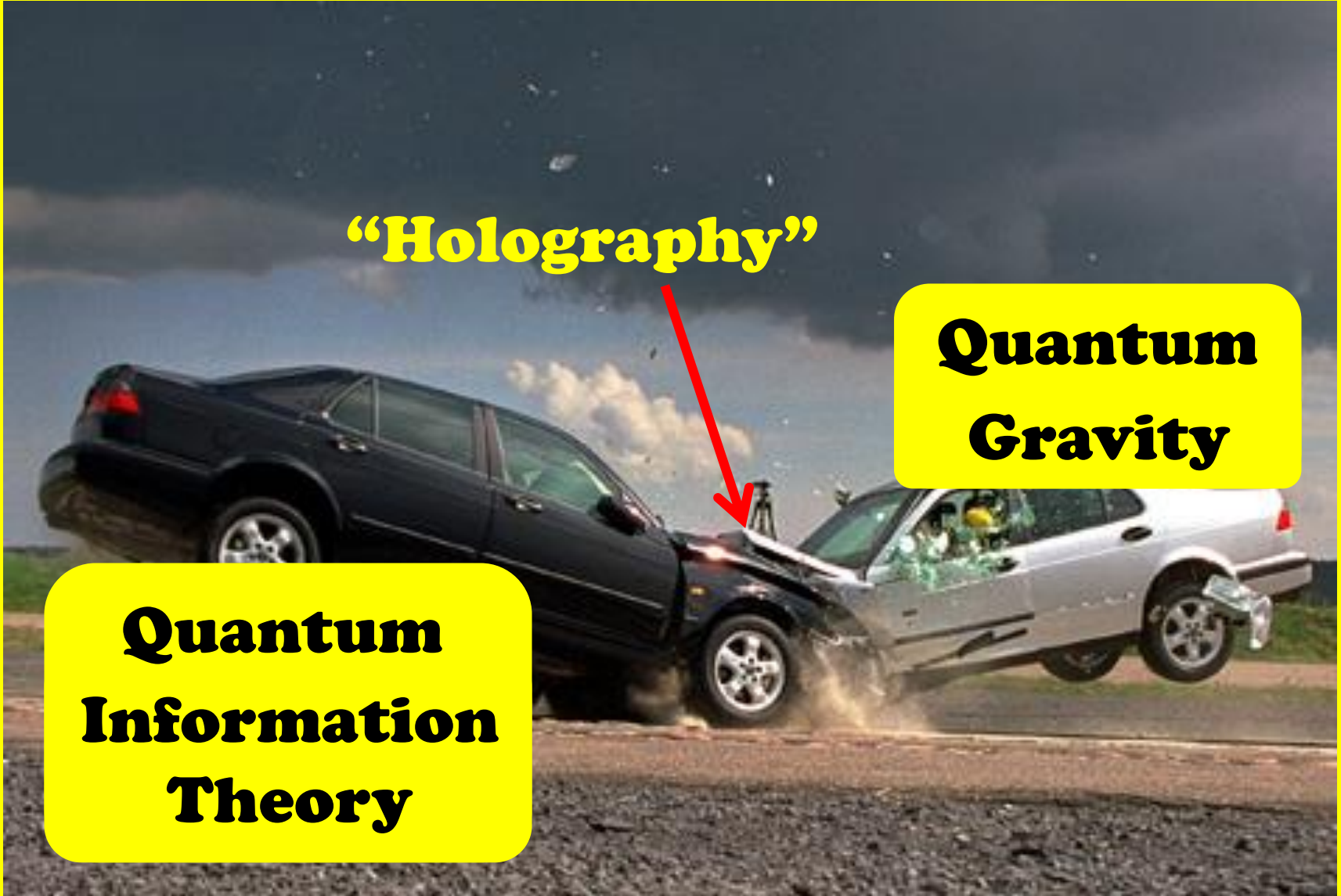


# **“It from Qubit”**: New Collision of Ideas

**“Holography”**

**Quantum  
Gravity**

**Quantum  
Information  
Theory**



~~Nexus~~  
~~Confluence~~  
Dialogue

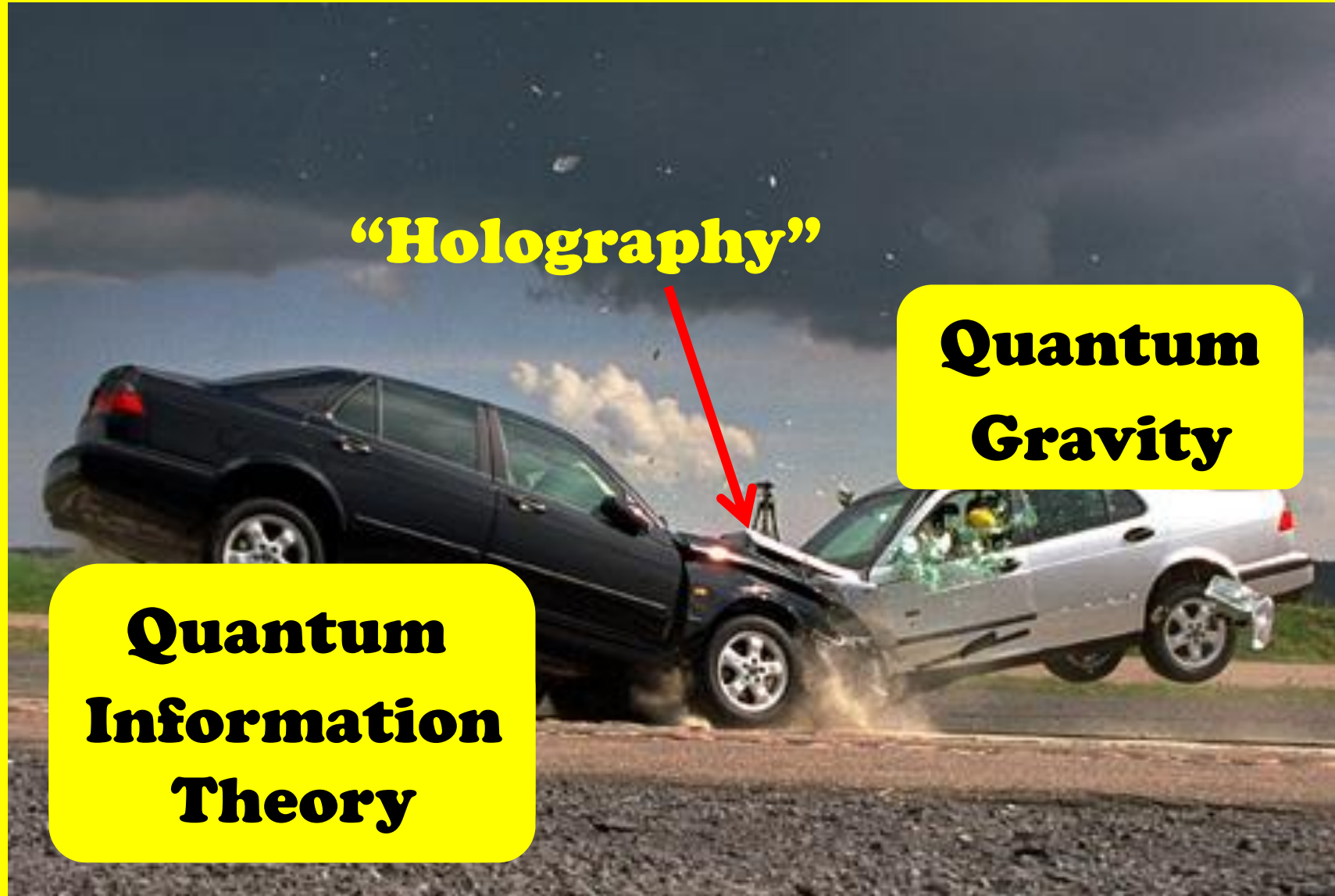
# **“It from Qubit”**: New ~~Collision~~ of Ideas

**“Holography”**

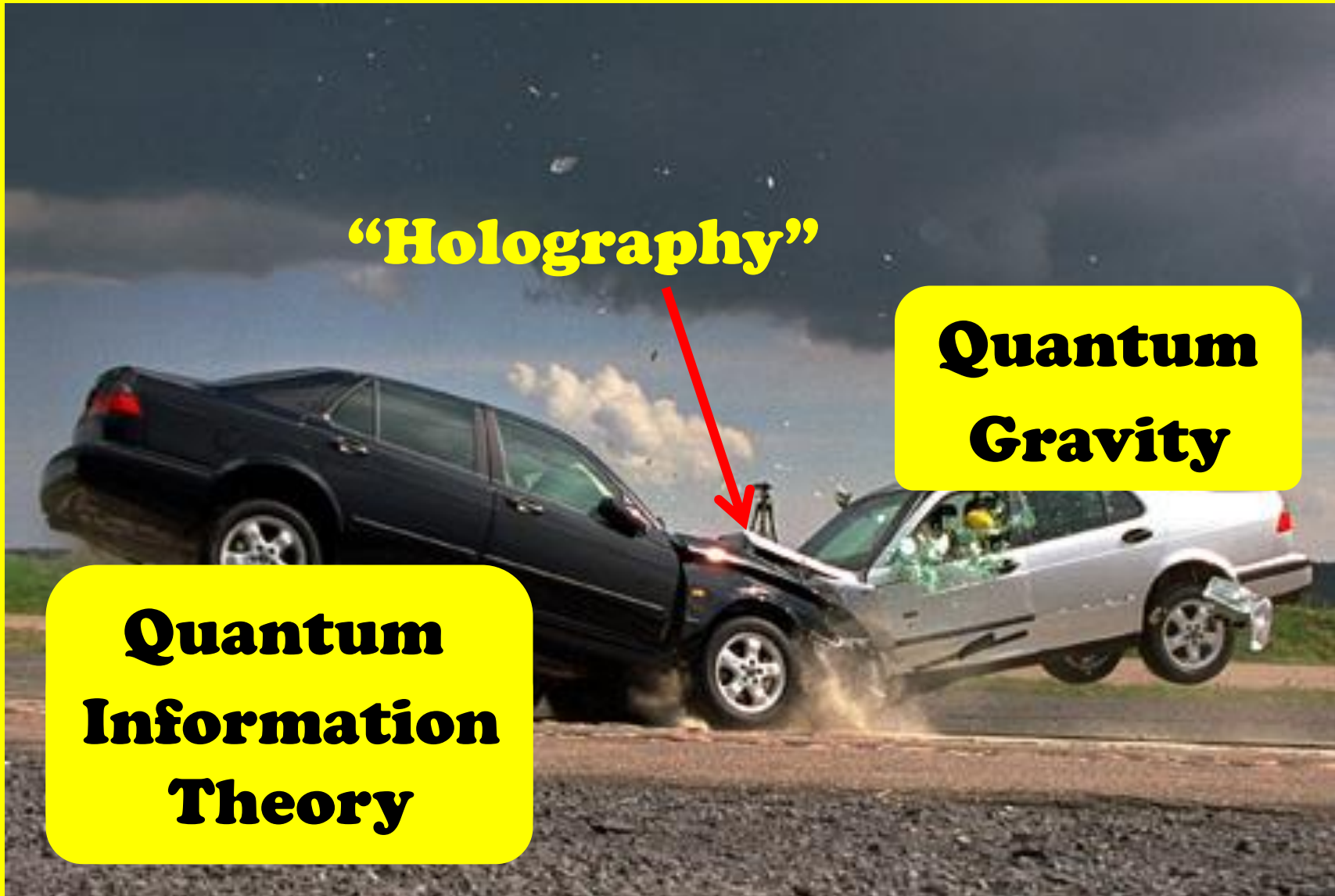


**Quantum Gravity**

**Quantum Information Theory**



# **“It from Qubit”**: New Collision of Ideas



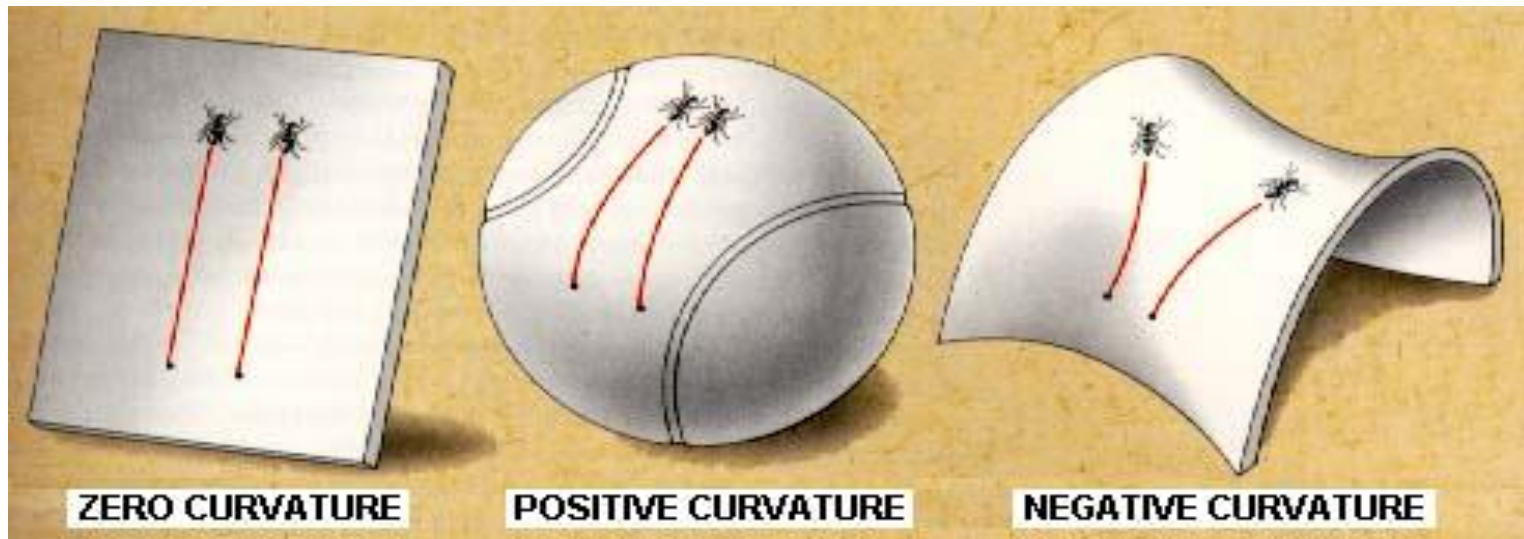
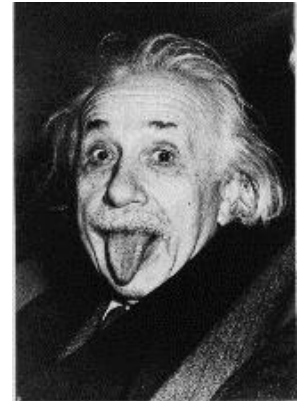
**“Holography”**

**Quantum  
Gravity**

**Quantum  
Information  
Theory**

# Einstein:

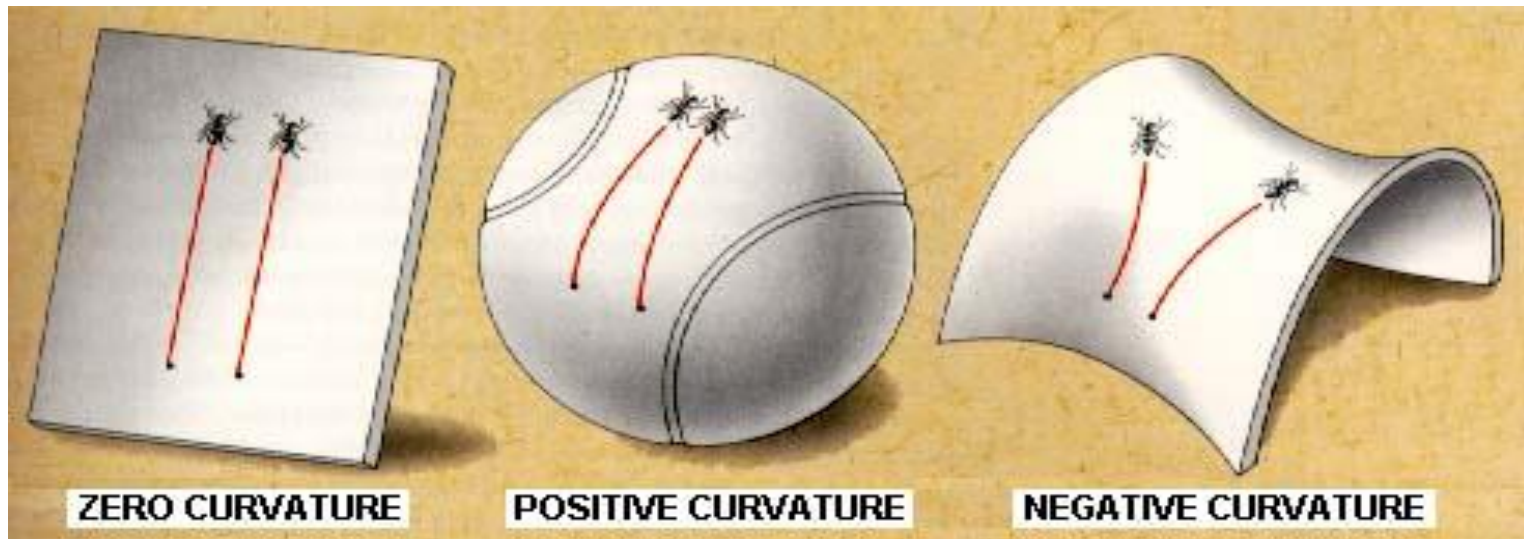
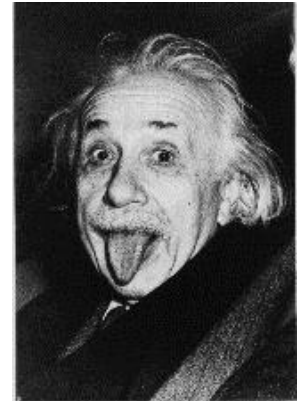
Gravity? It's all about geometry!



**General Relativity** is the geometric arena for physics on very large scales: planets, stars, galaxies, cosmology

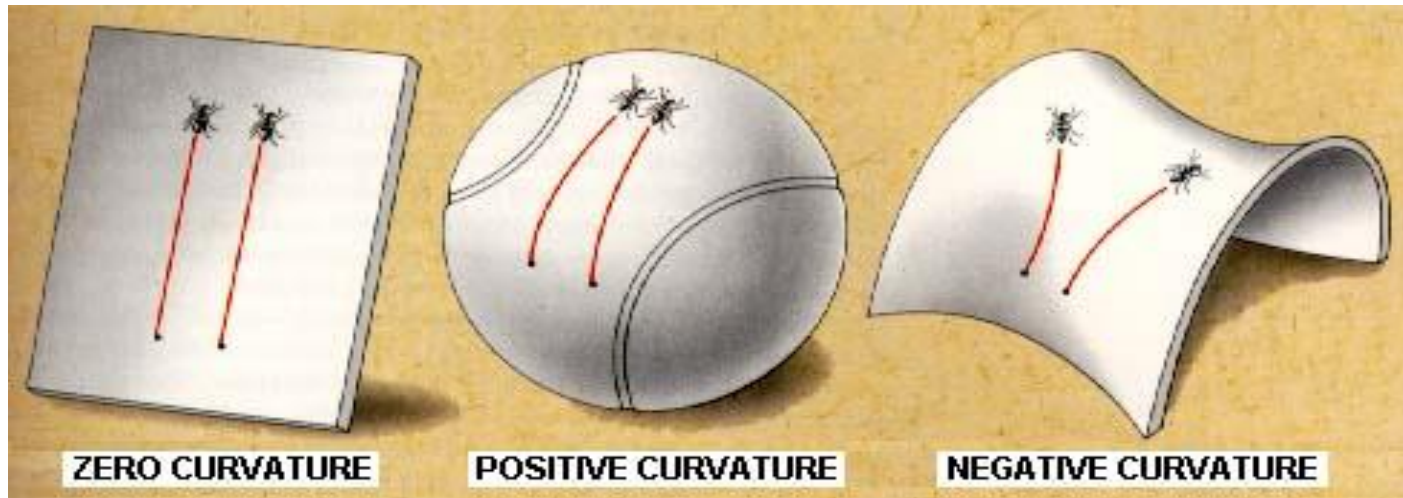
# Einstein:

Gravity? It's all about geometry!



**Spacetime** moves from simply stage for physical phenomena, to being both the stage and an active player in the dynamics

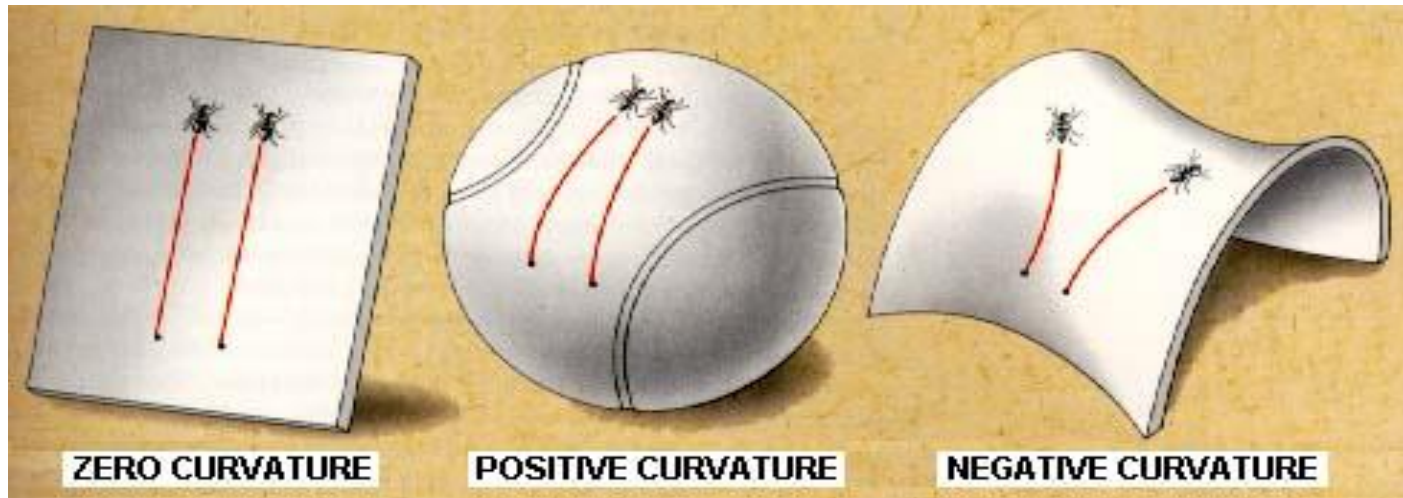
# General Relativity + Quantum Theory = Quantum Gravity?



+  $\hbar$

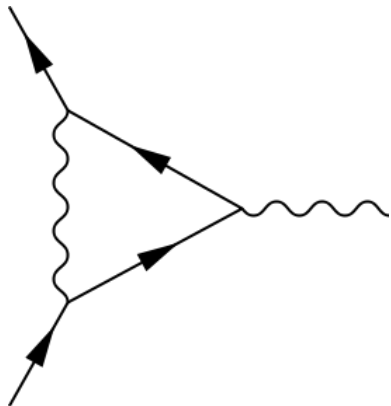


# General Relativity + Quantum Theory = Quantum Gravity?



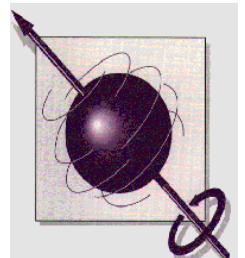
+  $\hbar$

- quantum fluctuations become manifest at small scales  
e.g., magnetic moment of the electron,  $\mu_e = g e \hbar / 4 m_e$ ,  
with  $g \approx 2$  but modified by quantum fluctuations

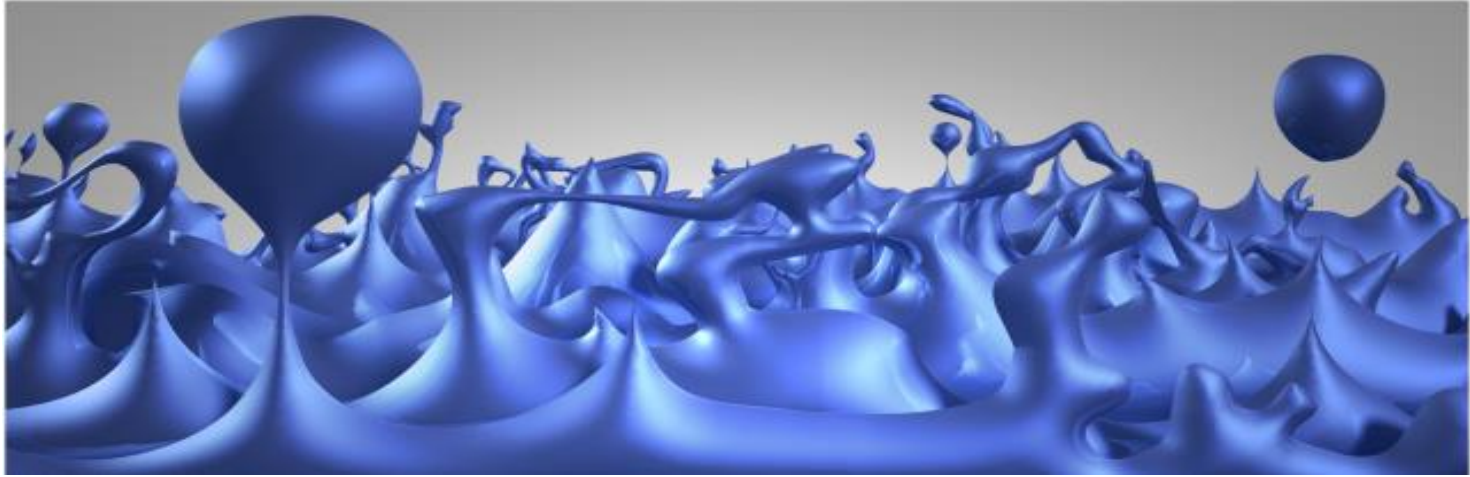


$$g_{\text{theory}} = 2.0023193043070$$

$$\left| \frac{\mu_{\text{theory}} - \mu_{\text{experiment}}}{\mu_{\text{experiment}}} \right| \leq 10^{-10}$$

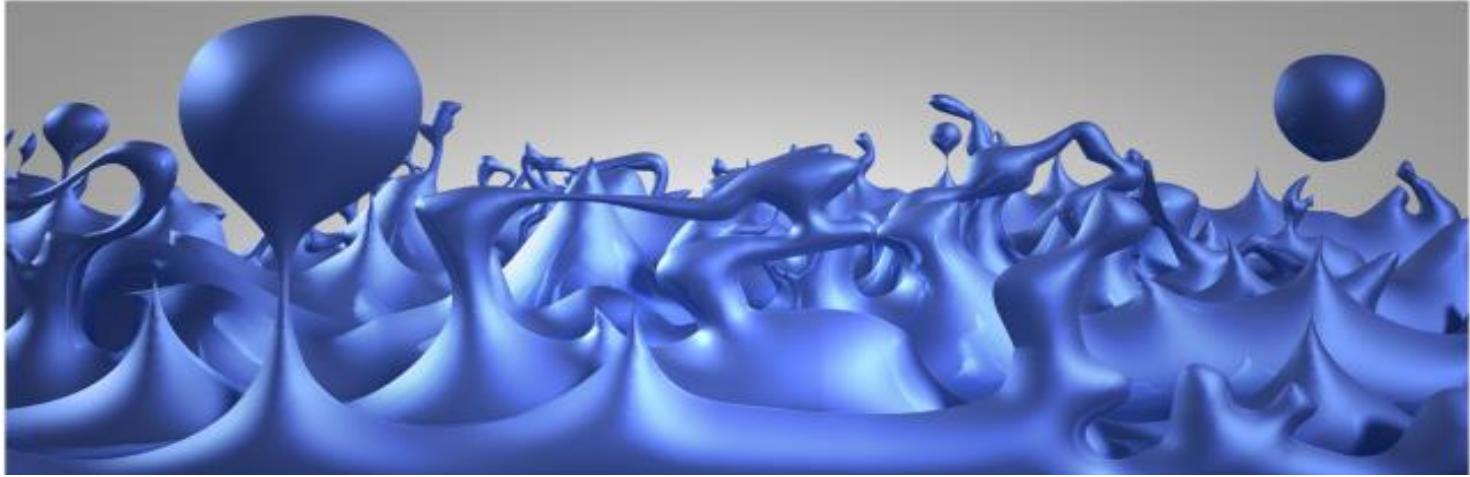


# General Relativity + Quantum Theory = Quantum Gravity?

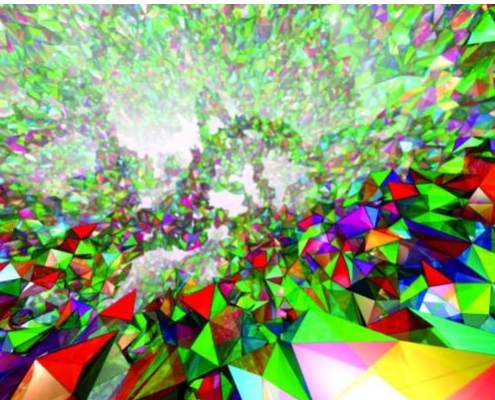


- spacetime geometry exhibits strong fluctuations when examined on very short distance scales
- how do we make sense of spacetime framework?

# General Relativity + Quantum Theory = Quantum Gravity?

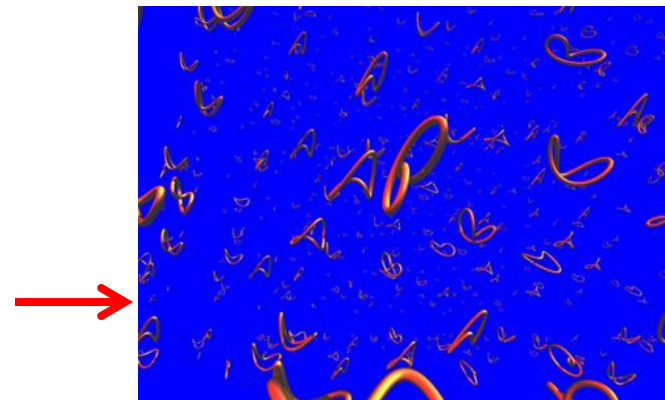


- spacetime geometry exhibits strong fluctuations when examined on very short distance scales
- how do we make sense of spacetime framework?

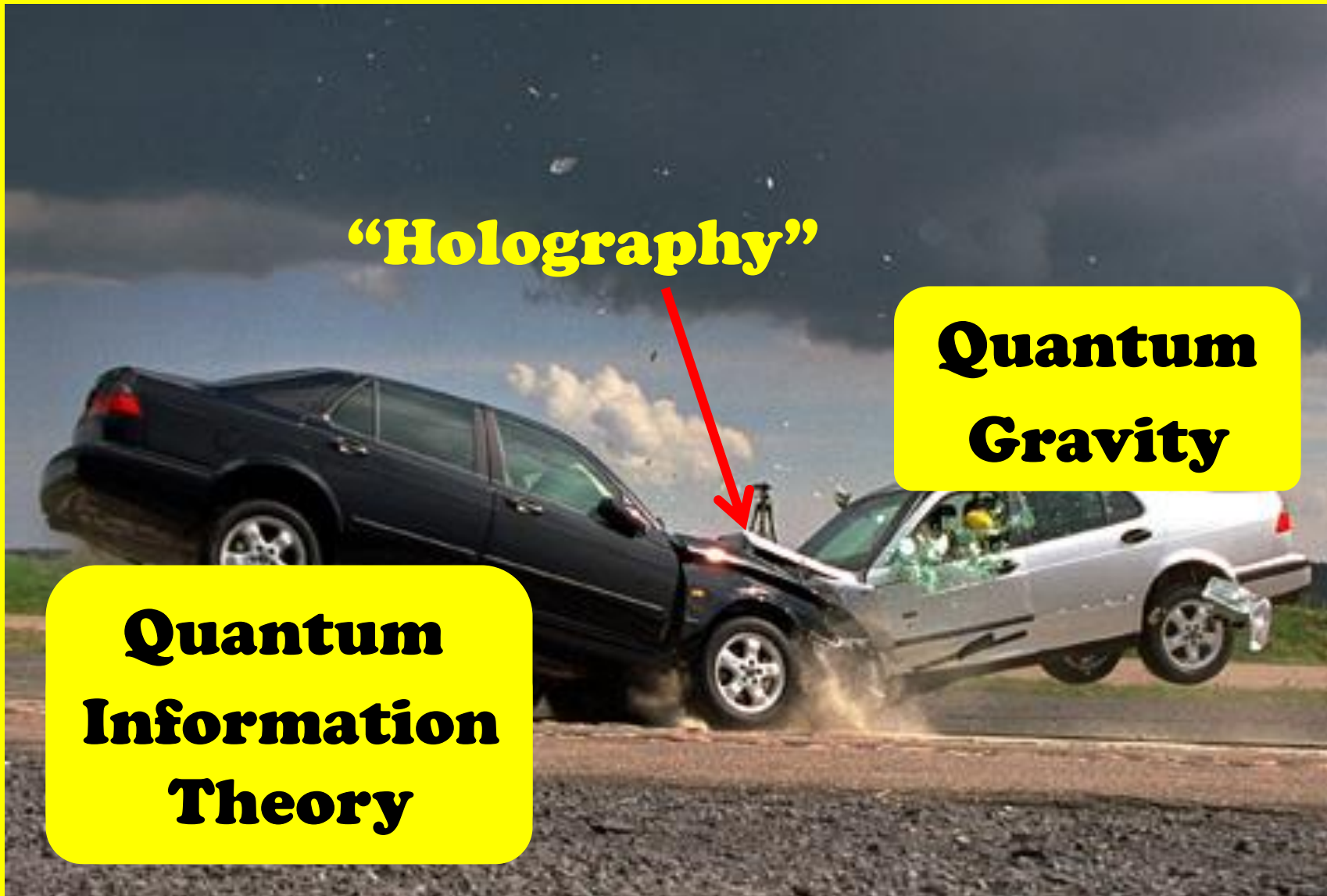


← modify geometry  
at short distances

modify spectrum  
at short distances →



# **“It from Qubit”**: New Collision of Ideas



**“Holography”**

**Quantum  
Gravity**

**Quantum  
Information  
Theory**

# Quantum Entanglement

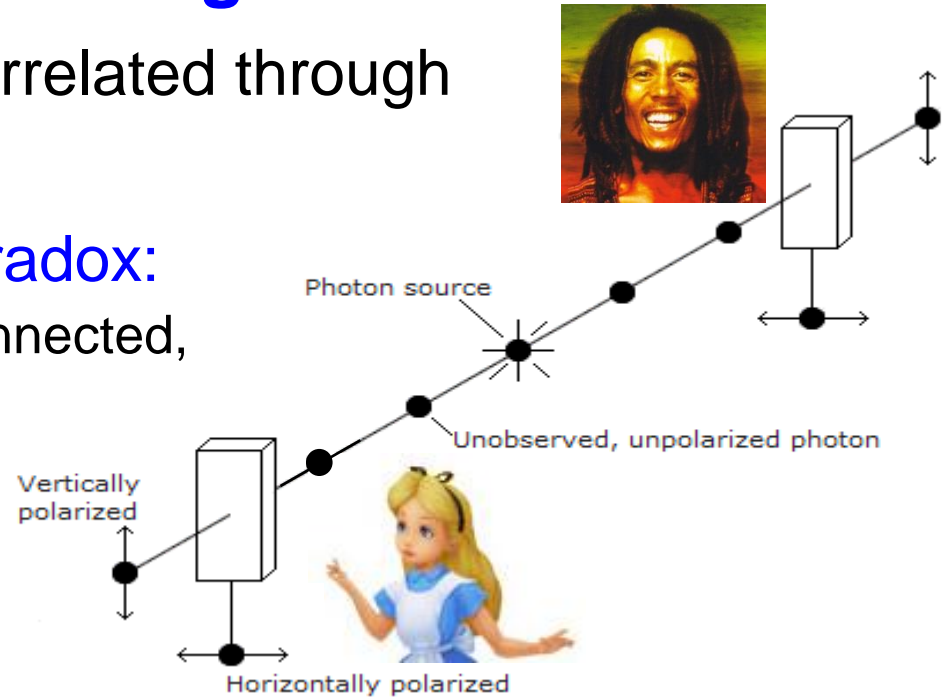
- different subsystems are correlated through global state of full system

## Einstein-Podolsky-Rosen Paradox:

- polarizations of pair of photons connected, no matter how far apart they travel

“*spukhafte Fernwirkung*” = spooky action at a distance

$$|\psi\rangle = \frac{1}{\sqrt{2}} \left( |\uparrow\downarrow\rangle + |\downarrow\uparrow\rangle \right)$$



# Quantum Entanglement

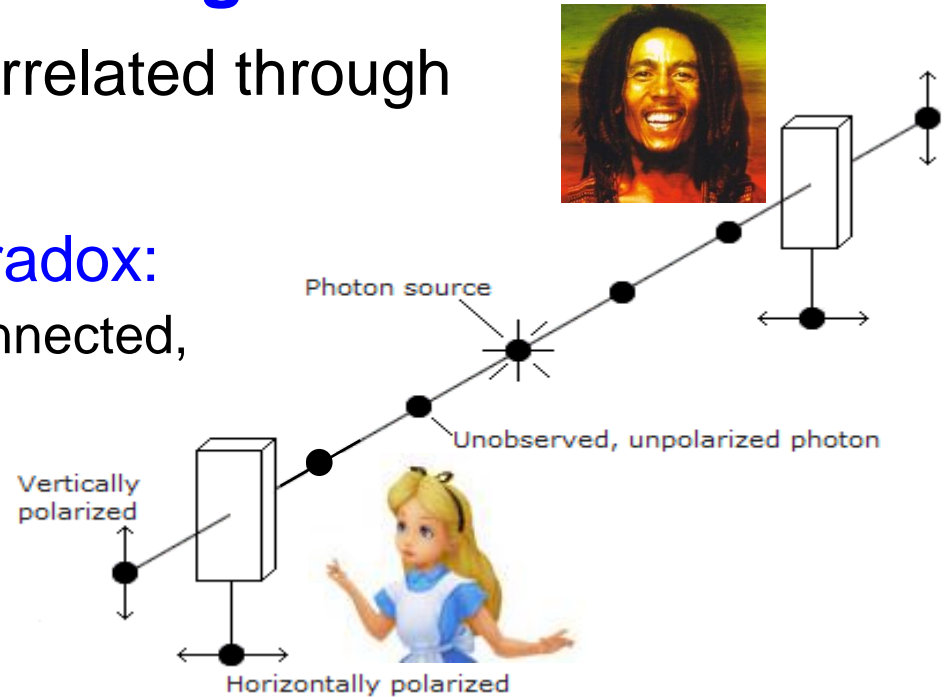
- different subsystems are correlated through global state of full system

## Einstein-Podolsky-Rosen Paradox:

- polarizations of pair of photons connected, no matter how far apart they travel

“*spukhafte Fernwirkung*” = spooky action at a distance

$$|\psi\rangle = \frac{1}{\sqrt{2}} \left( |\uparrow\downarrow\rangle + |\downarrow\uparrow\rangle \right)$$



**Quantum Information:** entanglement becomes a resource for (ultra)fast computations and (ultra)secure communications

**Condensed Matter:** key to “exotic” phases and phenomena, e.g., quantum Hall fluids, unconventional superconductors, quantum spin fluids, . . . .

# Quantum Entanglement

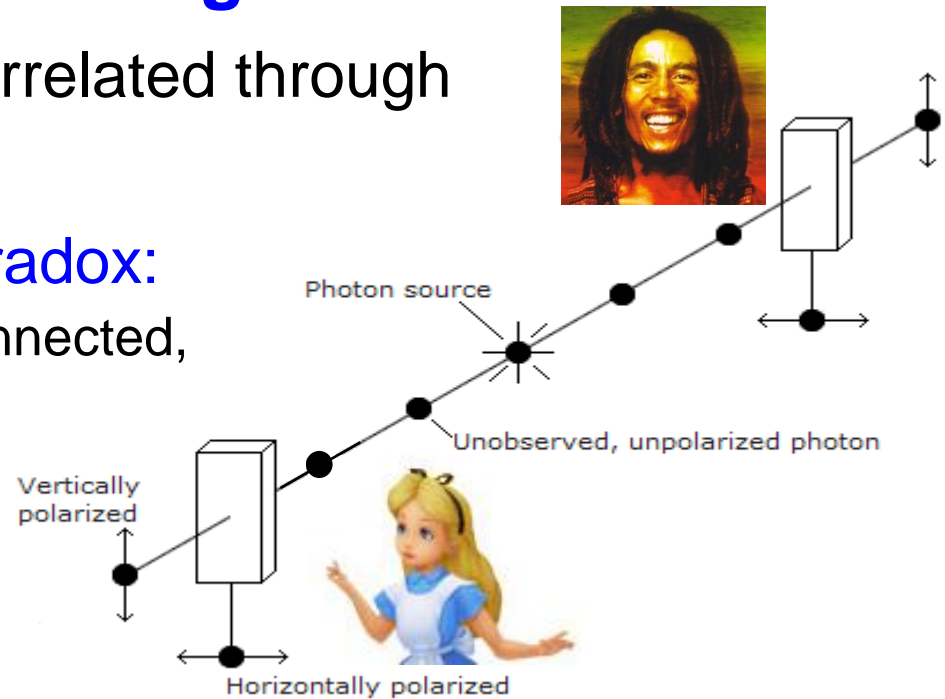
- different subsystems are correlated through global state of full system

## Einstein-Podolsky-Rosen Paradox:

- polarizations of pair of photons connected, no matter how far apart they travel

“*spukhafte Fernwirkung*” = spooky action at a distance

$$|\psi\rangle = \frac{1}{\sqrt{2}} \left( |\uparrow\downarrow\rangle + |\downarrow\uparrow\rangle \right)$$



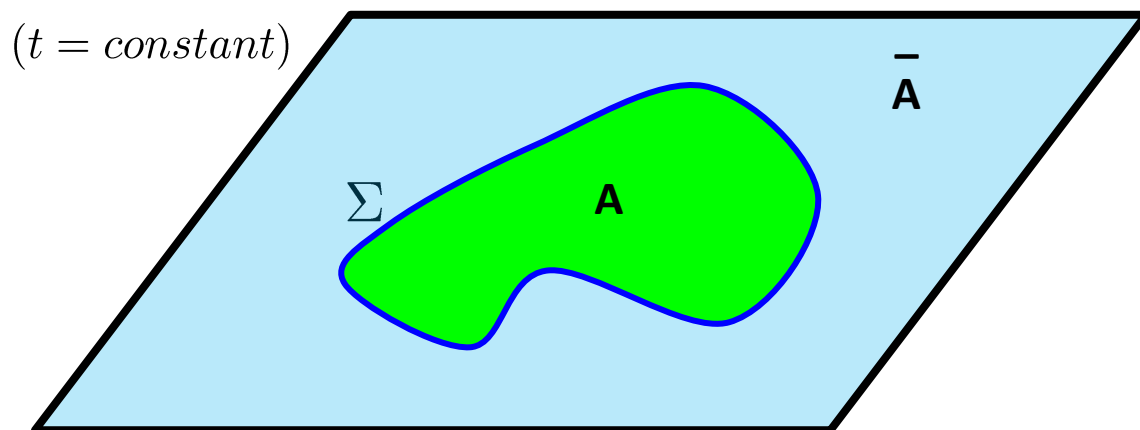
**Quantum Information:** entanglement becomes a resource for (ultra)fast computations and (ultra)secure communications

**Condensed Matter:** key to “exotic” phases and phenomena, e.g., quantum Hall fluids, unconventional superconductors, quantum spin fluids, . . . .

**Quantum Fields & Quantum Gravity**

## Entanglement Entropy in QFT

- general diagnostic to give a quantitative measure of entanglement using **entropy** to detect correlations between two subsystems
  - in QFT, typically introduce a (smooth) boundary or **entangling surface**  $\Sigma$  which divides the space into two separate regions
  - integrate out degrees of freedom in “outside” region
  - remaining dof are described by a density matrix  $\rho_A$
- calculate **von Neumann entropy**:  $S_{EE} = -\text{Tr} [\rho_A \log \rho_A]$





# Holography: AdS/CFT correspondence

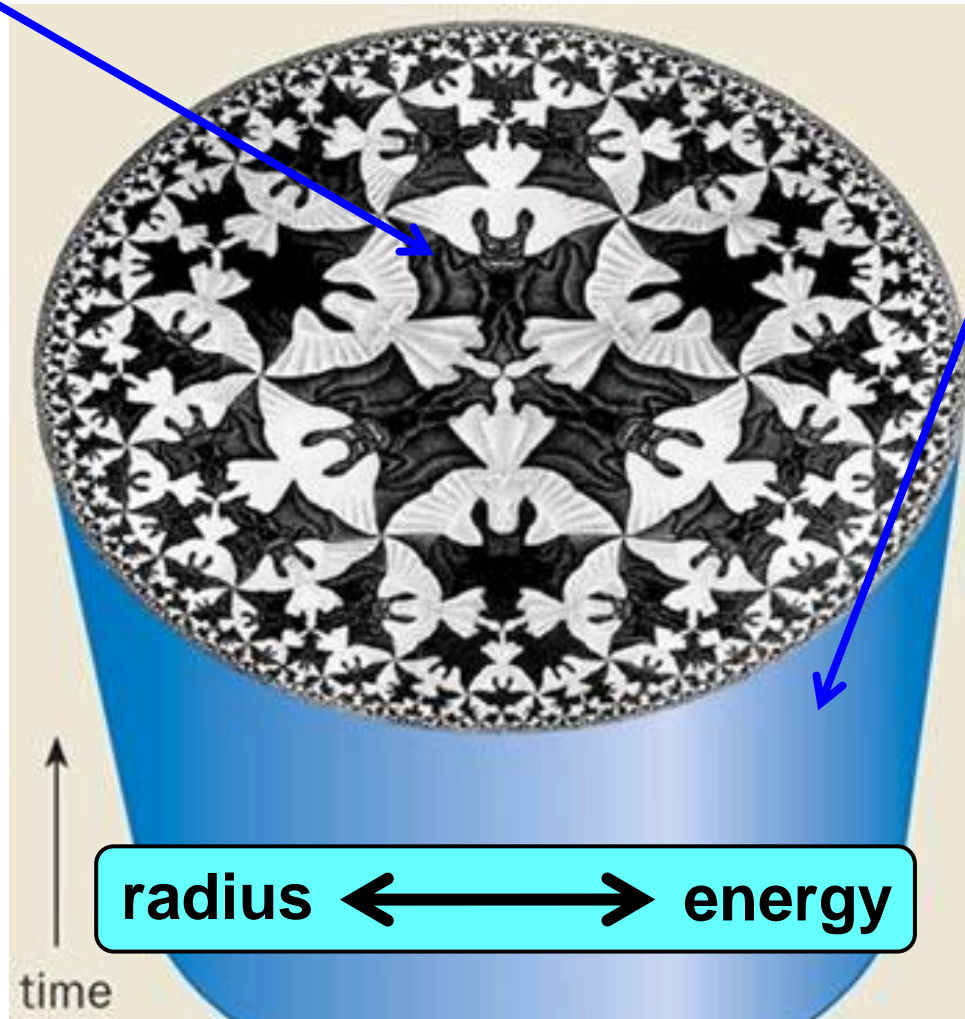
Bulk: gravity with negative  $\Lambda$   
in  $d+1$  dimensions

Boundary: quantum field theory  
without intrinsic scales  
in  $d$  dimensions

↔  
“holography”

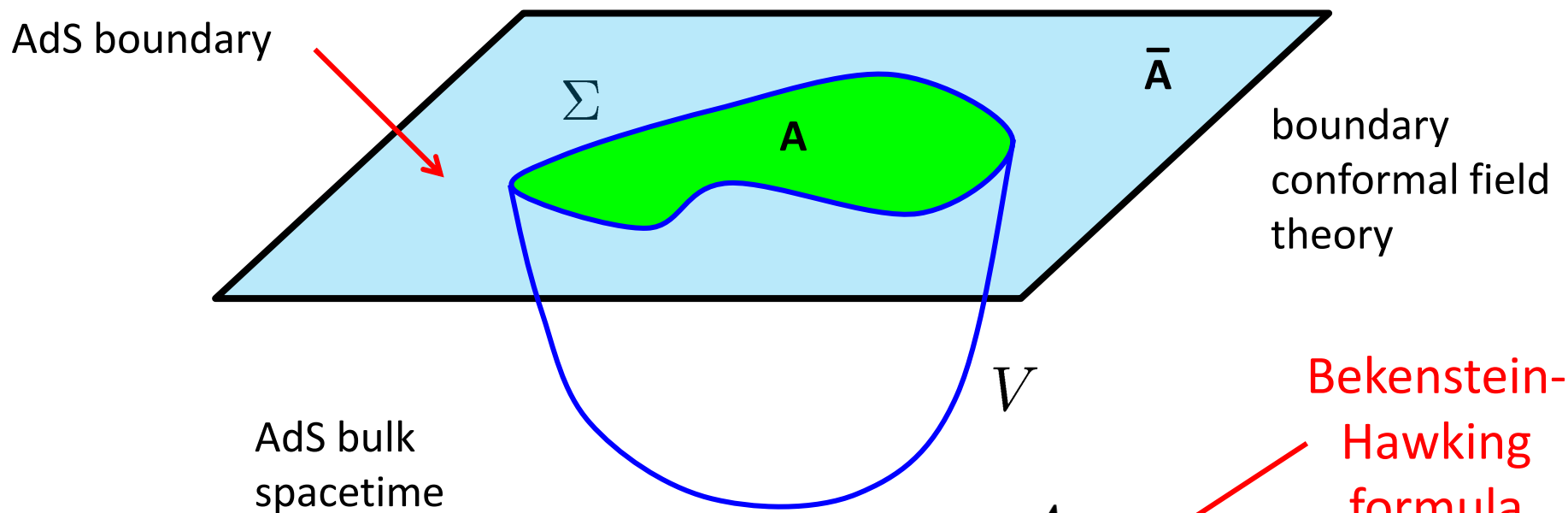
anti-de Sitter  
space

conformal  
field theory



(Maldacena '97)

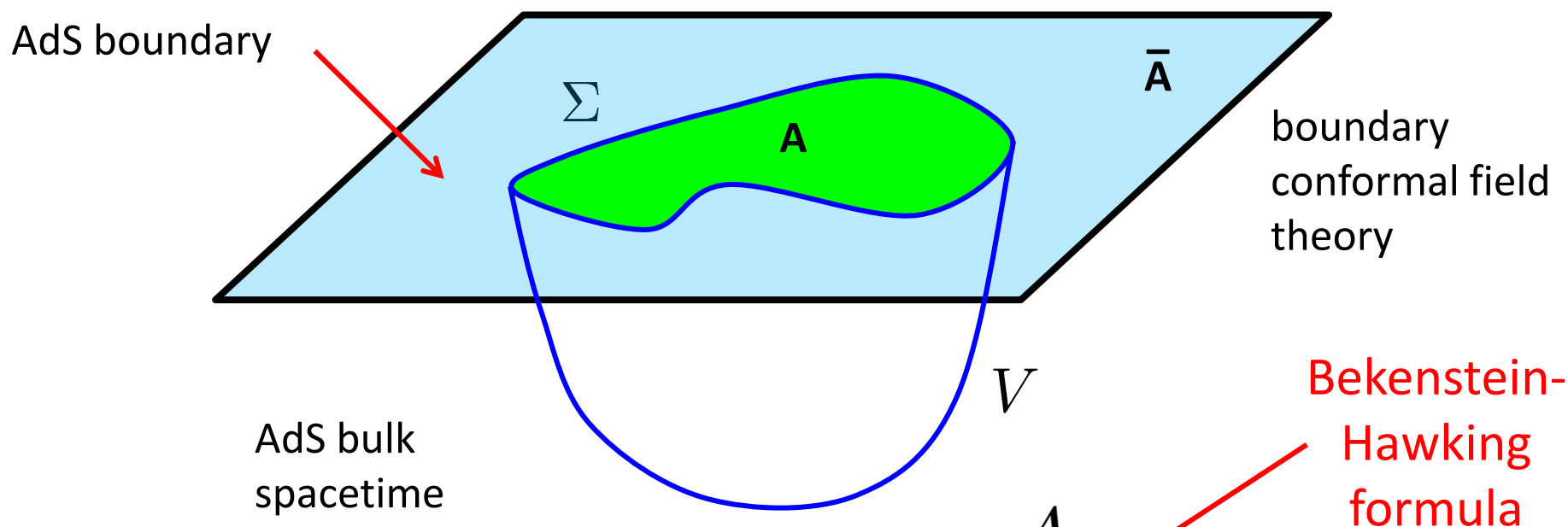
# Holographic Entanglement Entropy:



Bekenstein-Hawking formula

$$S(A) = \text{ext}_{V \sim A} \frac{A_V}{4G_N}$$

# Holographic Entanglement Entropy:

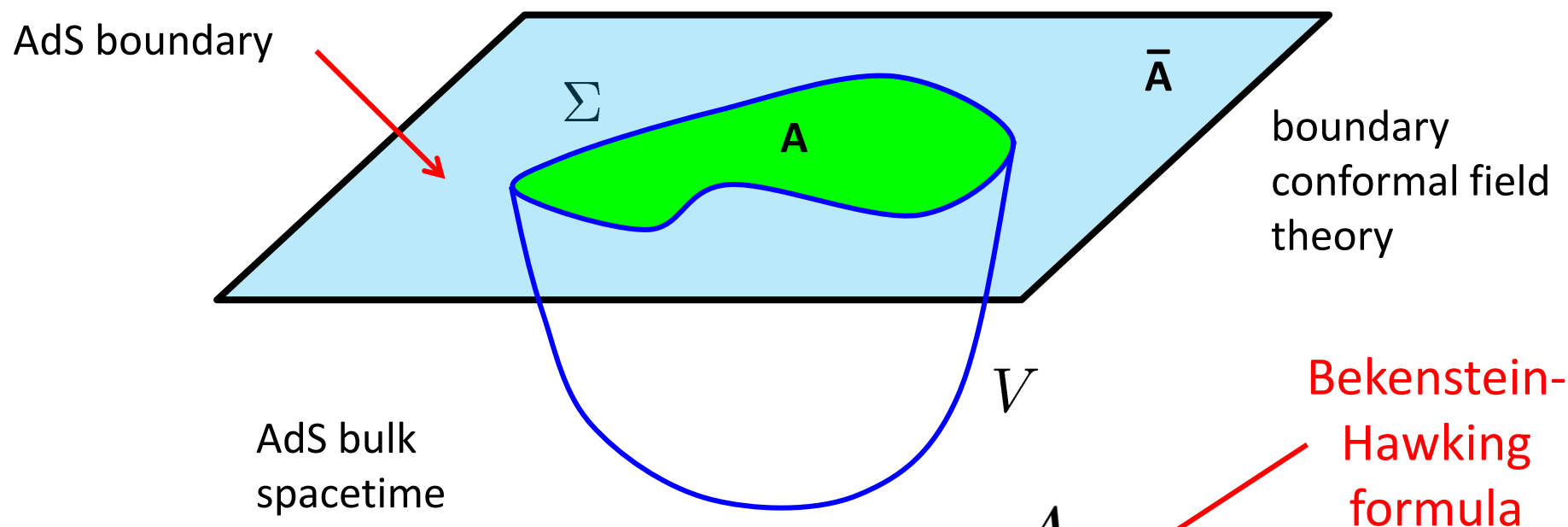


$$S(A) = \text{ext}_{V \sim A} \frac{A_V}{4G_N}$$

Bekenstein-Hawking formula

- conjecture  $\longrightarrow$  many detailed consistency tests  
(Ryu, Takayanagi, Hubeny, Rangamani, Headrick, Hung, Smolkin, RM, Faulkner, . . .)

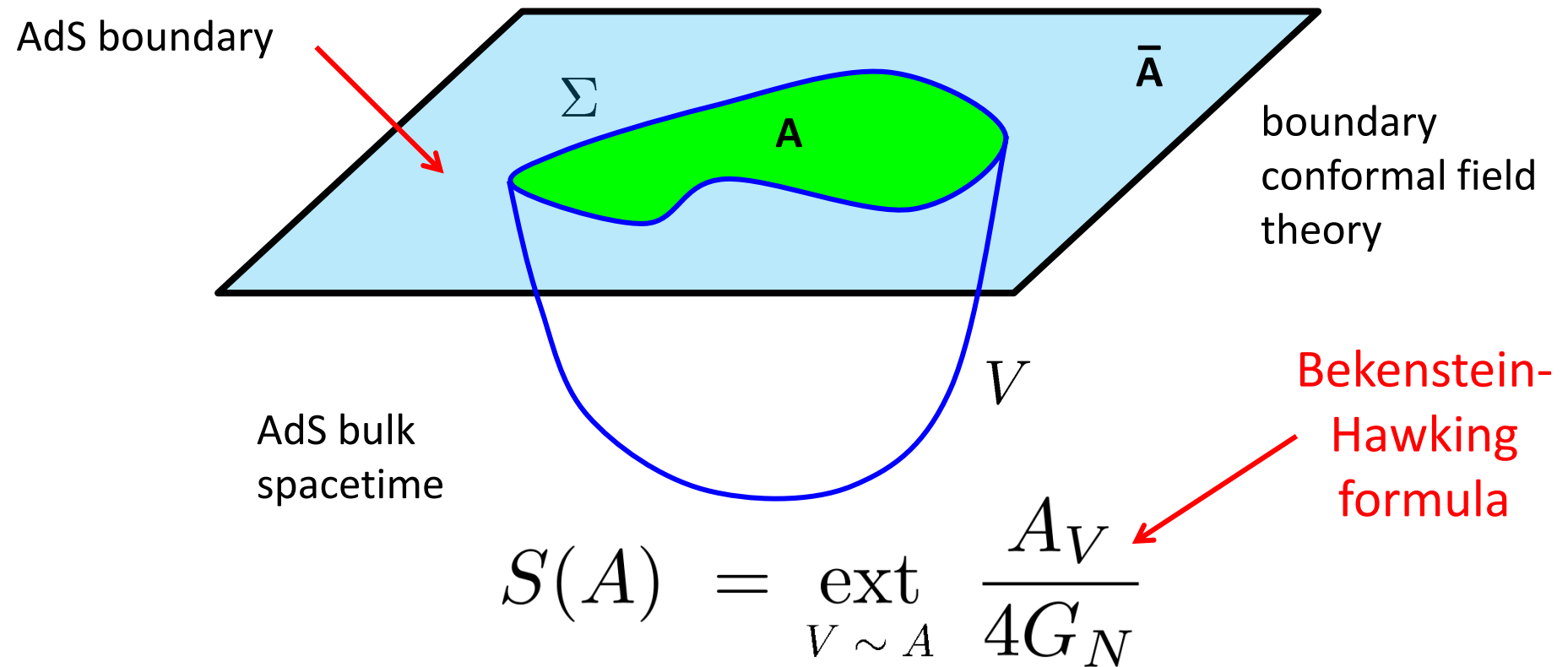
# Holographic Entanglement Entropy:



$$S(A) = \text{ext}_{V \sim A} \frac{A_V}{4G_N}$$

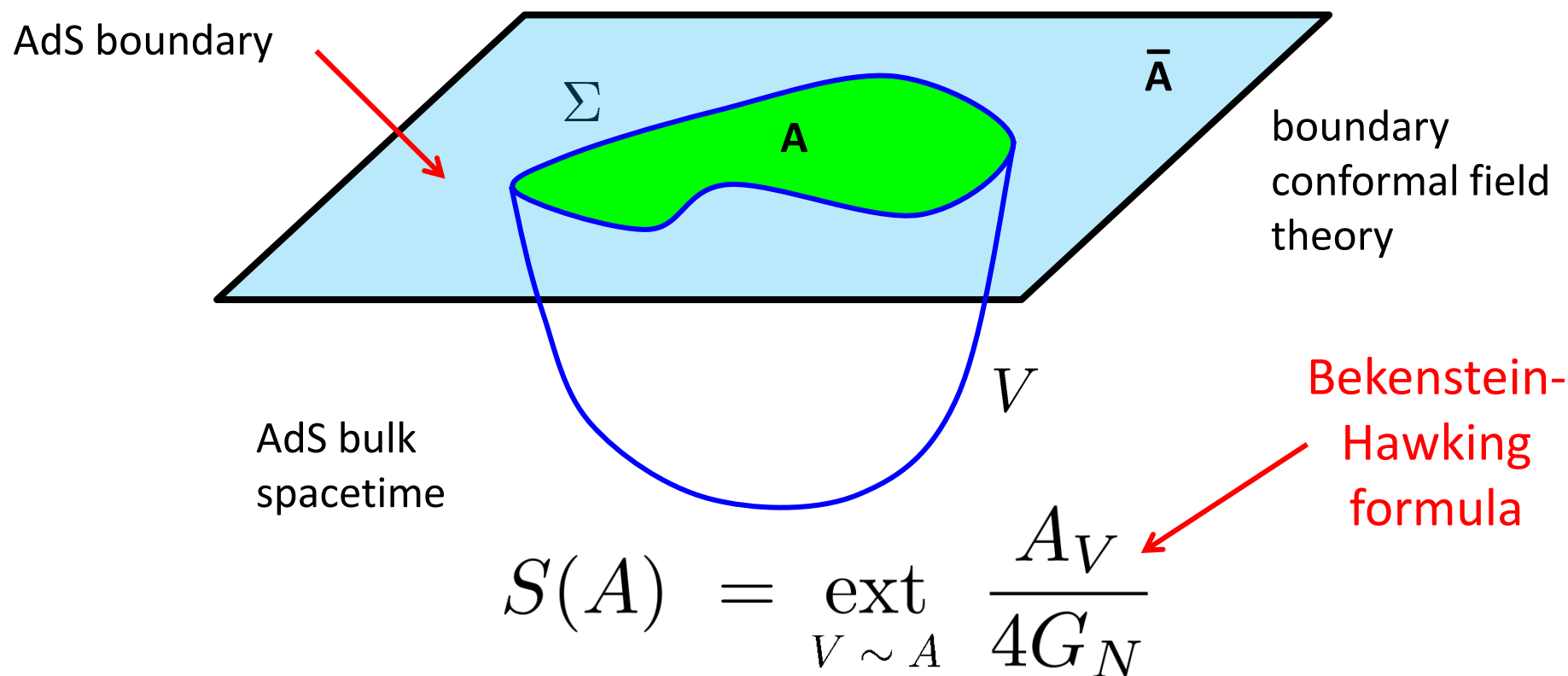
- conjecture  $\longrightarrow$  many detailed consistency tests  
(Ryu, Takayanagi, Hubeny, Rangamani, Headrick, Hung, Smolkin, RM, Faulkner, . . .)
- 2013 proof (for static geometries) (Maldacena & Lewkowycz)
- 2016 proof (for general geometries) (Dong, Lewkowycz & Rangamani)

# Holographic Entanglement Entropy:



- conjecture  $\longrightarrow$  many detailed consistency tests  
(Ryu, Takayanagi, Hubeny, Rangamani, Headrick, Hung, Smolkin, RM, Faulkner, . . .)
- 2013 proof (for static geometries) (Maldacena & Lewkowycz)
- 2016 proof (for general geometries) (Dong, Lewkowycz & Rangamani)
- holographic EE: fruitful forum for bulk-boundary dialogue

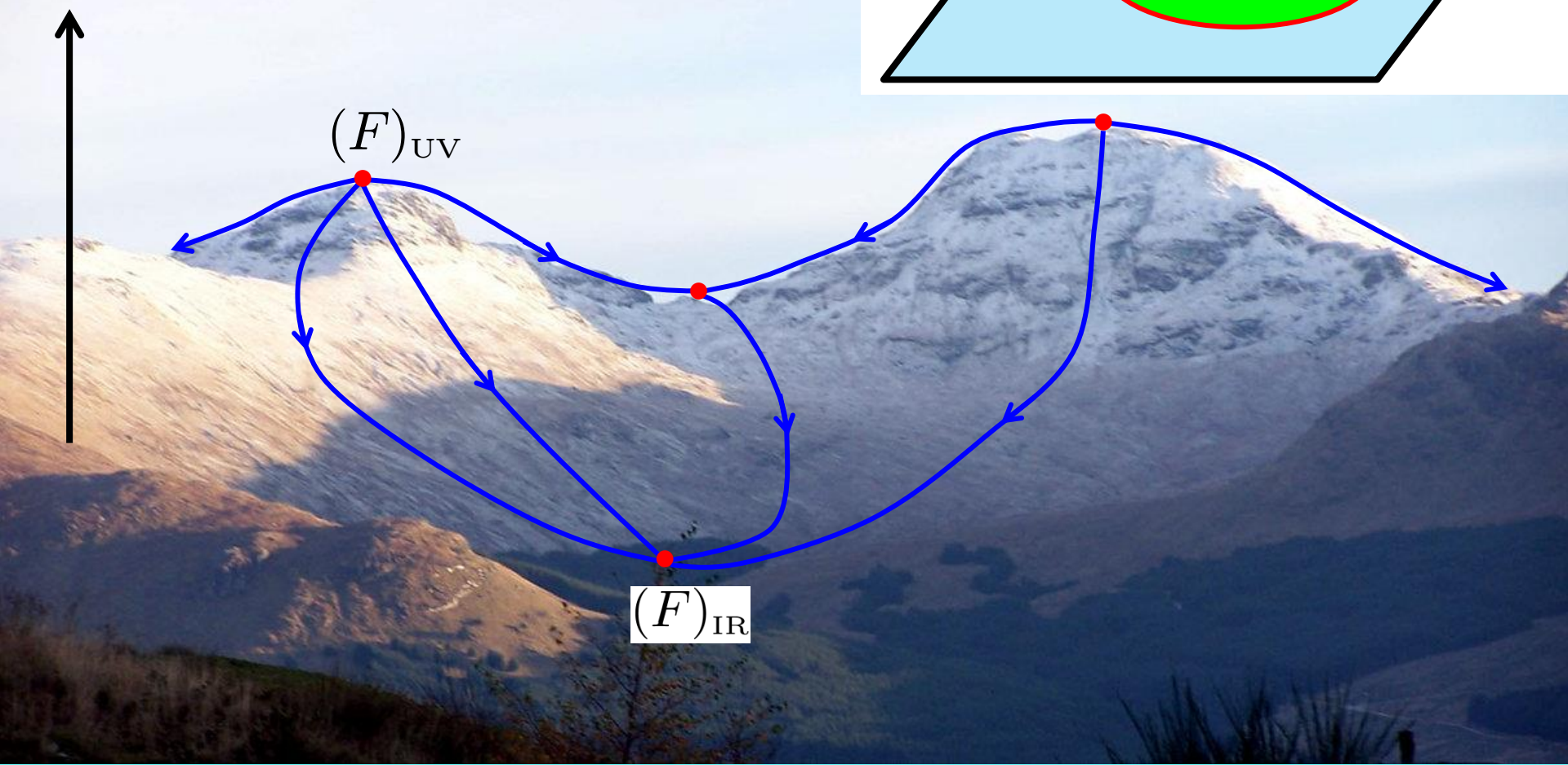
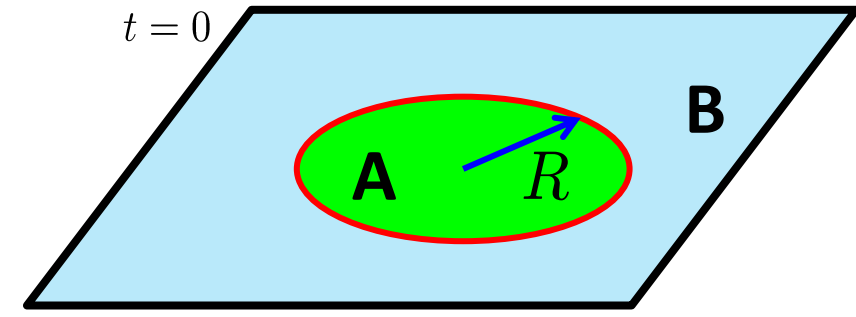
# Holographic Entanglement Entropy:



- holographic EE teaches us lessons about QFTs, eg,
  - diagnostic in RG flows and c-theorems, eg, F-theorem (Sinha & RM, .....

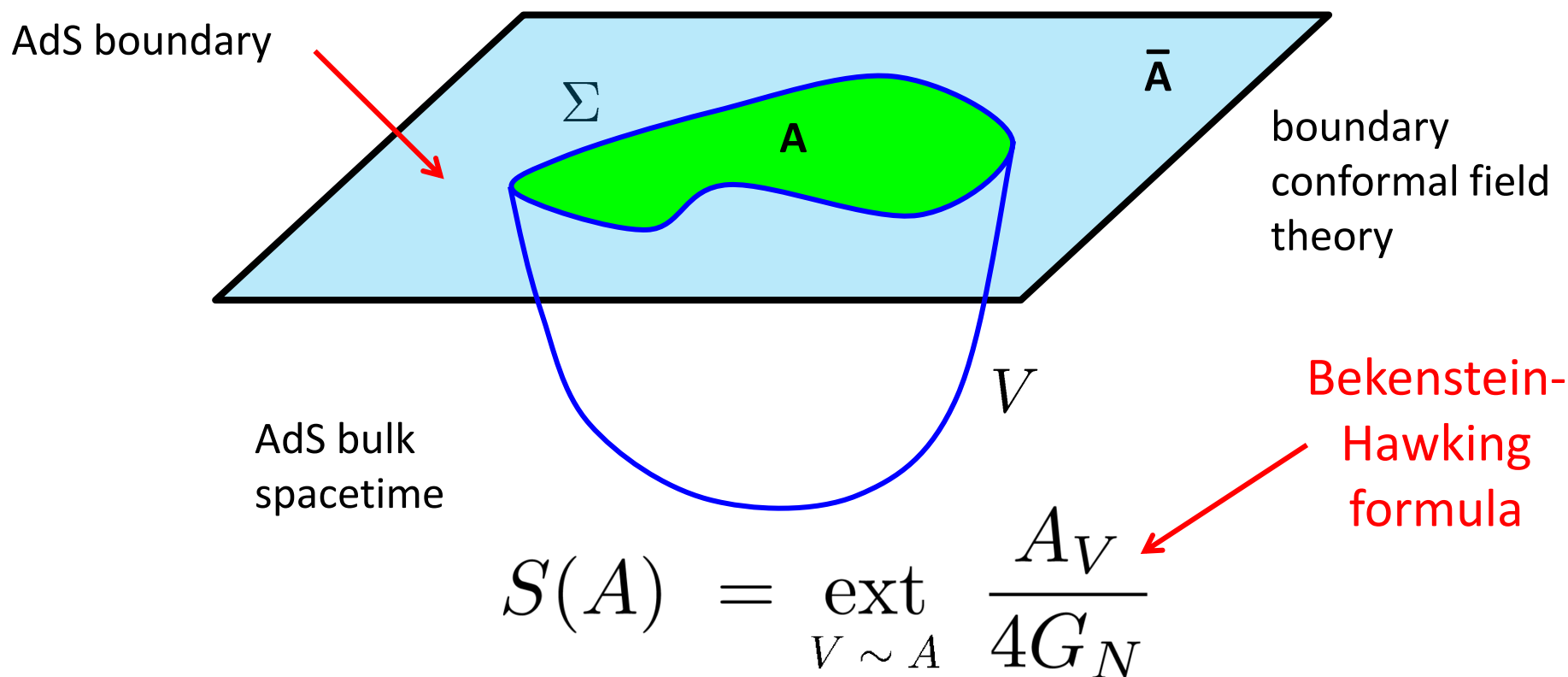
gravity/holography + EE  $\longrightarrow$  RG flows in (2+1)-dimensions

$$C(R) = R \partial_R S(R) - S(R)$$



F-theorem:  $(F)_{UV} \geq (F)_{IR}$

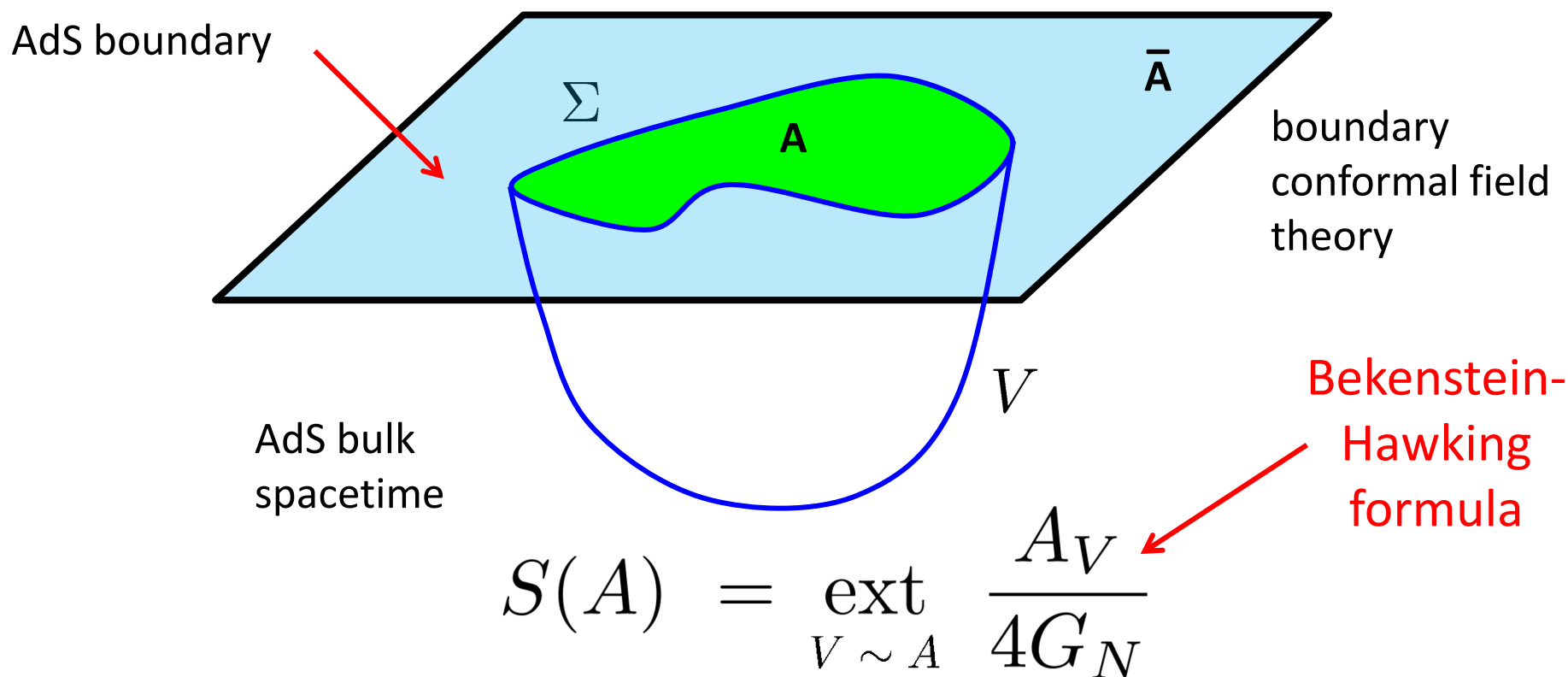
# Holographic Entanglement Entropy:



- holographic EE teaches us lessons about QFTs, eg,
  - diagnostic in RG flows and c-theorems, eg, F-theorem (Sinha & RM, .....
  - geometric properties of entanglement entropy in QFT's (Mezei, Perlmutter, Lewkowycz, Bueno, RM, Witczak-Krempa, ....)
  - diagnostic for quantum quenches/phase transitions (Lopez, Johnson, Balasubramanian, Bernamonti, Craps, Galli, ....)



# Holographic Entanglement Entropy:



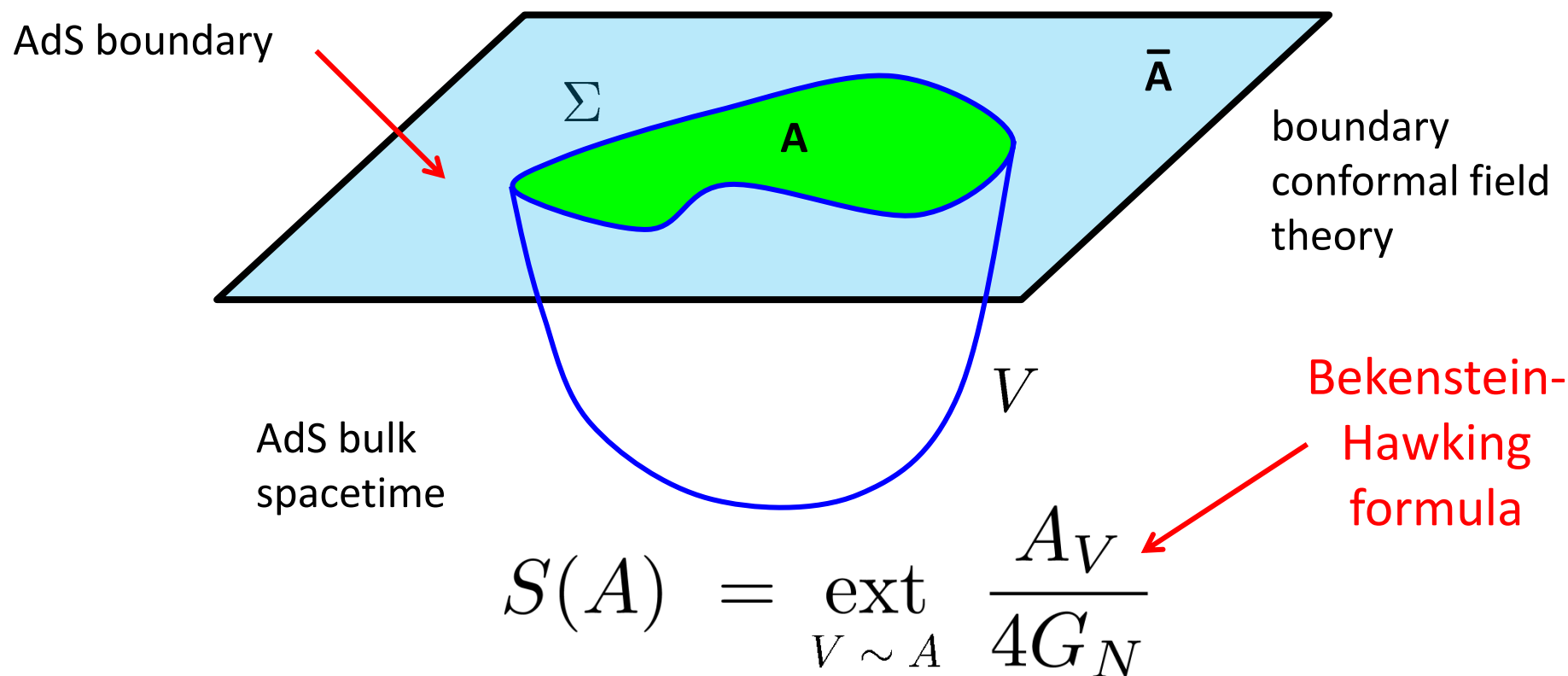
- holographic EE teaches us lessons about (quantum) gravity, eg,

→ BH formula applies beyond black holes/horizons

→ connectivity of spacetime requires entanglement

(van Raamsdonk)

# Holographic Entanglement Entropy:



- holographic EE teaches us lessons about (quantum) gravity, eg,

→ BH formula applies beyond black holes/horizons

→ connectivity of spacetime requires entanglement

(van Raamsdonk)

## Spacetime Geometry = Entanglement

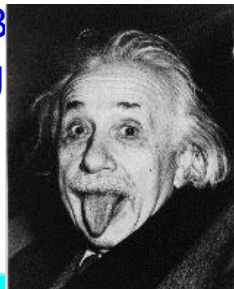
# Spacetime Geometry = Entanglement

- Bekenstein-Hawking formula: spacetime geometry encodes  $S_{\text{BH}}$
- black hole entropy is entanglement entropy (Sorkin, ....)
- use BH formula for holographic entanglement entropy (Ryu & Takayanagi; ....)
- connectivity of spacetime requires entanglement (van Raamsdonk)
- spacetime entanglement conjecture (Bianchi & RM)
- AdS spacetime as a tensor network (MERA) (Swingle, Vidal, ....)
- “ER = EPR” conjecture (Maldacena & Susskind)
- hole-ographic spacetime (Balasubramanian, Chowdhury, Czech, de Boer & Heller; RM, Rao & Sugishita; Czech, Dong & Sully; ....)



# Spacetime Geometry = Entanglement

- Bekenstein-Hawking formula: spacetime geometry encodes  $S_{\text{BH}}$
- black hole entropy is entanglement entropy (Sorkin, ....)
- use BH formula for holographic entanglement entropy (Ryu & Takayanagi; ....)
- connectivity of spacetime requires entanglement (van Raamsdonk)
- spacetime entanglement conjecture (Bianchi & RM)
- AdS spacetime as a tensor network (MERA) (Swingle, Vidal, ....)
- “ER = EPR” conjecture (Maldacena & Susskind)
- hole-ographic spacetime (Balasubramanian, Chowdhury, Czech, de B; RM, Rao & Sugishita; Czech, Dong)



**spacetime provides both the stage for physical phenomena and the agent which manifests gravitational dynamics**

# Gravitational Dynamics from Entanglement:

(Lashkari, McDermott & Van Raamsdonk; Swingle & Van Raamsdonk;  
Faulkner, Guica, Hartman, RM & Van Raamsdonk)

- entanglement entropy:  $S(\rho_A) = -\text{tr}(\rho_A \log \rho_A)$
- make a small perturbation of state:  $\tilde{\rho} = \rho_A + \delta\rho$

# Gravitational Dynamics from Entanglement:

(Lashkari, McDermott & Van Raamsdonk; Swingle & Van Raamsdonk;  
Faulkner, Guica, Hartman, RM & Van Raamsdonk)

• entanglement entropy:  $S(\rho_A) = -\text{tr}(\rho_A \log \rho_A)$

• make a small perturbation of state:  $\tilde{\rho} = \rho_A + \delta\rho$

$$\longrightarrow \delta S = -\text{tr}(\delta\rho \log \rho_A) - \text{tr}(\rho_A \rho_A^{-1} \delta\rho) + O(\delta\rho^2)$$

# Gravitational Dynamics from Entanglement:

(Lashkari, McDermott & Van Raamsdonk; Swingle & Van Raamsdonk;  
Faulkner, Guica, Hartman, RM & Van Raamsdonk)

• entanglement entropy:  $S(\rho_A) = -\text{tr}(\rho_A \log \rho_A)$

• make a small perturbation of state:  $\tilde{\rho} = \rho_A + \delta\rho$

$$\begin{aligned} \longrightarrow \delta S &= -\text{tr}(\delta\rho \log \rho_A) - \underbrace{\text{tr}(\rho_A \rho_A^{-1} \delta\rho)}_{= \text{Tr}(\delta\rho) = 0} + O(\delta\rho^2) \end{aligned}$$

# Gravitational Dynamics from Entanglement:

(Lashkari, McDermott & Van Raamsdonk; Swingle & Van Raamsdonk;  
Faulkner, Guica, Hartman, RM & Van Raamsdonk)

• entanglement entropy:  $S(\rho_A) = -\text{tr}(\rho_A \log \rho_A)$

• make a small perturbation of state:  $\tilde{\rho} = \rho_A + \delta\rho$

$$\begin{aligned} \longrightarrow \delta S &= -\text{tr}(\delta\rho \log \rho_A) - \underbrace{\text{tr}(\rho_A \rho_A^{-1} \delta\rho)}_{= \text{Tr}(\delta\rho) = 0} + O(\delta\rho^2) \\ &= -\text{tr}(\delta\rho \log \rho_A) + O(\delta\rho^2) \end{aligned}$$



# Gravitational Dynamics from Entanglement:

(Lashkari, McDermott & Van Raamsdonk; Swingle & Van Raamsdonk;  
Faulkner, Guica, Hartman, RM & Van Raamsdonk)

- entanglement entropy:  $S(\rho_A) = -\text{tr}(\rho_A \log \rho_A)$

- make a small perturbation of state:  $\tilde{\rho} = \rho_A + \delta\rho$

$$\begin{aligned} \longrightarrow \delta S &= -\text{tr}(\delta\rho \log \rho_A) - \underbrace{\text{tr}(\rho_A \rho_A^{-1} \delta\rho)}_{= \text{Tr}(\delta\rho) = 0} + O(\delta\rho^2) \\ &= -\text{tr}(\delta\rho \log \rho_A) + O(\delta\rho^2) \end{aligned}$$

- modular (or entanglement) Hamiltonian:  $\rho_A = \exp(-H_A)$

# Gravitational Dynamics from Entanglement:

(Lashkari, McDermott & Van Raamsdonk; Swingle & Van Raamsdonk;  
Faulkner, Guica, Hartman, RM & Van Raamsdonk)

• entanglement entropy:  $S(\rho_A) = -\text{tr}(\rho_A \log \rho_A)$

• make a small perturbation of state:  $\tilde{\rho} = \rho_A + \delta\rho$

$$\begin{aligned} \longrightarrow \delta S &= -\text{tr}(\delta\rho \log \rho_A) - \underbrace{\text{tr}(\rho_A \rho_A^{-1} \delta\rho)}_{= \text{Tr}(\delta\rho) = 0} + O(\delta\rho^2) \\ &= -\text{tr}(\delta\rho \log \rho_A) + O(\delta\rho^2) \end{aligned}$$

• modular (or entanglement) Hamiltonian:  $\rho_A = \exp(-H_A)$   
(Blanco, Casini, Hung & RM)

$$\delta S_A = \delta \langle H_A \rangle$$

“1<sup>st</sup> law” of entanglement entropy

# Gravitational Dynamics from Entanglement:

(Lashkari, McDermott & Van Raamsdonk; Swingle & Van Raamsdonk;  
Faulkner, Guica, Hartman, RM & Van Raamsdonk)

- entanglement entropy:  $S(\rho_A) = -\text{tr}(\rho_A \log \rho_A)$

- make a small perturbation of state:  $\tilde{\rho} = \rho_A + \delta\rho$

$$\begin{aligned} \longrightarrow \delta S &= -\text{tr}(\delta\rho \log \rho_A) - \underbrace{\text{tr}(\rho_A \rho_A^{-1} \delta\rho)}_{= \text{Tr}(\delta\rho) = 0} + O(\delta\rho^2) \\ &= -\text{tr}(\delta\rho \log \rho_A) + O(\delta\rho^2) \end{aligned}$$

- modular (or entanglement) Hamiltonian:  $\rho_A = \exp(-H_A)$   
(Blanco, Casini, Hung & RM)

$$\delta S_A = \delta \langle H_A \rangle$$

“1<sup>st</sup> law” of entanglement entropy

- this is **the** 1<sup>st</sup> law for thermal state:  $\rho_A = \exp(-H/T)$

“1<sup>st</sup> law” of entanglement entropy:  $\delta S_A = \delta \langle H_A \rangle$

- generally  $H_A$  “**nonlocal mess**” and flow is **not geometric**

$$H_A = \int d^{d-1}x \gamma_1^{\mu\nu}(x) T_{\mu\nu} + \int d^{d-1}x \int d^{d-1}y \gamma_2^{\mu\nu;\rho\sigma}(x,y) T_{\mu\nu} T_{\rho\sigma} + \dots$$

→ hence usefulness of first law is very limited, in general

“1<sup>st</sup> law” of entanglement entropy:  $\delta S_A = \delta \langle H_A \rangle$

- generally  $H_A$  “**nonlocal mess**” and flow is **not geometric**

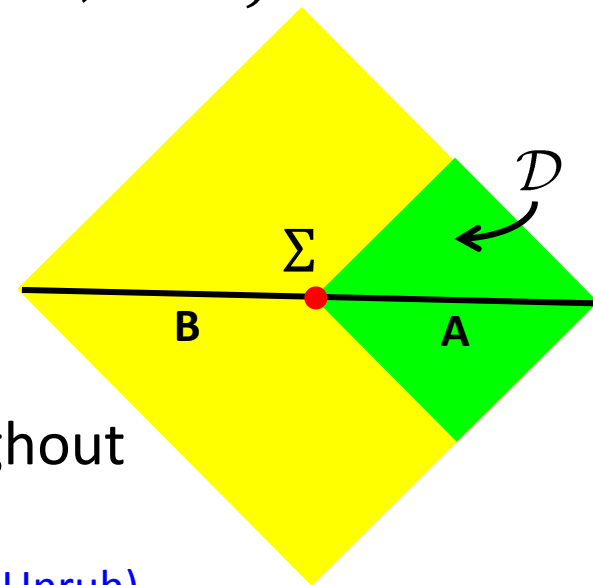
$$H_A = \int d^{d-1}x \gamma_1^{\mu\nu}(x) T_{\mu\nu} + \int d^{d-1}x \int d^{d-1}y \gamma_2^{\mu\nu;\rho\sigma}(x,y) T_{\mu\nu} T_{\rho\sigma} + \dots$$

→ hence usefulness of first law is very limited, in general

- famous exception: **Rindler wedge**
- any QFT in Minkowski vacuum; choose  $\Sigma = (x = 0, t = 0)$

$$H_A = 2\pi K \quad \leftarrow \text{boost generator}$$

$$= 2\pi \int_{A(x>0)} d^{d-2}y dx [x T_{tt}] + c'$$



- by causality,  $\rho_A$  and  $H_A$  describe physics throughout domain of dependence  $\mathcal{D}$

(Bisognano & Wichmann; Unruh)

“1<sup>st</sup> law” of entanglement entropy:  $\delta S_A = \delta \langle H_A \rangle$

- generally  $H_A$  “**nonlocal mess**” and flow is **not geometric**

$$H_A = \int d^{d-1}x \gamma_1^{\mu\nu}(x) T_{\mu\nu} + \int d^{d-1}x \int d^{d-1}y \gamma_2^{\mu\nu;\rho\sigma}(x,y) T_{\mu\nu} T_{\rho\sigma} + \dots$$

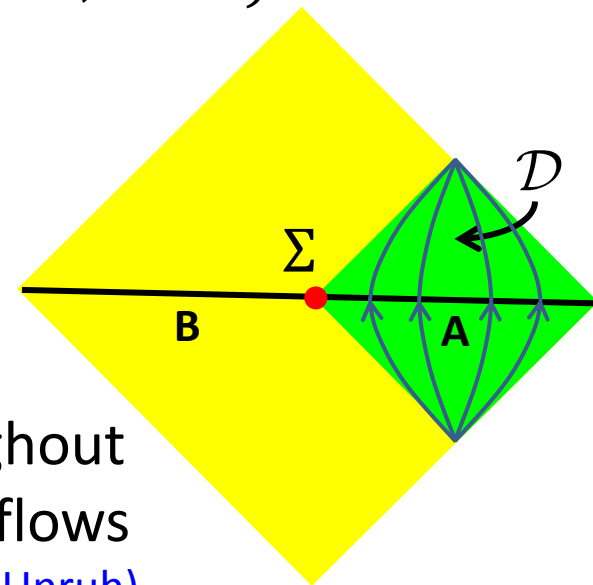
→ hence usefulness of first law is very limited, in general

- famous exception: **Rindler wedge**

- any QFT in Minkowski vacuum; choose  $\Sigma = (x = 0, t = 0)$

$$H_A = 2\pi K \quad \leftarrow \text{boost generator}$$

$$= 2\pi \int_{A(x>0)} d^{d-2}y dx [x T_{tt}] + c'$$



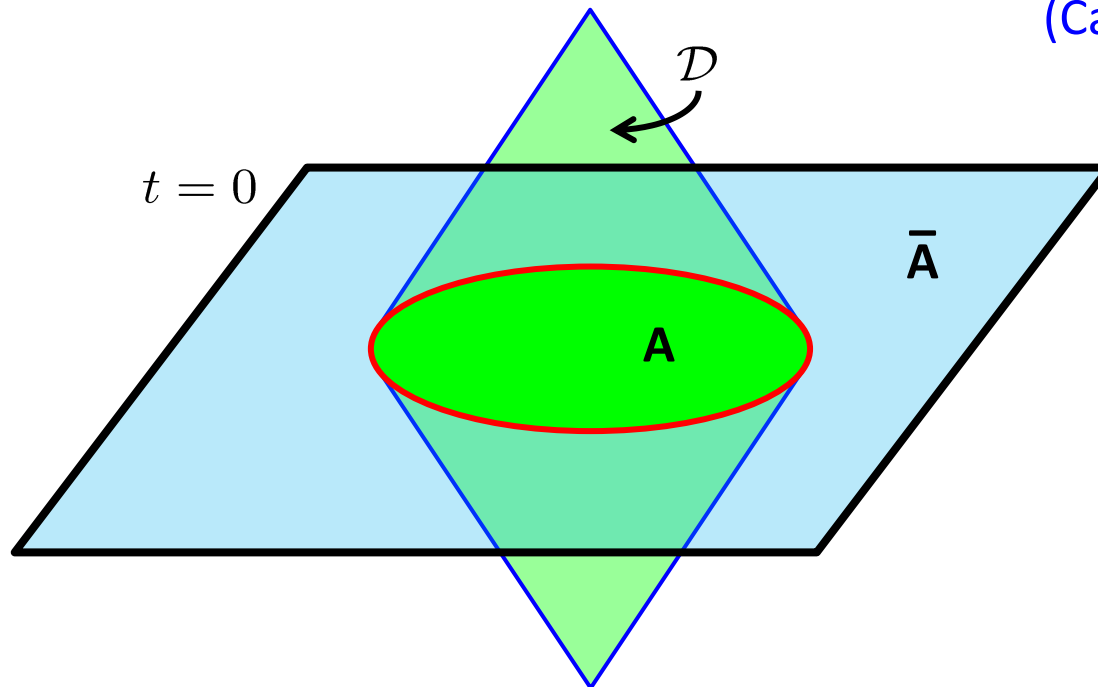
- by causality,  $\rho_A$  and  $H_A$  describe physics throughout domain of dependence  $\mathcal{D}$ ; eg, generate boost flows  
(Bisognano & Wichmann; Unruh)

“1<sup>st</sup> law” of entanglement entropy:  $\delta S_A = \delta \langle H_A \rangle$

- **another exception:** CFT in vacuum of d-dim. flat space and entangling surface which is  $S^{d-2}$  with radius R

$$H_A = 2\pi \int_A d^{d-1}y \frac{R^2 - |\vec{y}|^2}{2R} T_{tt}(\vec{y}) + c'$$

(Casini, Huerta & RM;  
Hislop & Longo)

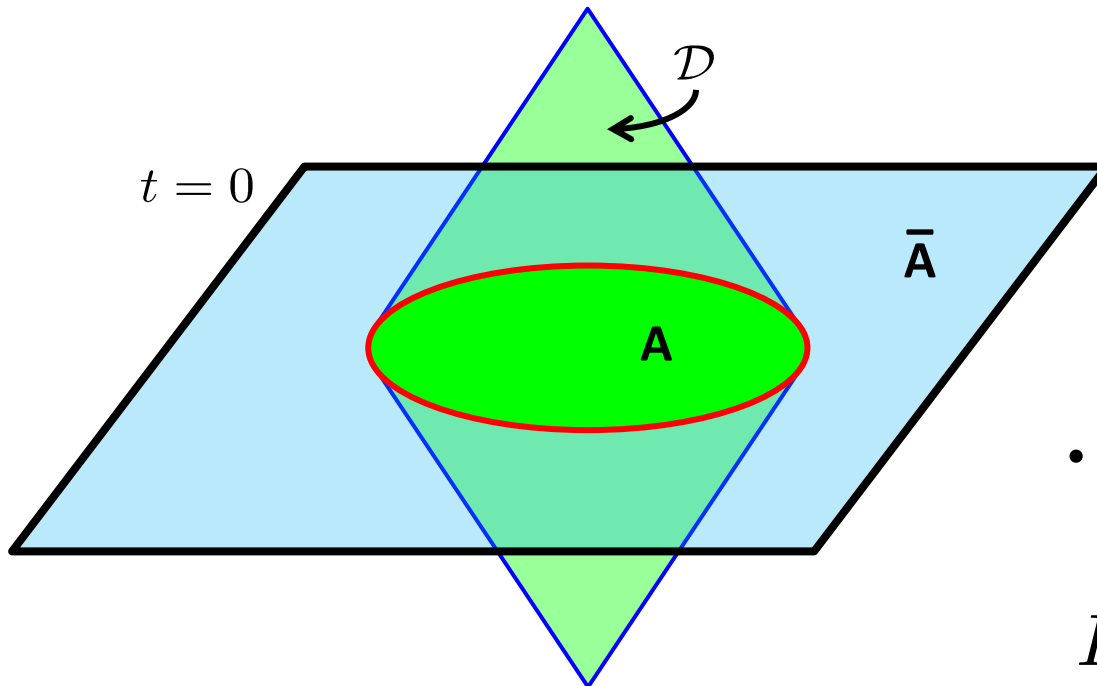


“1<sup>st</sup> law” of entanglement entropy:  $\delta S_A = \delta \langle H_A \rangle$

- **another exception:** CFT in vacuum of d-dim. flat space and entangling surface which is  $S^{d-2}$  with radius R

$$H_A = 2\pi \int_A d^{d-1}y \frac{R^2 - |\vec{y}|^2}{2R} T_{tt}(\vec{y}) + c'$$

(Casini, Huerta & RM;  
Hislop & Longo)



- construct with conformal transformation:

$$K = 2\pi \int_A d\Sigma^\mu T_{\mu\nu} K^\nu$$

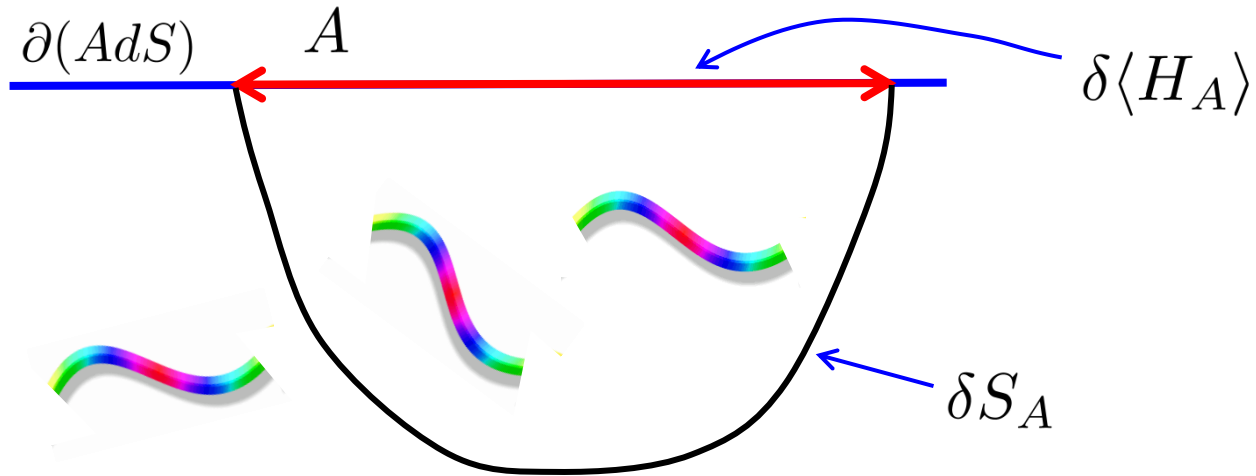


“1<sup>st</sup> law” of entanglement entropy:  $\delta S_A = \delta \langle H_A \rangle$

- $H_A$  has simple form for CFT and spherical entangling surface:

$$H_A = 2\pi \int_A d^{d-1}y \frac{R^2 - |\vec{y}|^2}{2R} T_{tt}(\vec{y}) + c'$$

- holographic realization:



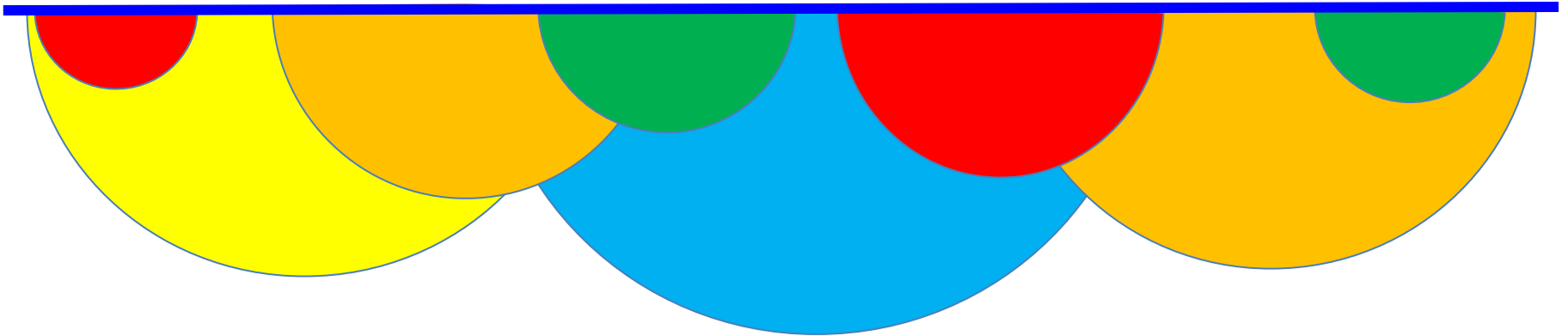
“1<sup>st</sup> law” of entanglement entropy:  $\delta S_A = \delta \langle H_A \rangle$

- $H_A$  has simple form for CFT and spherical entangling surface:

$$H_A = 2\pi \int_A d^{d-1}y \frac{R^2 - |\vec{y}|^2}{2R} T_{tt}(\vec{y}) + c'$$

- holographic realization:

$\partial(AdS)$



- apply 1<sup>st</sup> law for spheres of all sizes, positions and in all frames:

1<sup>st</sup> law of  $S_{EE}$



bulk geometry satisfies  
linearized Einstein eq's

“1<sup>st</sup> law” of entanglement entropy:  $\delta S_A = \delta \langle H_A \rangle$

- $H_A$  has simple form for CFT and spherical entangling surface:

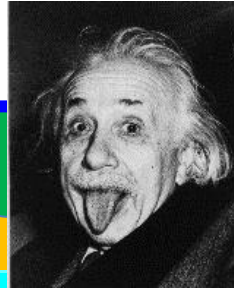
$$H_A = 2\pi \int_A d^{d-1}y \frac{R^2 - |\vec{y}|^2}{2R} T_{tt}(\vec{y}) + c'$$

- holographic realization:

$\partial(AdS)$

entanglement

~~spacetime~~ provides both the stage for physical phenomena and the agent which manifests gravitational dynamics



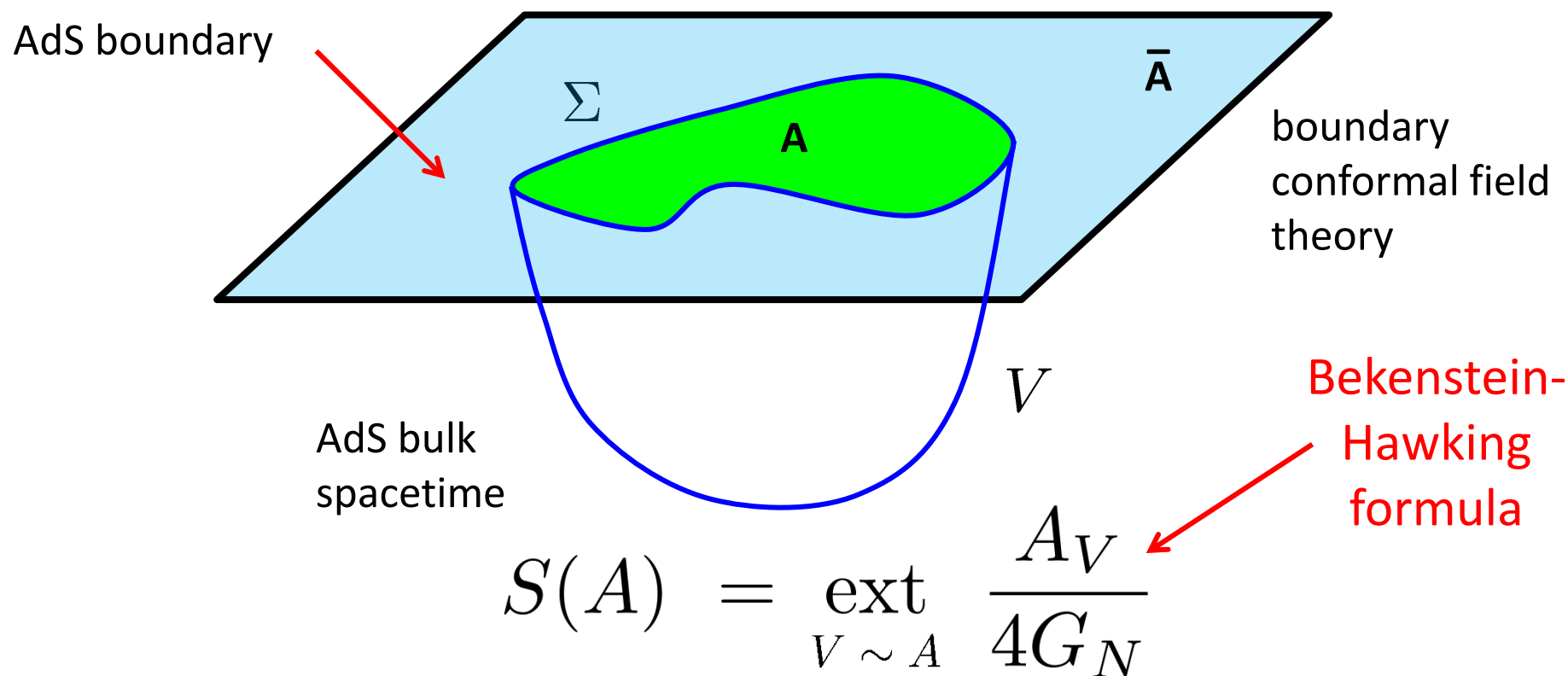
- apply 1<sup>st</sup> law for spheres of all sizes, positions and in all frames:

1<sup>st</sup> law of  $S_{EE}$



bulk geometry satisfies linearized Einstein eq's

# Holographic Entanglement Entropy:



• holographic EE teaches us lessons about (quantum) gravity, eg,

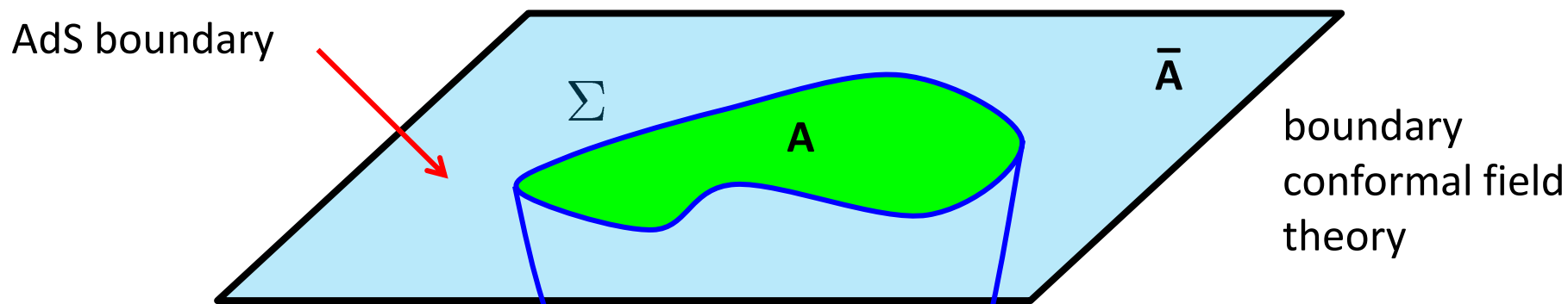
→ BH formula applies beyond black holes/horizons

→ connectivity of spacetime requires entanglement

(van Raamsdonk)

## Spacetime Geometry = Entanglement

# Holographic Entanglement Entropy:



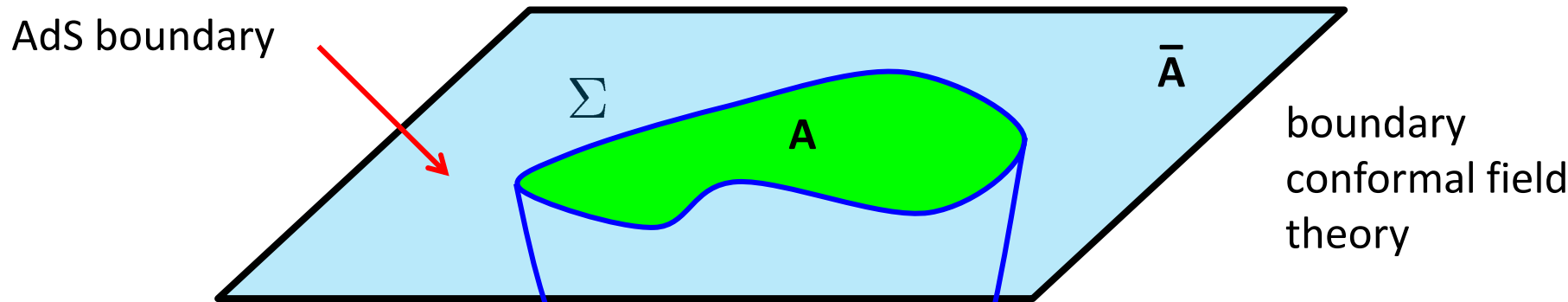
**Susskind: Entanglement is not enough!**

$$V \sim A \frac{c^3}{4G_N}$$

- holographic EE teaches us lessons about (quantum) gravity, eg,
  - BH formula applies beyond black holes/horizons
  - connectivity of spacetime requires entanglement (van Raamsdonk)

**Spacetime Geometry = Entanglement**

# Holographic Entanglement Entropy:



**Susskind: Entanglement is not enough!**

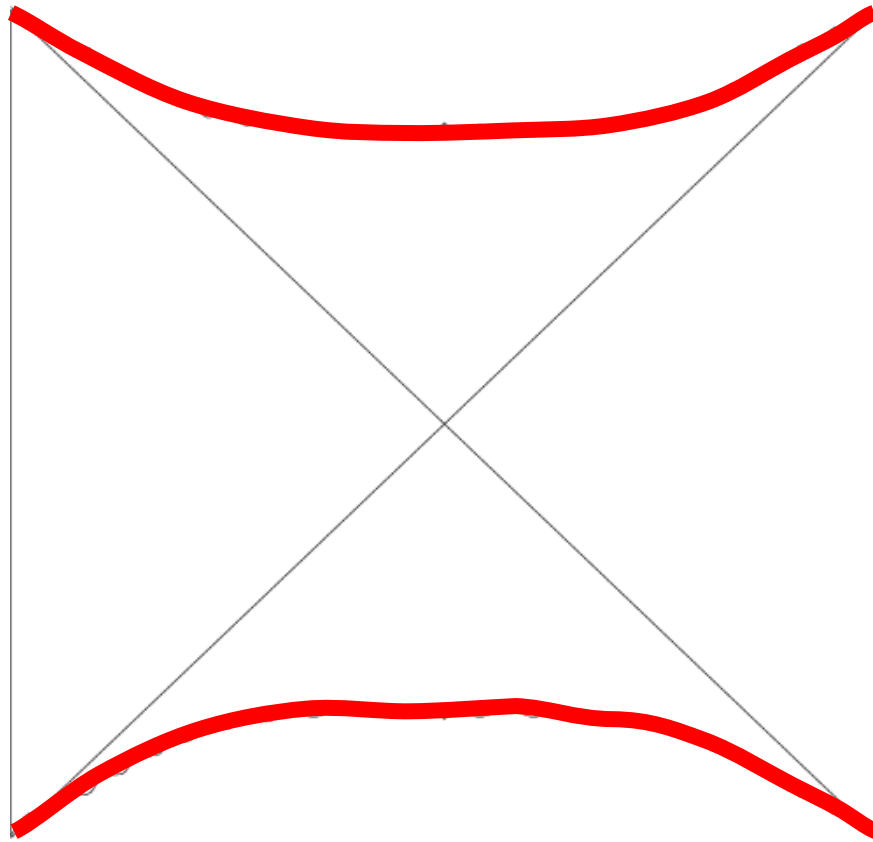
$$S(A) = \frac{c}{3} \ln \frac{V}{4G_N}$$

- holographic EE teaches us lessons about (quantum) gravity, eg,
  - BH formula applies beyond black holes/horizons
  - connectivity of spacetime requires entanglement (van Raamsdonk)

**Spacetime Geometry = Entanglement**

# Susskind: Entanglement<sup>Entropy</sup> is not enough!

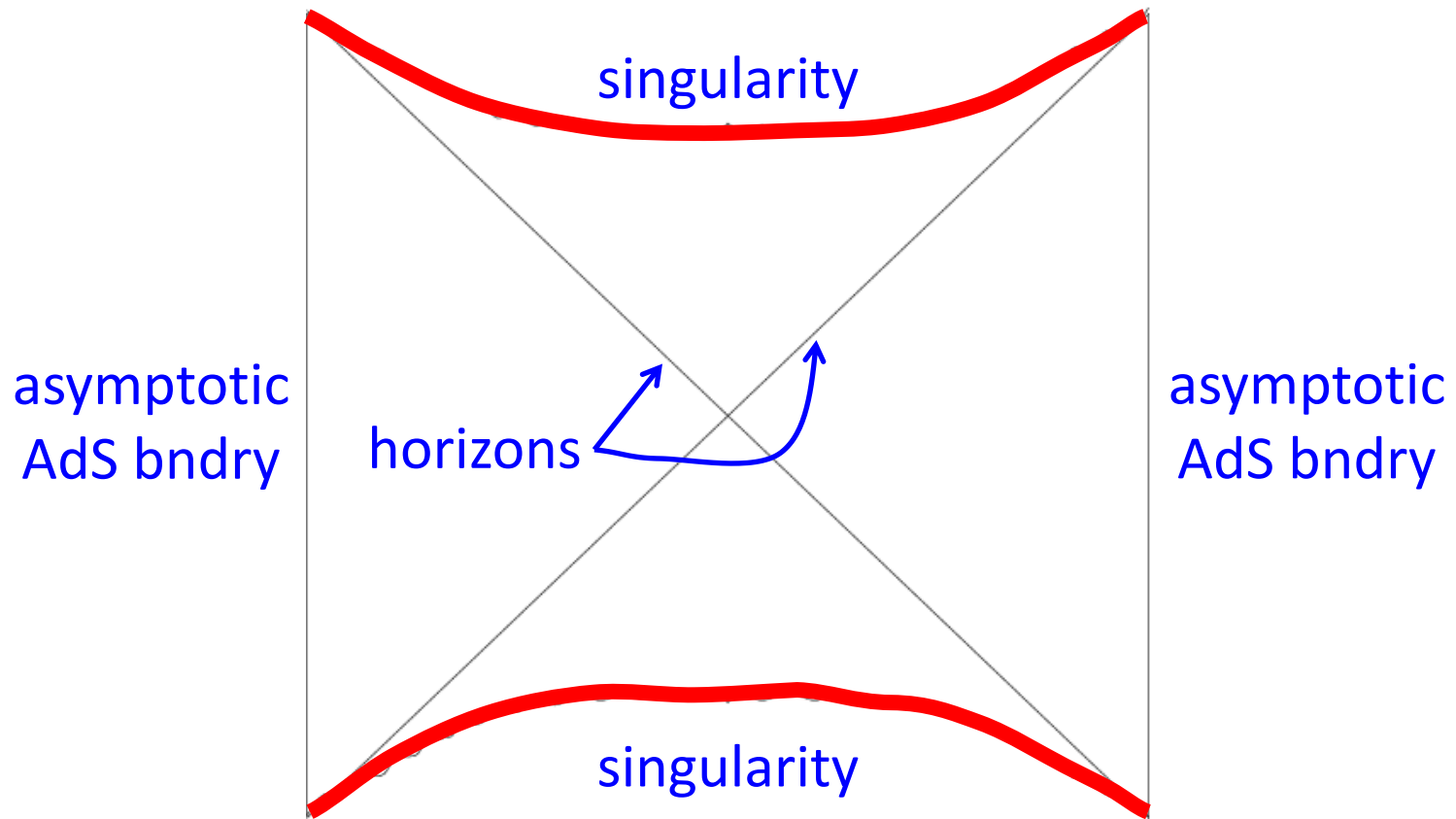
- “to understand the rich geometric structures that exist behind the horizon and which are predicted by general relativity.”



$$ds^2 = - \left( \frac{r^2}{L^2} + 1 - \frac{\mu}{r^{d-2}} \right) dt^2 + \frac{dr^2}{\frac{r^2}{L^2} + 1 - \frac{\mu}{r^{d-2}}} + r^2 d\Omega_{d-1}^2$$

# Susskind: Entanglement<sup>Entropy</sup> is not enough!

- “to understand the rich geometric structures that exist behind the horizon and which are predicted by general relativity.”

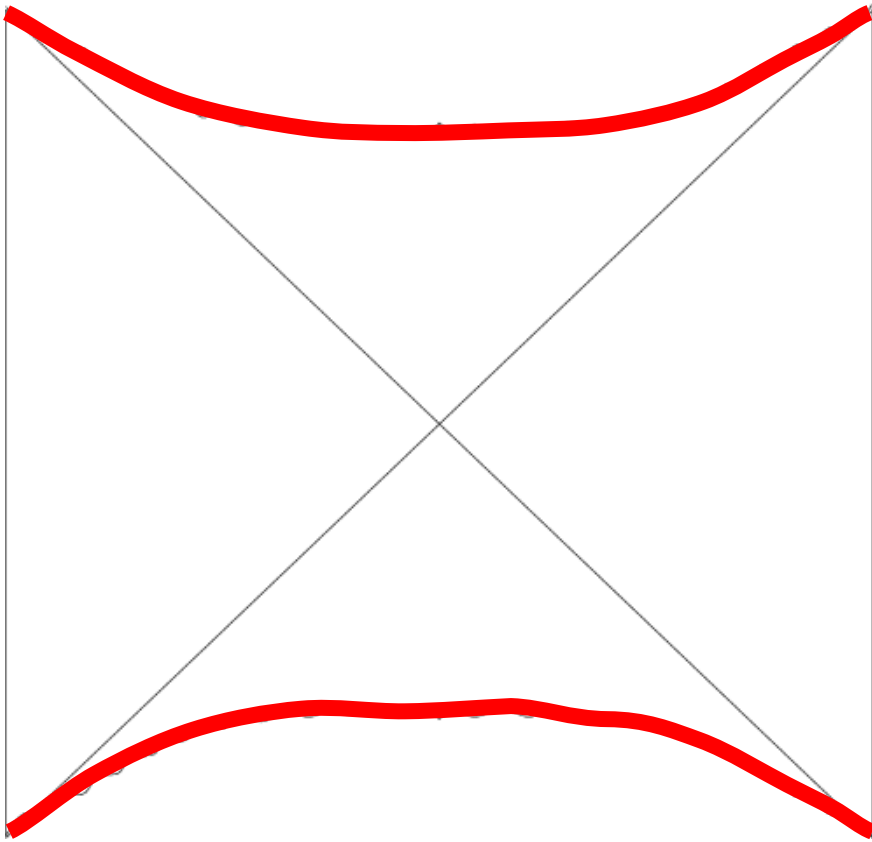


$$ds^2 = - \left( \frac{r^2}{L^2} + 1 - \frac{\mu}{r^{d-2}} \right) dt^2 + \frac{dr^2}{\frac{r^2}{L^2} + 1 - \frac{\mu}{r^{d-2}}} + r^2 d\Omega_{d-1}^2$$



# Susskind: Entanglement<sup>Entropy</sup> is not enough!

- “to understand the rich geometric structures that exist behind the horizon and which are predicted by general relativity.”

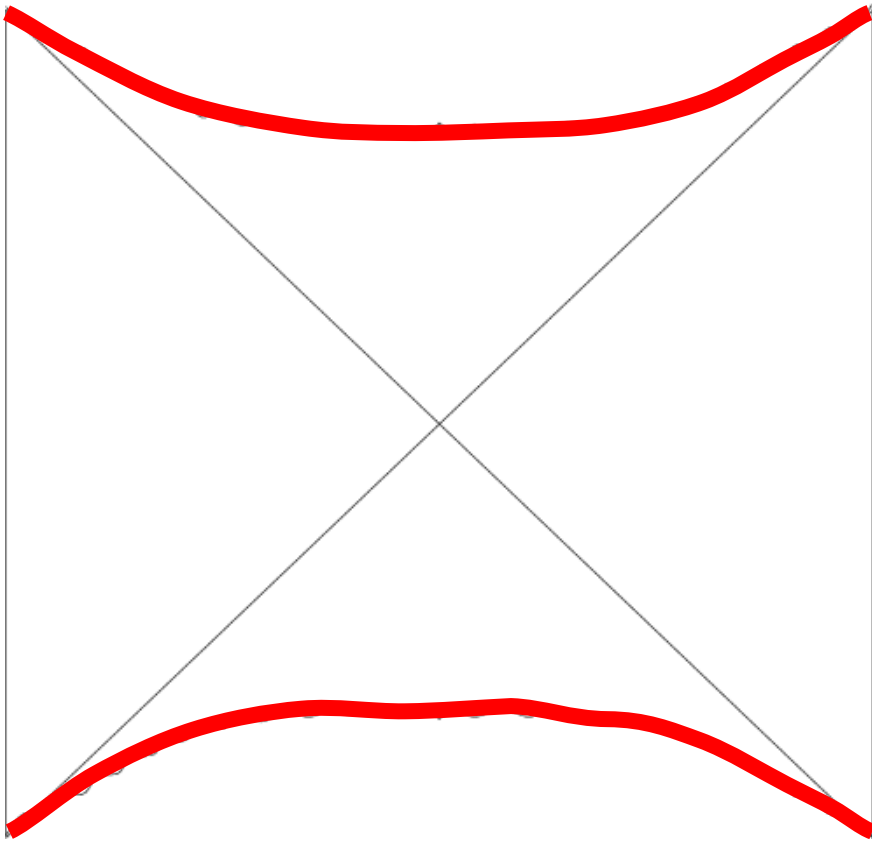


$$|\text{TFD}\rangle \simeq \sum_{\alpha} e^{-E_{\alpha}/(2T)} |E_{\alpha}\rangle_L |E_{\alpha}\rangle_R$$

- pure state:  $S_{EE} = 0$

# Susskind: Entanglement<sup>Entropy</sup> is not enough!

- “to understand the rich geometric structures that exist behind the horizon and which are predicted by general relativity.”



$$|\text{TFD}\rangle \simeq \sum_{\alpha} e^{-E_{\alpha}/(2T)} |E_{\alpha}\rangle_L |E_{\alpha}\rangle_R$$

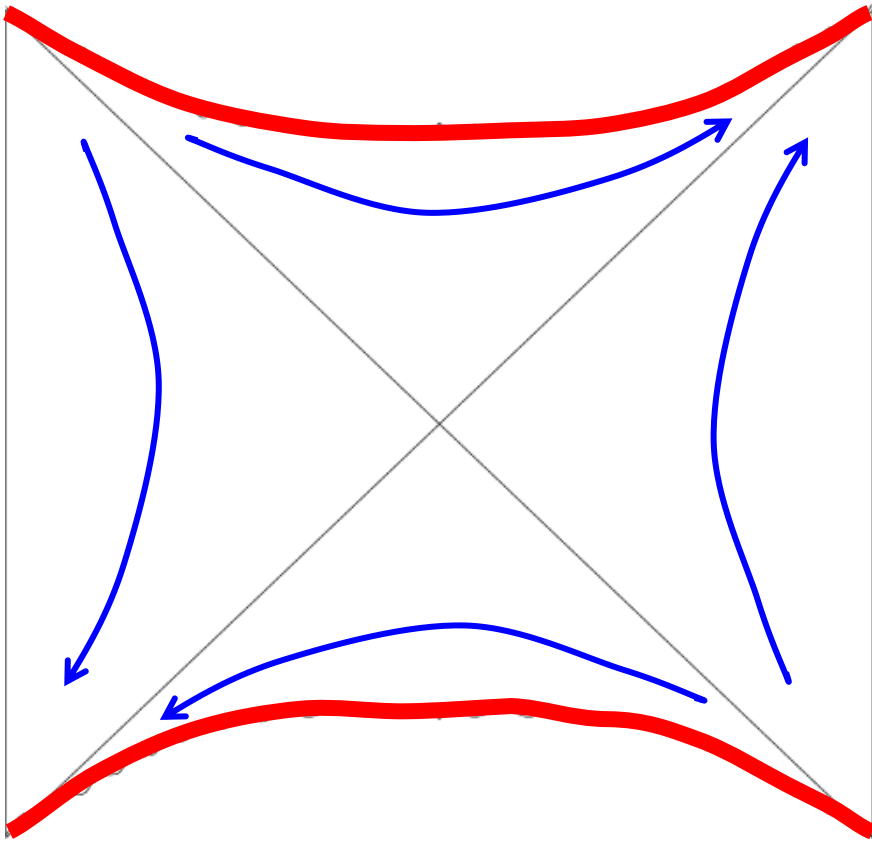
- pure state:  $S_{EE} = 0$

$$\begin{aligned} \rho_R &= \text{Tr}_L |\text{TFD}\rangle \langle \text{TFD}| \\ &\simeq \sum_{\alpha} e^{-E_{\alpha}/T} |E_{\alpha}\rangle_R \langle E_{\alpha}|_R \end{aligned}$$

- mixed state:  $S_{EE} = A_H/4G$

# Susskind: Entanglement<sup>Entropy</sup> is not enough!

- “to understand the rich geometric structures that exist behind the horizon and which are predicted by general relativity.”



$$|\text{TFD}\rangle \simeq \sum_{\alpha} e^{-E_{\alpha}/(2T)} |E_{\alpha}\rangle_L |E_{\alpha}\rangle_R$$

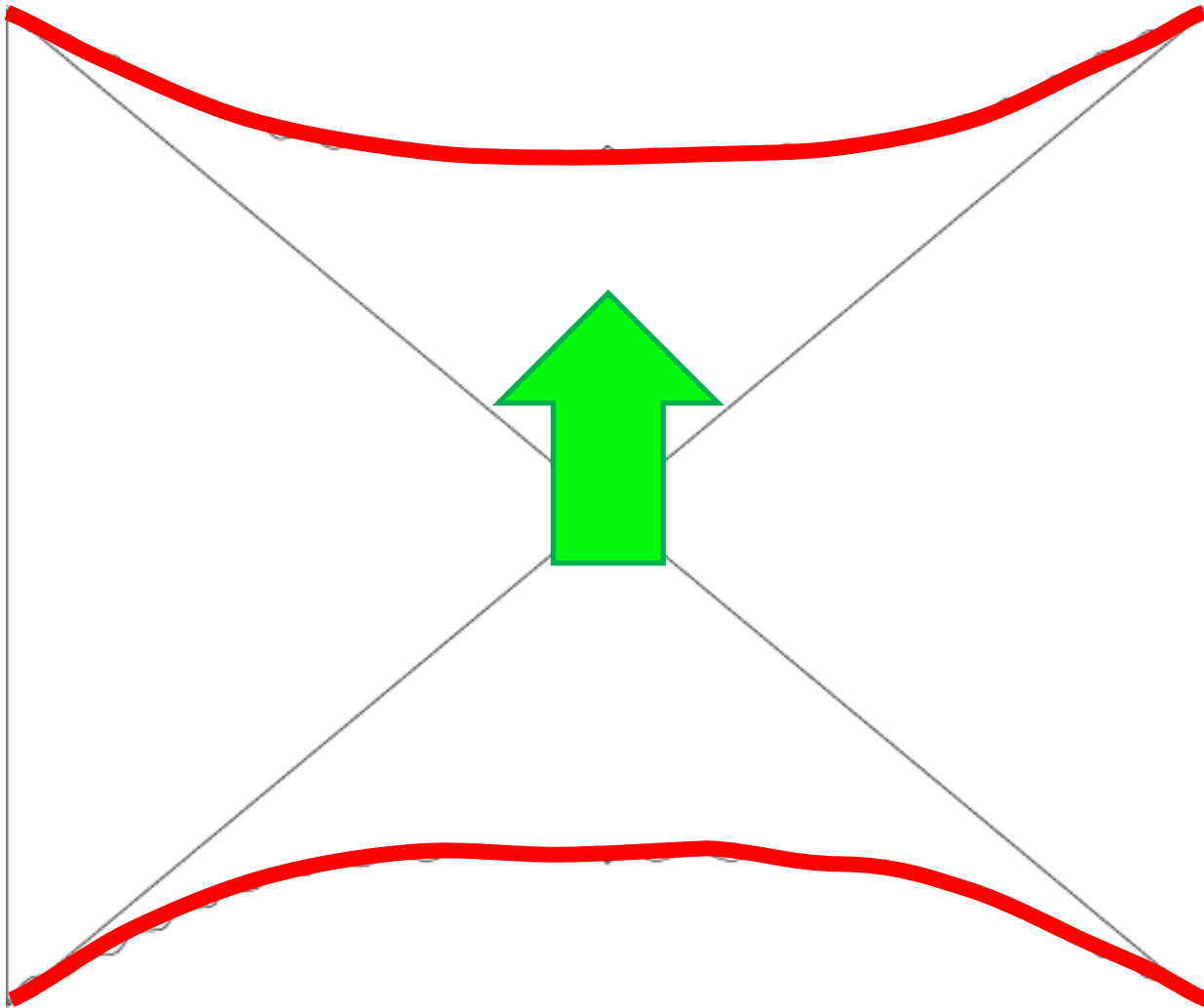
- pure state:  $S_{EE} = 0$

$$\begin{aligned} \rho_R &= \text{Tr}_L |\text{TFD}\rangle \langle \text{TFD}| \\ &\simeq \sum_{\alpha} e^{-E_{\alpha}/T} |E_{\alpha}\rangle_R \langle E_{\alpha}|_R \end{aligned}$$

- mixed state:  $S_{EE} = A_H/4G$

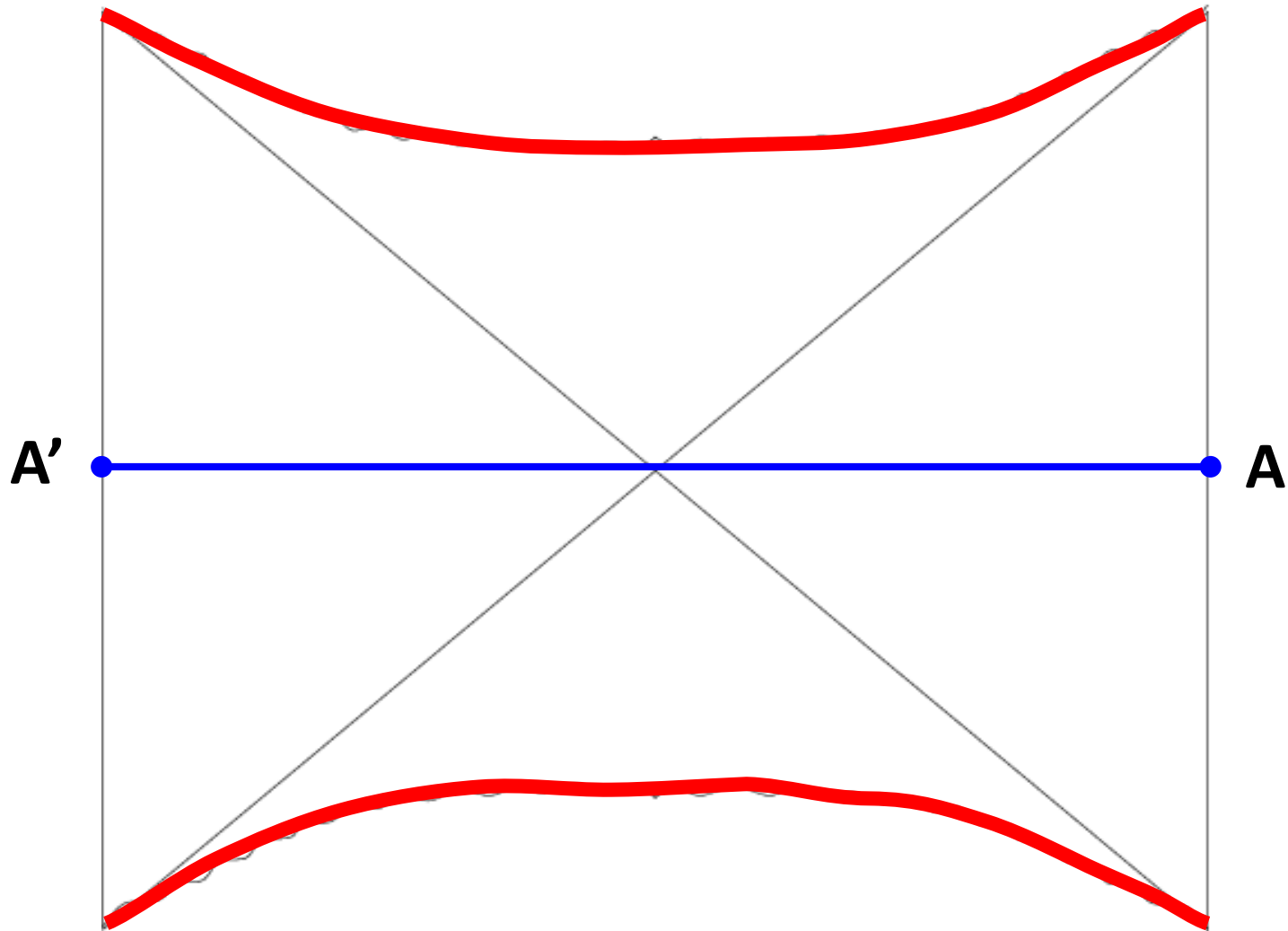
# Susskind: Entanglement<sup>Entropy</sup> is not enough!

- “to understand the rich geometric structures that exist behind the horizon and which are predicted by general relativity.”



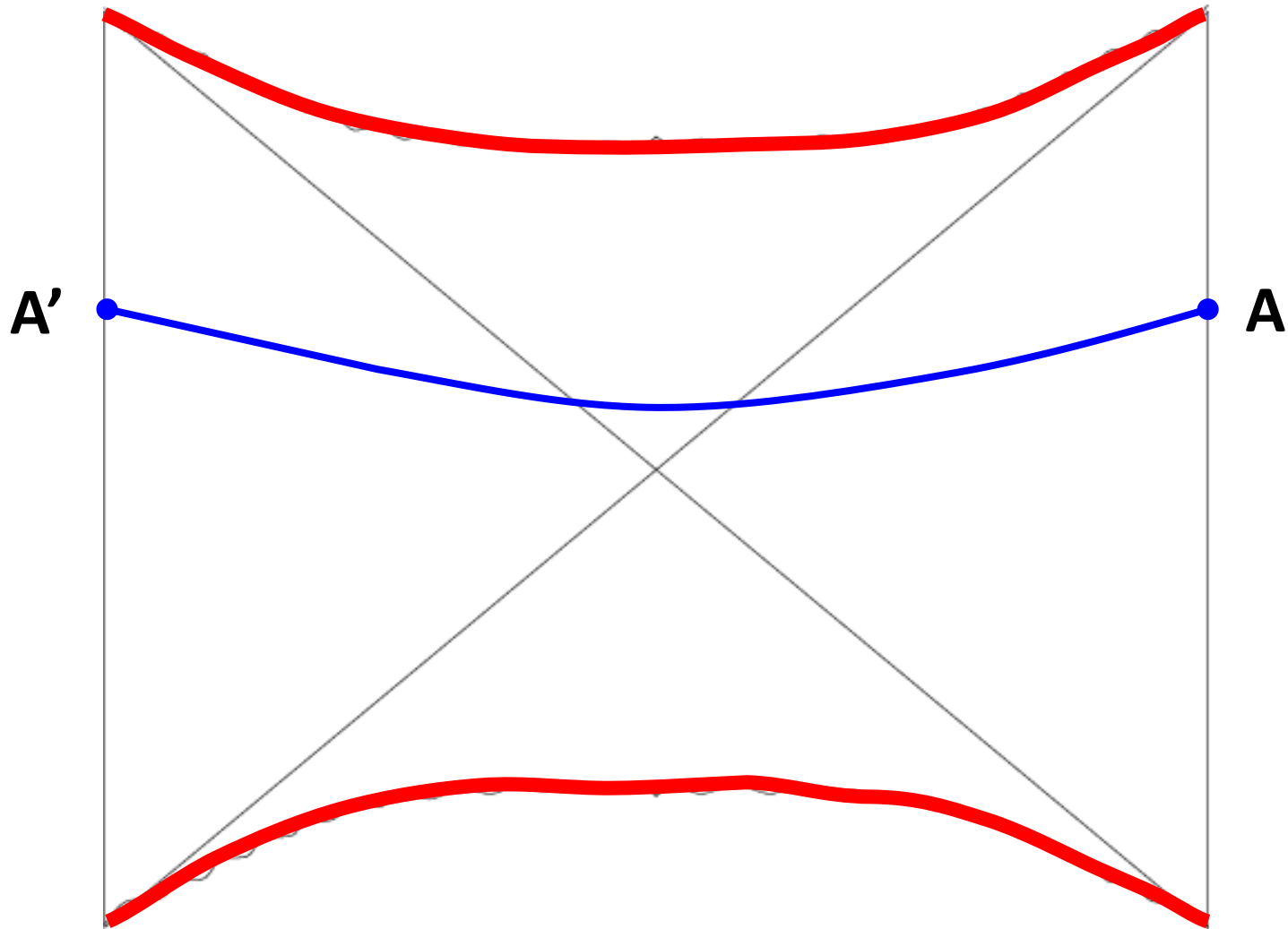
# Susskind: Entanglement<sup>Entropy</sup> is not enough!

- “to understand the rich geometric structures that exist behind the horizon and which are predicted by general relativity.”



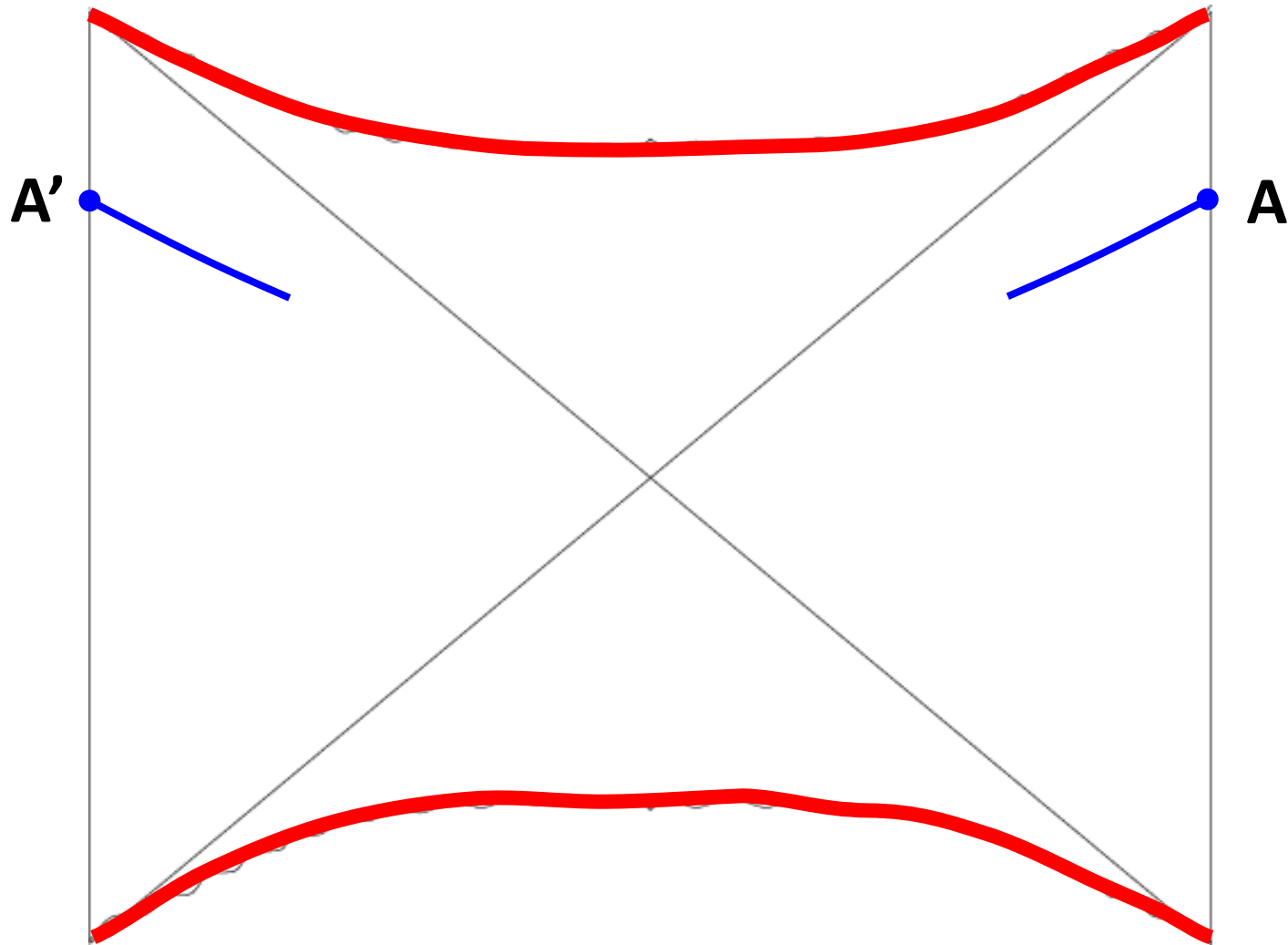
# Susskind: Entanglement<sup>Entropy</sup> is not enough!

- “to understand the rich geometric structures that exist behind the horizon and which are predicted by general relativity.”



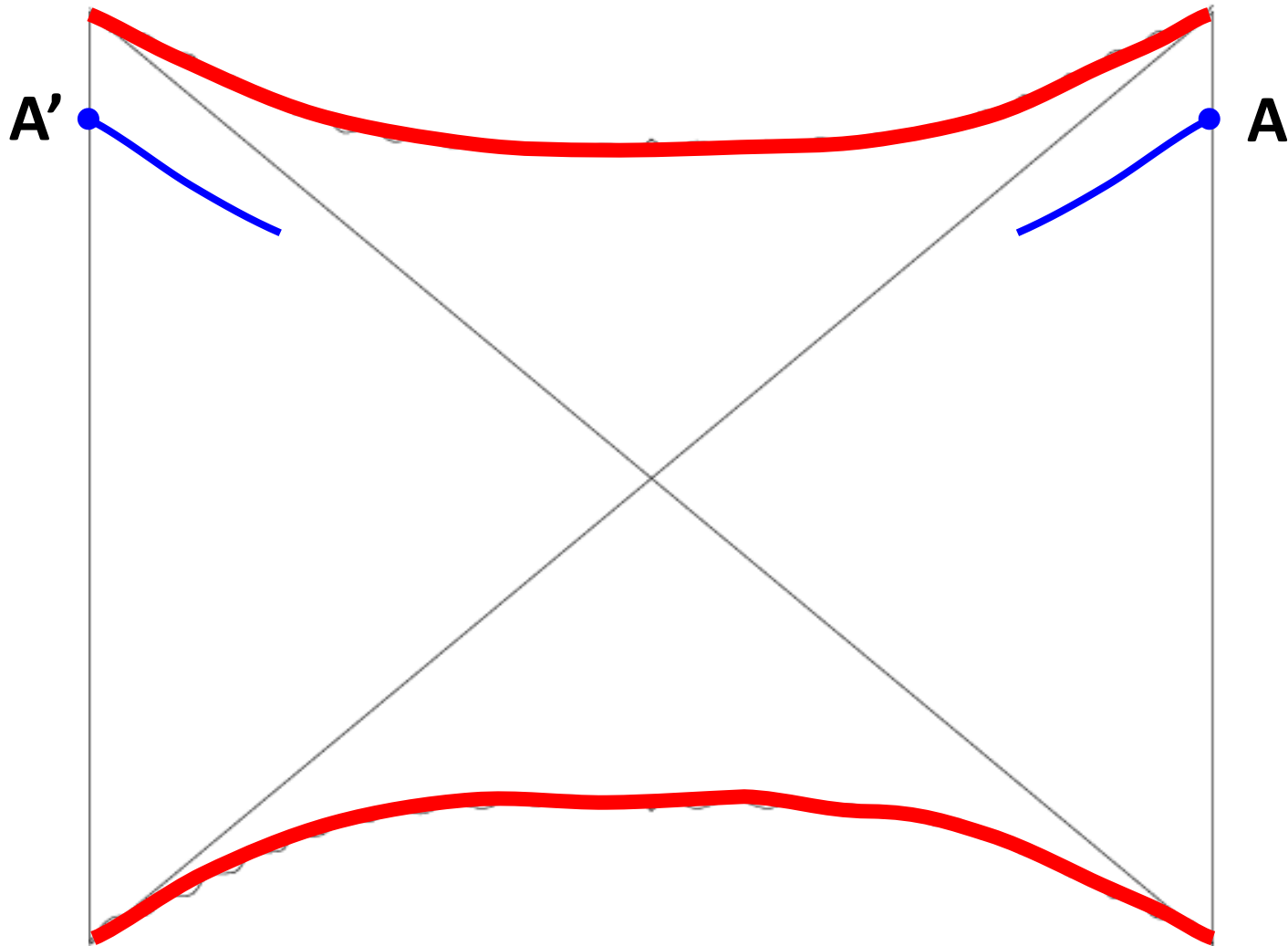
# Susskind: Entanglement<sup>Entropy</sup> is not enough!

- “to understand the rich geometric structures that exist behind the horizon and which are predicted by general relativity.”



# Susskind: Entanglement<sup>Entropy</sup> is not enough!

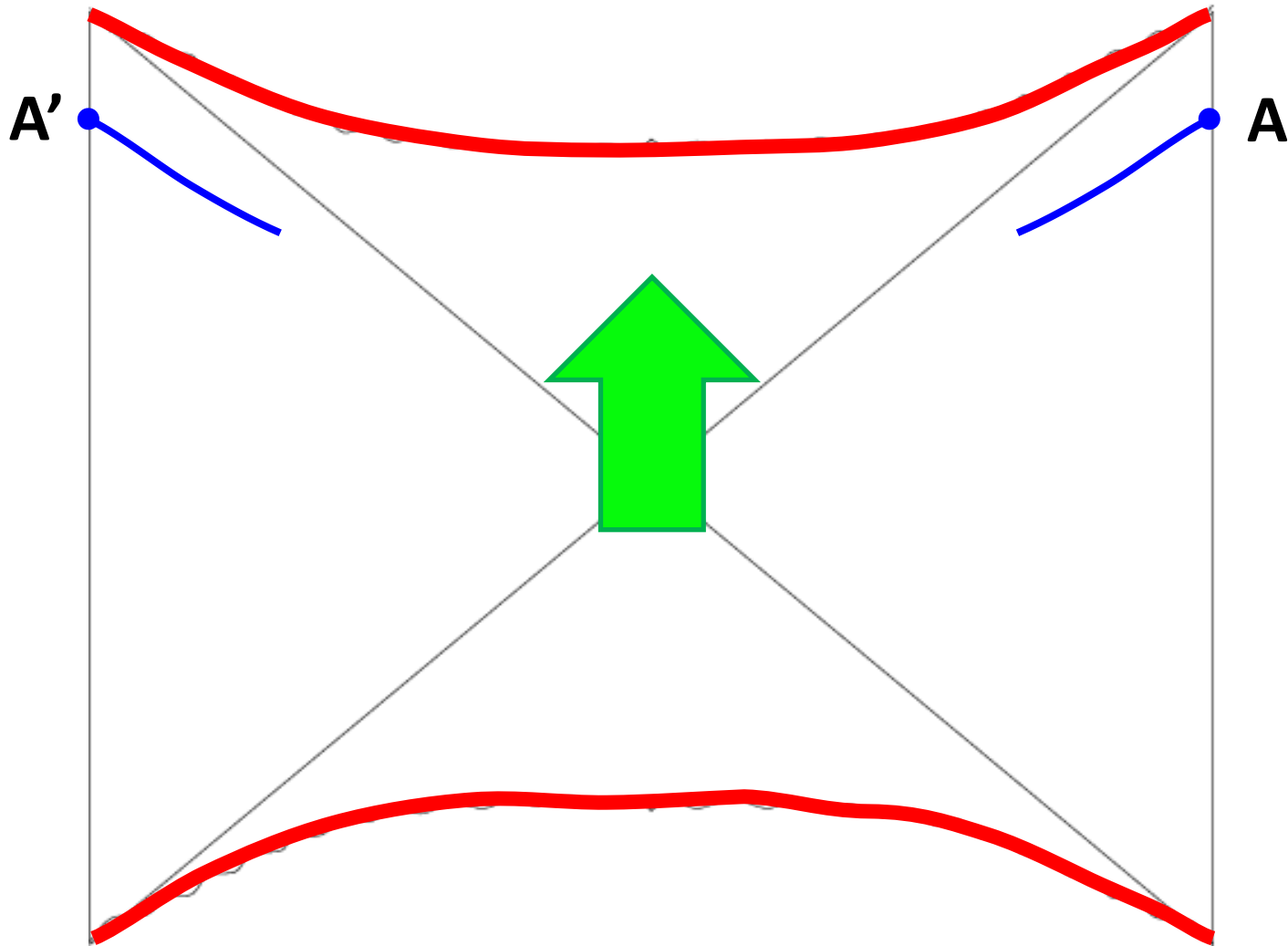
- “to understand the rich geometric structures that exist behind the horizon and which are predicted by general relativity.”





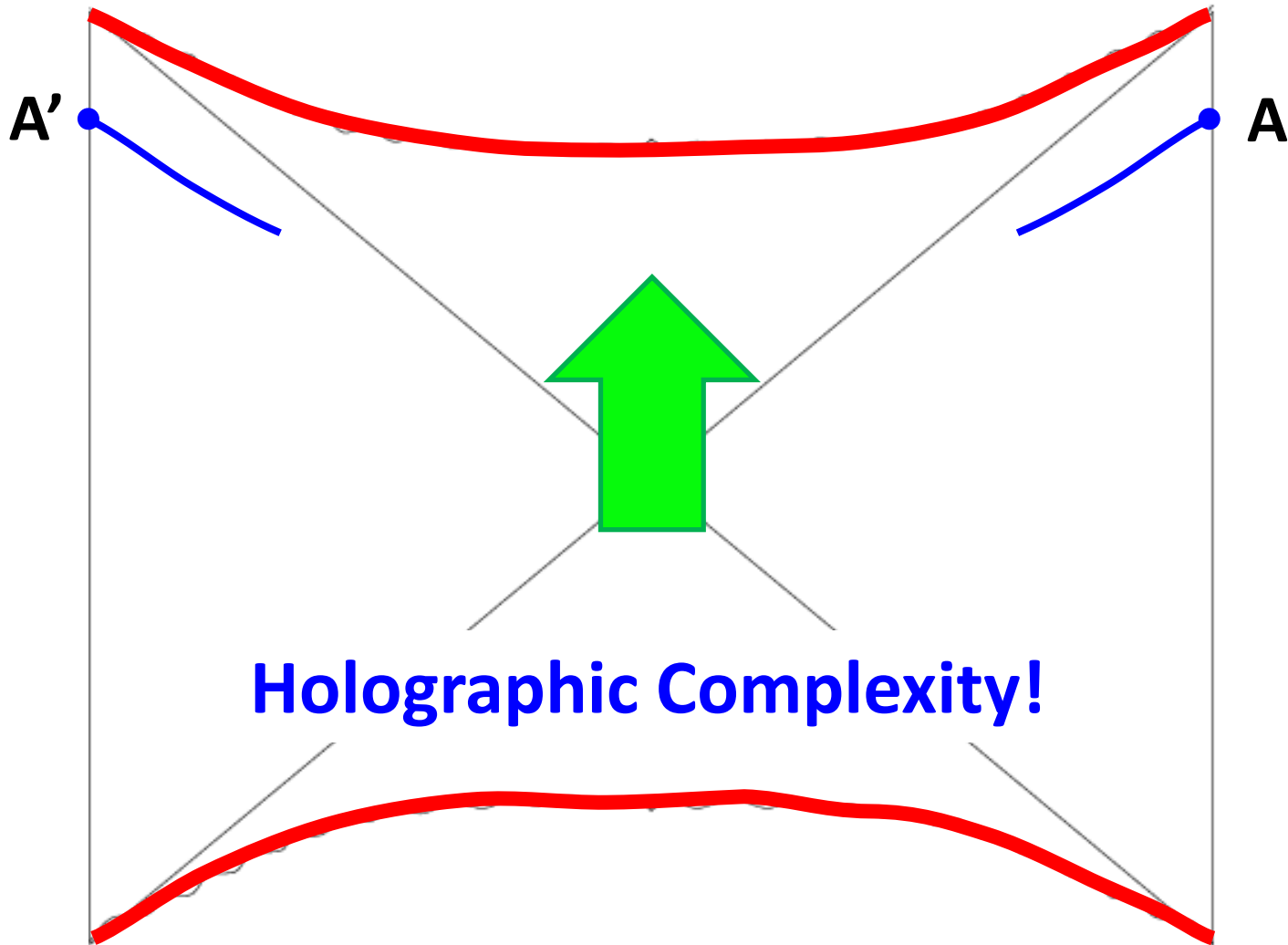
# Susskind: Entanglement<sup>Entropy</sup> is not enough!

- “to understand the rich geometric structures that exist behind the horizon and which are predicted by general relativity.”



# Susskind: Entanglement<sup>Entropy</sup> is not enough!

- “to understand the rich geometric structures that exist behind the horizon and which are predicted by general relativity.”



## Complexity?

- recall  $S_{EE}$  only probes the **eigenvalues** of the density matrix

$$S_{EE} = -\text{Tr} [\rho_A \log \rho_A] = -\sum \lambda_i \log \lambda_i$$

## Complexity?

- recall  $S_{EE}$  only probes the **eigenvalues** of the density matrix

$$S_{EE} = -\text{Tr} [\rho_A \log \rho_A] = -\sum \lambda_i \log \lambda_i$$

- would like a new probe “sensitive to phases”

$$|\text{TFD}\rangle \simeq \sum_{\alpha} e^{-E_{\alpha}/(2T) - iE_{\alpha}(t_L + t_R)} |E_{\alpha}\rangle_L |E_{\alpha}\rangle_R$$

## Complexity:

- computational complexity: how difficult is it to implement a task? eg, how difficult is it to prepare a particular state?

## Complexity:

- computational complexity: how difficult is it to implement a task? eg, how difficult is it to prepare a particular state?
- quantum circuit model:

$$|\psi\rangle = U |\psi_0\rangle$$

## Complexity:

- computational complexity: how difficult is it to implement a task? eg, how difficult is it to prepare a particular state?
- quantum circuit model:

$$|\psi\rangle = U |\psi_0\rangle$$

simple reference state

eg,  $|00000 \dots 0\rangle$

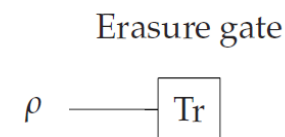
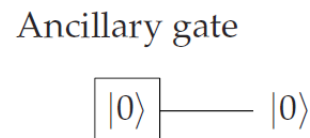
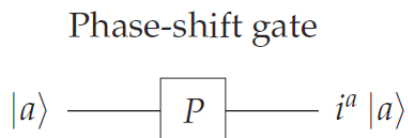
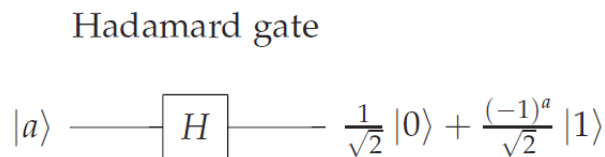
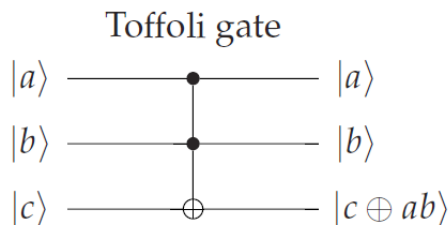
# Complexity:

- computational complexity: how difficult is it to implement a task? eg, how difficult is it to prepare a particular state?
- quantum circuit model:

$$|\psi\rangle = U |\psi_0\rangle$$

unitary operator  
built from set of  
simple gates

simple reference state  
eg,  $|00000 \dots 0\rangle$





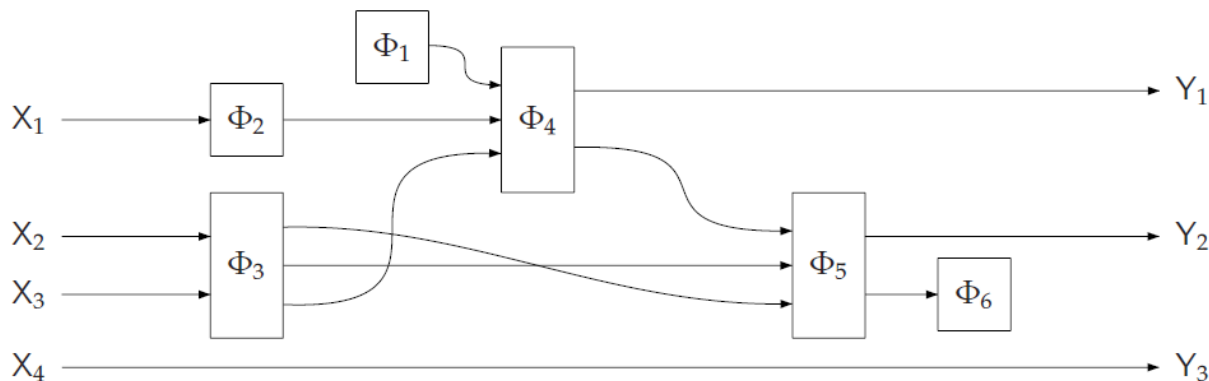
# Complexity:

- computational complexity: how difficult is it to implement a task? eg, how difficult is it to prepare a particular state?
- quantum circuit model:

$$|\psi\rangle = U |\psi_0\rangle$$

unitary operator  
built from set of  
simple gates

simple reference state  
eg,  $|00000 \dots 0\rangle$



## Complexity:

- computational complexity: how difficult is it to implement a task? eg, how difficult is it to prepare a particular state?
- quantum circuit model:

$$|\psi\rangle = U |\psi_0\rangle$$

unitary operator  $\rightarrow$  built from set of simple gates

$\leftarrow$  simple reference state  
eg,  $|00000 \dots 0\rangle$

tolerance:  $\left| |\psi\rangle - |\psi\rangle_{\text{Target}} \right|^2 \cdot \varepsilon$

## Complexity:

- computational complexity: how difficult is it to implement a task? eg, how difficult is it to prepare a particular state?
- quantum circuit model:

$$|\psi\rangle = U |\psi_0\rangle$$

unitary operator  $\xrightarrow{\quad}$  simple reference state  
built from set of  
simple gates eg,  $|00000 \dots 0\rangle$

tolerance:  $\left| |\psi\rangle - |\psi\rangle_{\text{Target}} \right|^2 \cdot \varepsilon$

- **complexity** = minimum number of gates required to prepare the desired state

## Complexity:

- computational complexity: how difficult is it to implement a task? eg, how difficult is it to prepare a particular state?
- quantum circuit model:

$$|\psi\rangle = U |\psi_0\rangle$$

unitary operator  $\xrightarrow{\quad}$  simple reference state  
built from set of eg,  $|00000 \dots 0\rangle$   
simple gates

tolerance:  $\left| |\psi\rangle - |\psi\rangle_{\text{Target}} \right|^2 \cdot \varepsilon$

- **complexity** = minimum number of gates required to prepare the desired state
- does the answer depend on the choices?? **YES!!**
- compare to “circuit depth” for spin chain

## Complexity:

- computational complexity: how difficult is it to implement a task? eg, how difficult is it to prepare a particular state?
- quantum circuit model:

$$|\psi\rangle = U |\psi_0\rangle$$

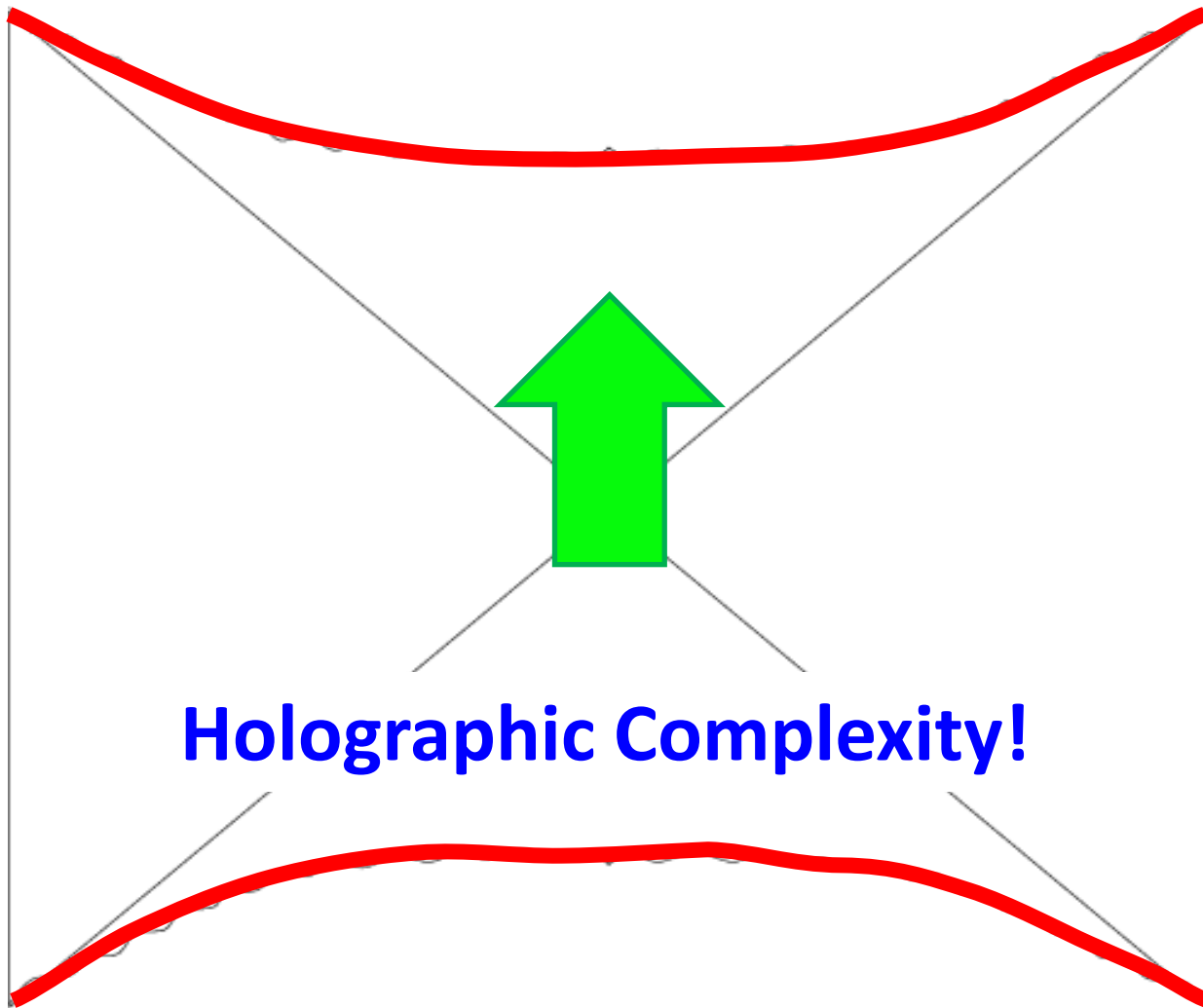
unitary operator  $\xrightarrow{\quad}$  simple reference state  
built from set of eg,  $|00000 \dots 0\rangle$   
simple gates

tolerance:  $\left| |\psi\rangle - |\psi\rangle_{\text{Target}} \right|^2 \cdot \varepsilon$

- **complexity** = minimum number of gates required to prepare the desired state
- does the answer depend on the choices?? **YES!!**
- but what does this really mean in quantum field theory? **???**

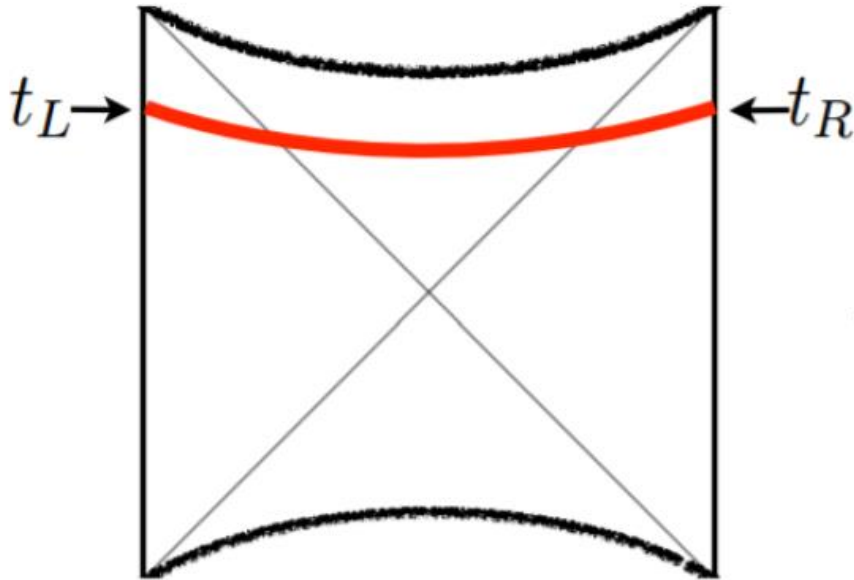
# Susskind: Entanglement<sup>Entropy</sup> is not enough!

- “to understand the rich geometric structures that exist behind the horizon and which are predicted by general relativity.”



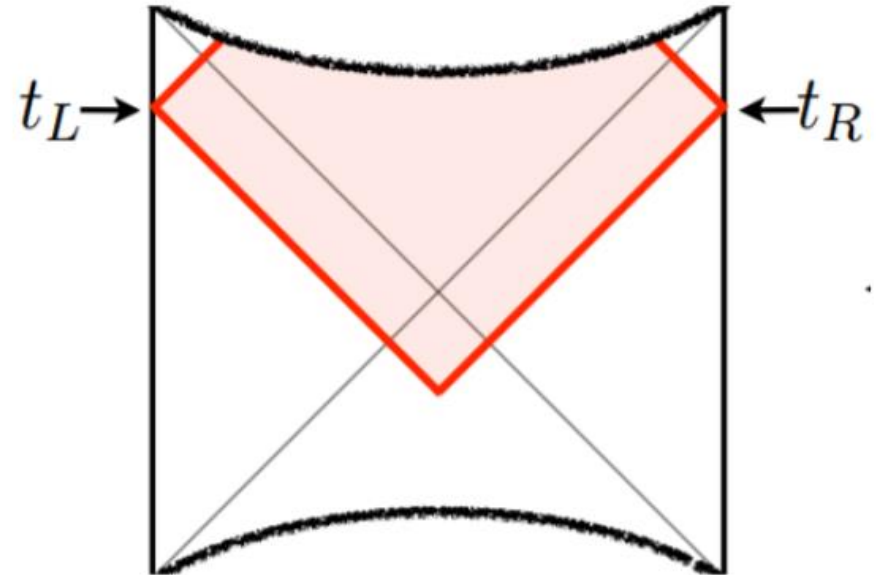
# A Tale of Two Dualities: Holographic Complexity

Complexity = Volume



$$\mathcal{C}_V(\Sigma) = \max_{\Sigma=\partial\mathcal{B}} \left[ \frac{\mathcal{V}(\mathcal{B})}{G_N \ell} \right]$$

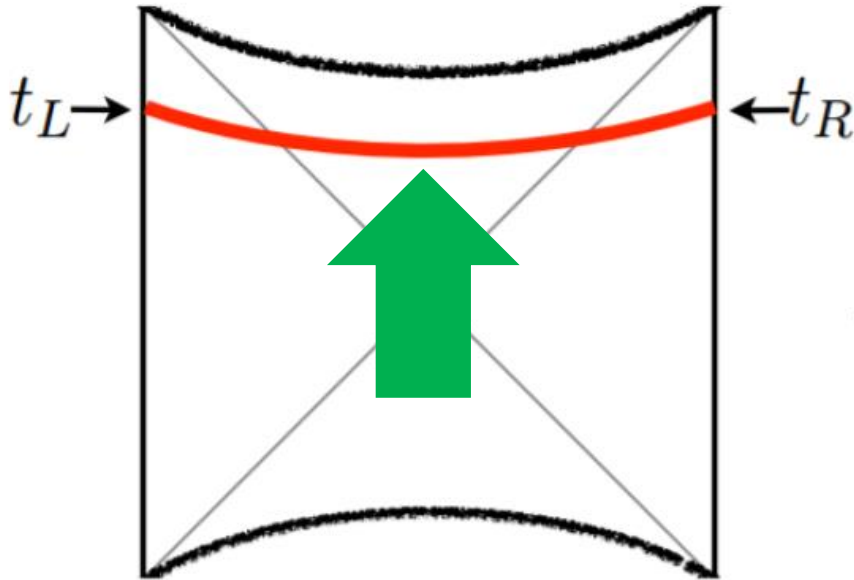
Complexity = Action



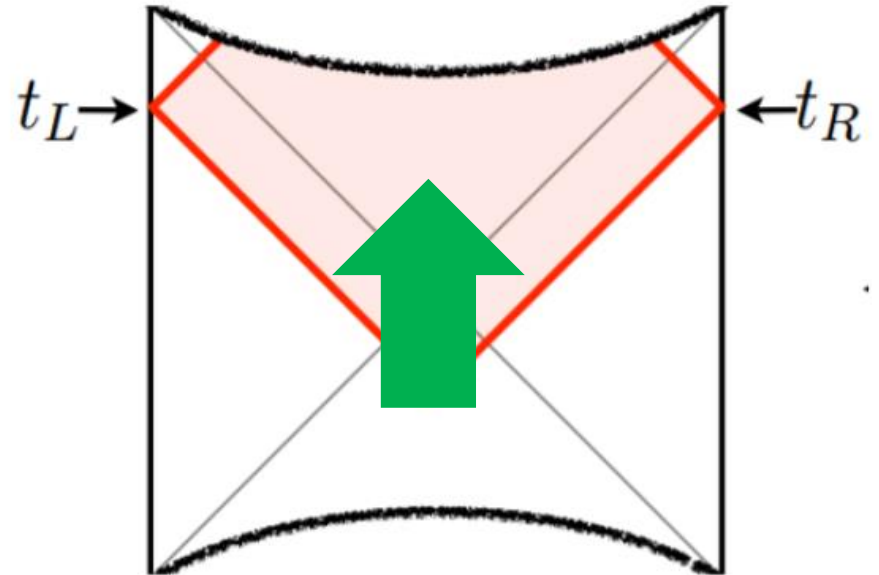
$$\mathcal{C}_A(\Sigma) = \frac{I_{\text{WDW}}}{\pi \hbar}$$

# A Tale of Two Dualities: Holographic Complexity

Complexity = Volume



Complexity = Action



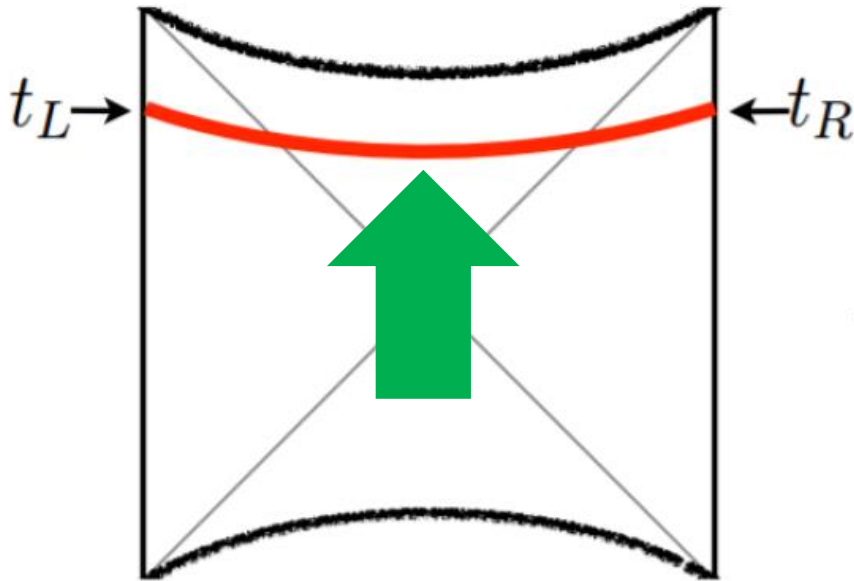
$$\mathcal{C}_V(\Sigma) = \max_{\Sigma=\partial\mathcal{B}} \left[ \frac{\mathcal{V}(\mathcal{B})}{G_N \ell} \right]$$

$$\mathcal{C}_A(\Sigma) = \frac{I_{\text{WDW}}}{\pi \hbar}$$

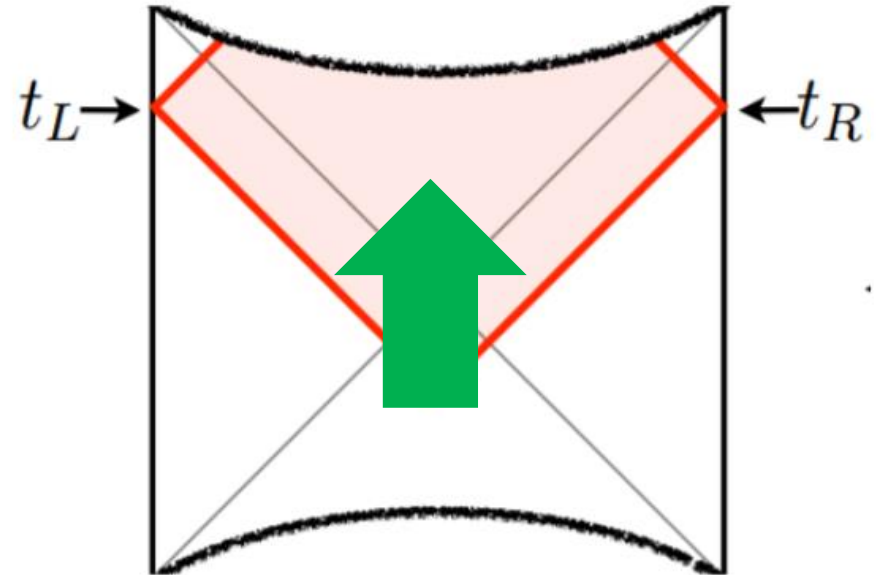


# A Tale of Two Dualities: Holographic Complexity

Complexity = Volume



Complexity = Action



$$\left. \frac{d\mathcal{C}_V}{dt} \right|_{t \rightarrow \infty} = \frac{8\pi}{d-1} M \quad (\text{planar})$$

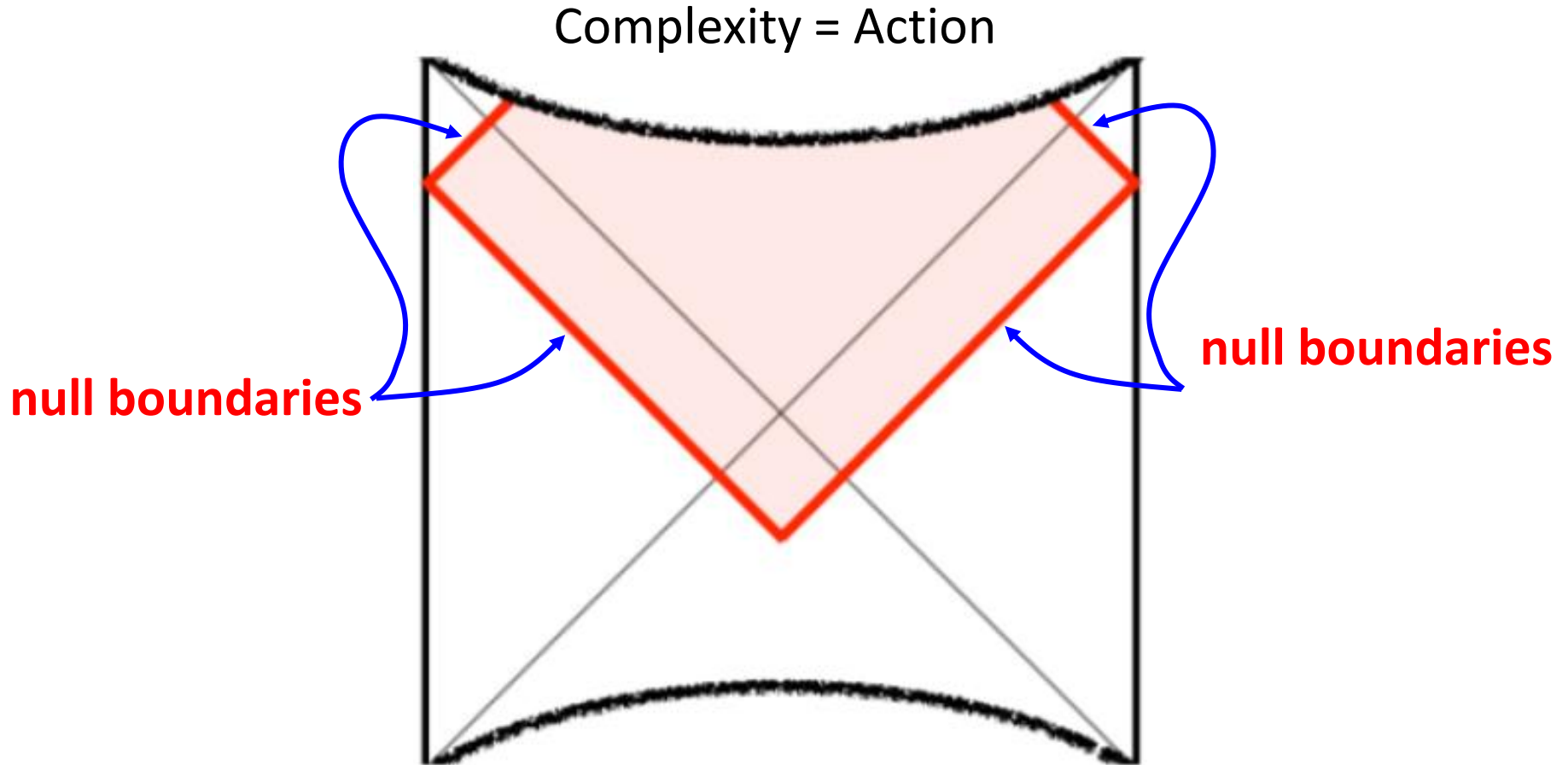
( $d$  = boundary dimension)

$$\left. \frac{d\mathcal{C}_A}{dt} \right|_{t \rightarrow \infty} = \frac{2M}{\pi}$$

(universal; Lloyd bound)

Team Lenny, including Brown, Roberts, Swingle, Stanford, Susskind & Zhao

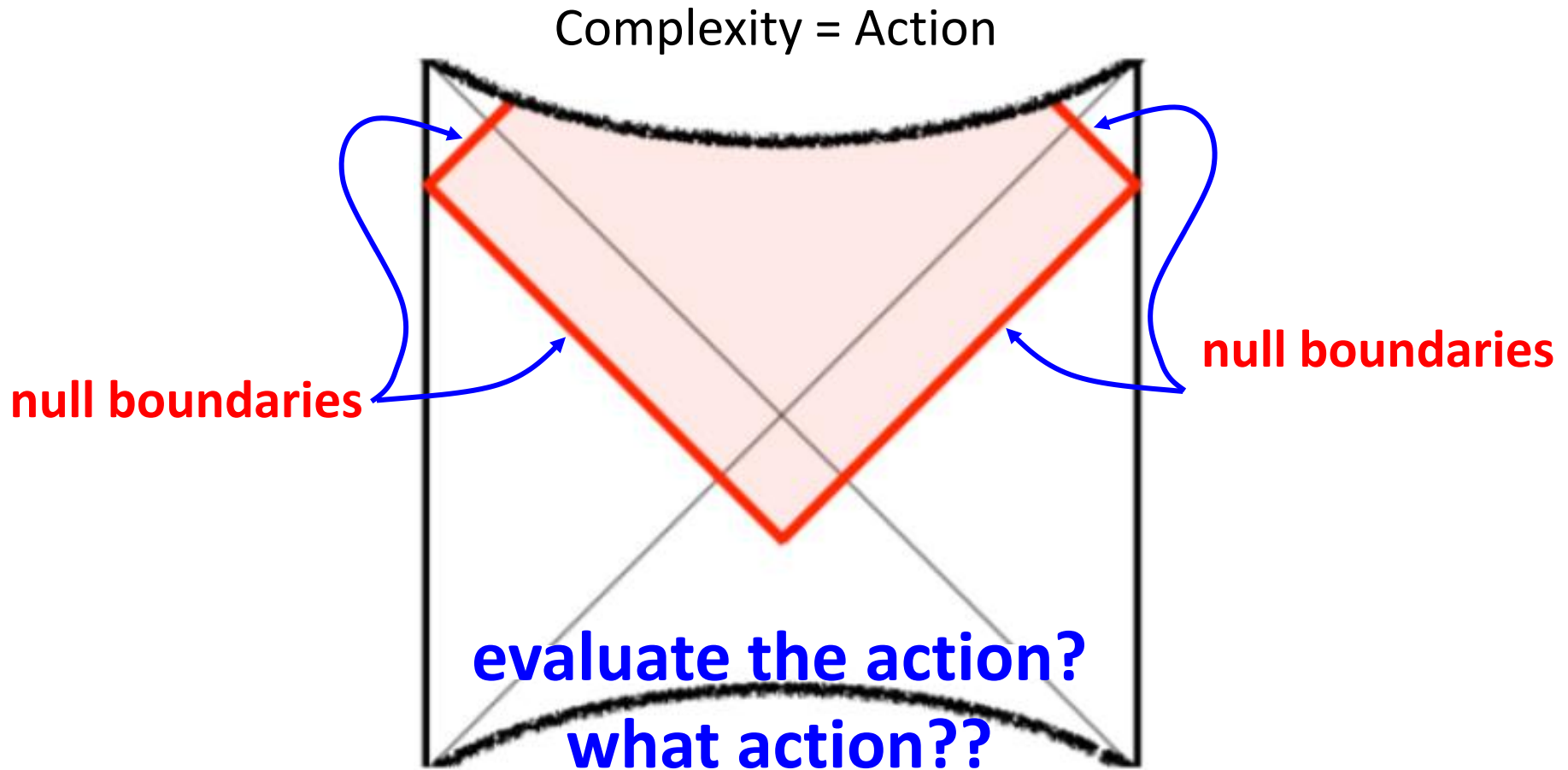
# Complexity = Action:



Wheeler-DeWitt patch:

domain of dependence of Cauchy surface ending on  
boundary time slice

# Complexity = Action:



Wheeler-DeWitt patch:

domain of dependence of Cauchy surface ending on  
boundary time slice

# Complexity = Action:

- gravitational action:

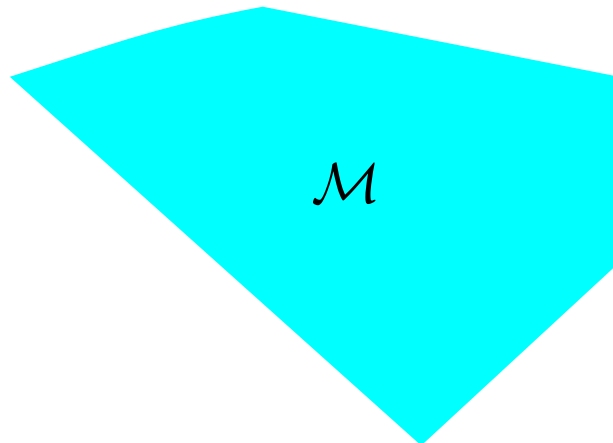
$$I = \frac{1}{16\pi G_N} \int_{\mathcal{M}} d^{d+1}x \sqrt{-g} \left( R + \frac{d(d-1)}{L^2} \right)$$

# Complexity = Action:

- gravitational action:

$$I = \frac{1}{16\pi G_N} \int_{\mathcal{M}} d^{d+1}x \sqrt{-g} \left( R + \frac{d(d-1)}{L^2} \right)$$

- well-defined action principle requires boundary terms



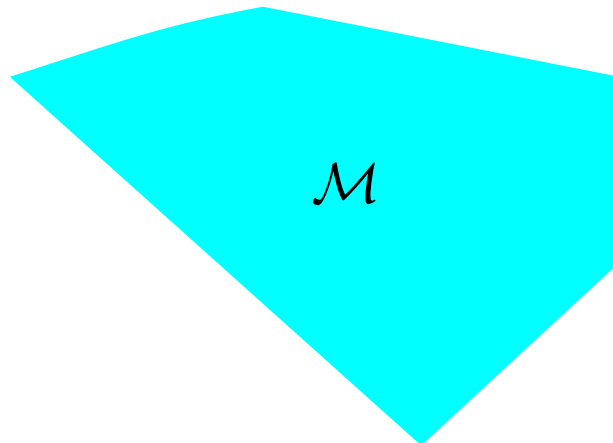
# Complexity = Action:

- gravitational action:

$$I = \frac{1}{16\pi G_N} \int_{\mathcal{M}} d^{d+1}x \sqrt{-g} \left( R + \frac{d(d-1)}{L^2} \right)$$

$$\text{eg, } - \int_{\mathcal{M}} (\nabla\Phi)^2 = \int_{\mathcal{M}} \Phi \nabla^2 \Phi - \int_{\partial\mathcal{M}} \Phi \mathbf{n} \cdot \nabla \Phi$$

- well-defined action principle requires boundary terms



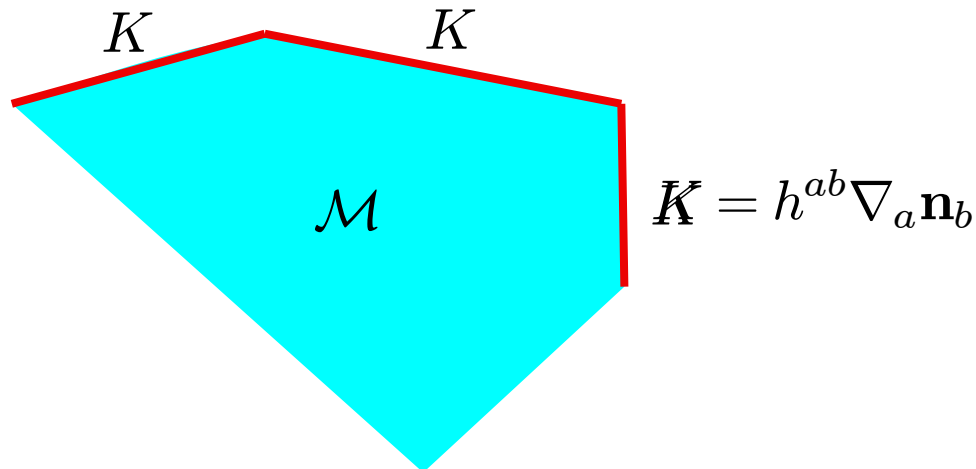
# Complexity = Action:

- gravitational action:

$$I = \frac{1}{16\pi G_N} \int_{\mathcal{M}} d^{d+1}x \sqrt{-g} \left( R + \frac{d(d-1)}{L^2} \right) + \frac{1}{8\pi G_N} \int_{\mathcal{B}} d^d x \sqrt{|h|} K \quad \text{Gibbons-Hawking-York}$$

$$\text{eg, } - \int_{\mathcal{M}} (\nabla\Phi)^2 = \int_{\mathcal{M}} \Phi \nabla^2 \Phi - \int_{\partial\mathcal{M}} \Phi \mathbf{n} \cdot \nabla \Phi$$

- well-defined action principle requires boundary terms



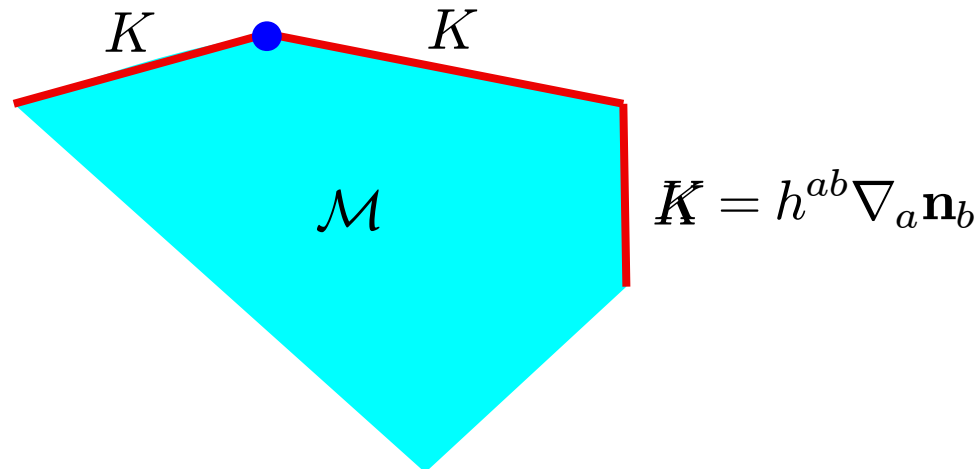
# Complexity = Action:

- gravitational action:

$$I = \frac{1}{16\pi G_N} \int_{\mathcal{M}} d^{d+1}x \sqrt{-g} \left( R + \frac{d(d-1)}{L^2} \right) + \frac{1}{8\pi G_N} \int_{\mathcal{B}} d^d x \sqrt{|h|} K \quad \text{Gibbons-Hawking-York}$$

$$\text{eg, } - \int_{\mathcal{M}} (\nabla\Phi)^2 = \int_{\mathcal{M}} \Phi \nabla^2 \Phi - \int_{\partial\mathcal{M}} \Phi \mathbf{n} \cdot \nabla \Phi$$

- well-defined action principle requires boundary terms





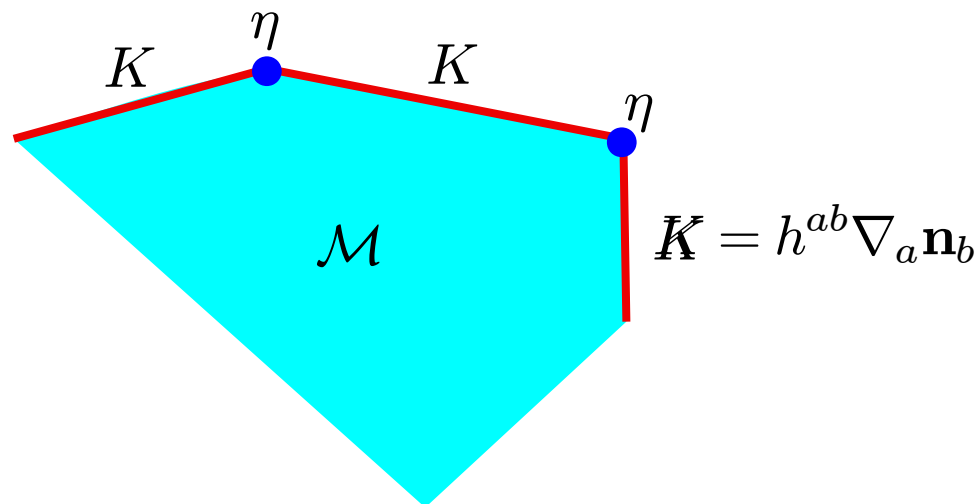
# Complexity = Action:

- gravitational action:

$$I = \frac{1}{16\pi G_N} \int_{\mathcal{M}} d^{d+1}x \sqrt{-g} \left( R + \frac{d(d-1)}{L^2} \right) + \frac{1}{8\pi G_N} \int_{\mathcal{B}} d^d x \sqrt{|h|} K + \frac{1}{8\pi G_N} \int_{\Sigma} d^{d-1}x \sqrt{\sigma} \eta$$

Hayward

- well-defined action principle requires boundary terms

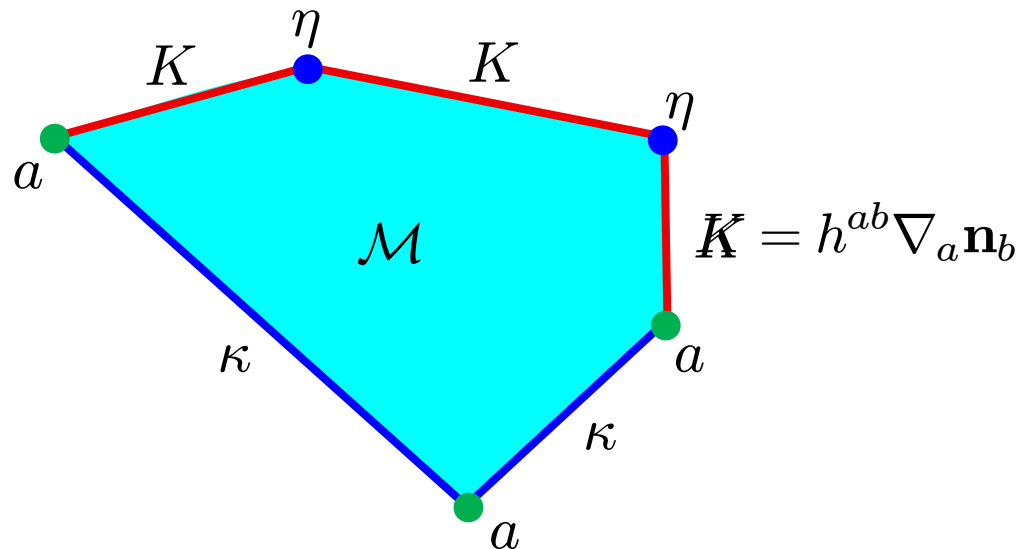


# Complexity = Action:

- gravitational action:

$$\begin{aligned}
 I = & \frac{1}{16\pi G_N} \int_{\mathcal{M}} d^{d+1}x \sqrt{-g} \left( R + \frac{d(d-1)}{L^2} \right) \\
 & + \frac{1}{8\pi G_N} \int_{\mathcal{B}} d^d x \sqrt{|h|} K \quad + \quad \frac{1}{8\pi G_N} \int_{\Sigma} d^{d-1}x \sqrt{\sigma} \eta \\
 \text{eg, } & \frac{1}{8\pi G_N} \int_{\mathcal{B}'} d\lambda d^{d-1}\theta \sqrt{\gamma} \kappa \quad + \quad \frac{1}{8\pi G_N} \int_{\Sigma'} d^{d-1}x \sqrt{\sigma} a
 \end{aligned}$$

- well-defined action principle requires boundary terms



# Complexity = Action:

- gravitational action:

$$\begin{aligned}
 I = & \frac{1}{16\pi G_N} \int_{\mathcal{M}} d^{d+1}x \sqrt{-g} \left( R + \frac{d(d-1)}{L^2} \right) \\
 & + \frac{1}{8\pi G_N} \int_{\mathcal{B}} d^d x \sqrt{|h|} K \quad + \frac{1}{8\pi G_N} \int_{\Sigma} d^{d-1}x \sqrt{\sigma} \eta \\
 & - \frac{1}{8\pi G_N} \int_{\mathcal{B}'} d\lambda d^{d-1}\theta \sqrt{\gamma} \kappa \quad + \frac{1}{8\pi G_N} \int_{\Sigma'} d^{d-1}x \sqrt{\sigma} a
 \end{aligned}$$

**+ ???????**

- ambiguities: total derivatives, extra boundary terms, . . . .

# Complexity = Action:

- gravitational action:

$$\begin{aligned}
 I = & \frac{1}{16\pi G_N} \int_{\mathcal{M}} d^{d+1}x \sqrt{-g} \left( R + \frac{d(d-1)}{L^2} \right) \\
 & + \frac{1}{8\pi G_N} \int_{\mathcal{B}} d^d x \sqrt{|h|} K \quad + \frac{1}{8\pi G_N} \int_{\Sigma} d^{d-1}x \sqrt{\sigma} \eta \\
 & - \frac{1}{8\pi G_N} \int_{\mathcal{B}'} d\lambda d^{d-1}\theta \sqrt{\gamma} \kappa \quad + \frac{1}{8\pi G_N} \int_{\Sigma'} d^{d-1}x \sqrt{\sigma} a
 \end{aligned}$$

**+ ???????**

- ambiguities: total derivatives, extra boundary terms, . . . .

**→ ambiguities in circuit complexity?**

# Complexity = Action:

- gravitational action:

$$\begin{aligned}
 I = & \frac{1}{16\pi G_N} \int_{\mathcal{M}} d^{d+1}x \sqrt{-g} \left( R + \frac{d(d-1)}{L^2} \right) \\
 & + \frac{1}{8\pi G_N} \int_{\mathcal{B}} d^d x \sqrt{|h|} K \quad + \quad \frac{1}{8\pi G_N} \int_{\Sigma} d^{d-1}x \sqrt{\sigma} \eta \\
 & - \frac{1}{8\pi G_N} \int_{\mathcal{B}'} d\lambda d^{d-1}\theta \sqrt{\gamma} \kappa \quad + \quad \frac{1}{8\pi G_N} \int_{\Sigma'} d^{d-1}x \sqrt{\sigma} a
 \end{aligned}$$

# Complexity = Action:

- gravitational action:

$$\begin{aligned}
 I = & \frac{1}{16\pi G_N} \int_{\mathcal{M}} d^{d+1}x \sqrt{-g} \left( R + \frac{d(d-1)}{L^2} \right) \\
 & + \frac{1}{8\pi G_N} \int_{\mathcal{B}} d^d x \sqrt{|h|} K \quad + \quad \frac{1}{8\pi G_N} \int_{\Sigma} d^{d-1}x \sqrt{\sigma} \eta \\
 & - \frac{1}{8\pi G_N} \int_{\mathcal{B}'} d\lambda d^{d-1}\theta \sqrt{\gamma} \kappa \quad + \quad \frac{1}{8\pi G_N} \int_{\Sigma'} d^{d-1}x \sqrt{\sigma} a
 \end{aligned}$$

- ambiguities:

$$1) \quad k^\nu \nabla_\nu k^\mu = \boldsymbol{\kappa} k^\mu$$

# Complexity = Action:

- gravitational action:

$$\begin{aligned}
 I = & \frac{1}{16\pi G_N} \int_{\mathcal{M}} d^{d+1}x \sqrt{-g} \left( R + \frac{d(d-1)}{L^2} \right) \\
 & + \frac{1}{8\pi G_N} \int_{\mathcal{B}} d^d x \sqrt{|h|} K \quad + \quad \frac{1}{8\pi G_N} \int_{\Sigma} d^{d-1}x \sqrt{\sigma} \eta \\
 & - \frac{1}{8\pi G_N} \int_{\mathcal{B}'} d\lambda d^{d-1}\theta \sqrt{\gamma} \kappa \quad + \quad \frac{1}{8\pi G_N} \int_{\Sigma'} d^{d-1}x \sqrt{\sigma} a
 \end{aligned}$$

- ambiguities:

$$1) \quad k^\nu \nabla_\nu k^\mu = \boldsymbol{\kappa} k^\mu \quad \longrightarrow \quad \text{affinely parameterize: } \boldsymbol{\kappa} = 0$$

# Complexity = Action:

- gravitational action:

$$\begin{aligned}
 I = & \frac{1}{16\pi G_N} \int_{\mathcal{M}} d^{d+1}x \sqrt{-g} \left( R + \frac{d(d-1)}{L^2} \right) \\
 & + \frac{1}{8\pi G_N} \int_{\mathcal{B}} d^d x \sqrt{|h|} K \quad + \quad \frac{1}{8\pi G_N} \int_{\Sigma} d^{d-1}x \sqrt{\sigma} \eta \\
 & - \frac{1}{8\pi G_N} \int_{\mathcal{B}'} d\lambda d^{d-1}\theta \sqrt{\gamma} \kappa \quad + \quad \frac{1}{8\pi G_N} \int_{\Sigma'} d^{d-1}x \sqrt{\sigma} a
 \end{aligned}$$

- ambiguities:

$$1) \quad k^\nu \nabla_\nu k^\mu = \boldsymbol{\kappa} k^\mu \quad \longrightarrow \quad \text{affinely parameterize: } \boldsymbol{\kappa} = 0$$

**Aside:** make choices without referring to particular metric/coordinates, which allow for meaningful comparison of different states/geometries



# Complexity = Action:

- gravitational action:

$$\begin{aligned}
 I = & \frac{1}{16\pi G_N} \int_{\mathcal{M}} d^{d+1}x \sqrt{-g} \left( R + \frac{d(d-1)}{L^2} \right) \\
 & + \frac{1}{8\pi G_N} \int_{\mathcal{B}} d^d x \sqrt{|h|} K \quad + \quad \frac{1}{8\pi G_N} \int_{\Sigma} d^{d-1}x \sqrt{\sigma} \eta \\
 & - \frac{1}{8\pi G_N} \int_{\mathcal{B}'} d\lambda d^{d-1}\theta \sqrt{\gamma} \kappa \quad + \quad \frac{1}{8\pi G_N} \int_{\Sigma'} d^{d-1}x \sqrt{\sigma} a
 \end{aligned}$$

- ambiguities:

$$1) \quad k^\nu \nabla_\nu k^\mu = \boldsymbol{\kappa} k^\mu \quad \longrightarrow \quad \text{affinely parameterize: } \boldsymbol{\kappa} = 0$$

$$2) \quad a = \mathbf{a}_0 + \varepsilon \begin{cases} \log |\mathbf{k}_1 \cdot \mathbf{n}_2| \\ \log |\mathbf{k}_1 \cdot \mathbf{k}_2/2| \end{cases}$$

# Complexity = Action:

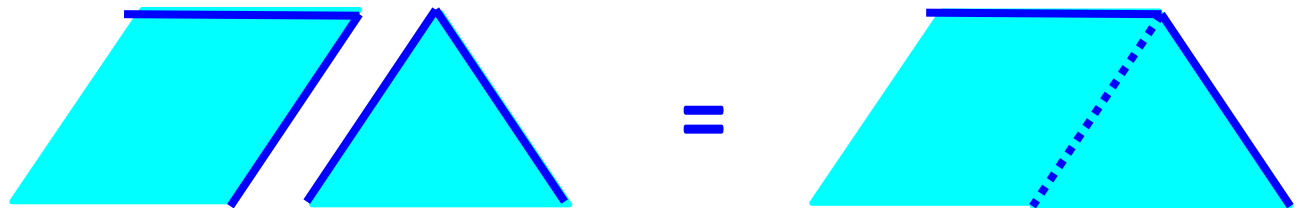
- gravitational action:

$$\begin{aligned}
 I = & \frac{1}{16\pi G_N} \int_{\mathcal{M}} d^{d+1}x \sqrt{-g} \left( R + \frac{d(d-1)}{L^2} \right) \\
 & + \frac{1}{8\pi G_N} \int_{\mathcal{B}} d^d x \sqrt{|h|} K \quad + \quad \frac{1}{8\pi G_N} \int_{\Sigma} d^{d-1}x \sqrt{\sigma} \eta \\
 & - \frac{1}{8\pi G_N} \int_{\mathcal{B}'} d\lambda d^{d-1}\theta \sqrt{\gamma} \kappa \quad + \quad \frac{1}{8\pi G_N} \int_{\Sigma'} d^{d-1}x \sqrt{\sigma} a
 \end{aligned}$$

- ambiguities:

$$1) \quad k^\nu \nabla_\nu k^\mu = \boldsymbol{\kappa} k^\mu \quad \longrightarrow \quad \text{affinely parameterize: } \boldsymbol{\kappa} = 0$$

$$2) \quad a = \mathbf{a}_0 + \varepsilon \begin{cases} \log |\mathbf{k}_1 \cdot \mathbf{n}_2| \\ \log |\mathbf{k}_1 \cdot \mathbf{k}_2/2| \end{cases} \longrightarrow \quad \text{additivity of grav. action: } \mathbf{a}_0 = 0$$



# Complexity = Action:

- gravitational action:

$$\begin{aligned}
 I = & \frac{1}{16\pi G_N} \int_{\mathcal{M}} d^{d+1}x \sqrt{-g} \left( R + \frac{d(d-1)}{L^2} \right) \\
 & + \frac{1}{8\pi G_N} \int_{\mathcal{B}} d^d x \sqrt{|h|} K \quad + \quad \frac{1}{8\pi G_N} \int_{\Sigma} d^{d-1}x \sqrt{\sigma} \eta \\
 & - \frac{1}{8\pi G_N} \int_{\mathcal{B}'} d\lambda d^{d-1}\theta \sqrt{\gamma} \kappa \quad + \quad \frac{1}{8\pi G_N} \int_{\Sigma'} d^{d-1}x \sqrt{\sigma} a
 \end{aligned}$$

- ambiguities:

1)  $k^\nu \nabla_\nu k^\mu = \boldsymbol{\kappa} k^\mu \longrightarrow$  affinely parameterize:  $\boldsymbol{\kappa} = 0$

2)  $a = \mathbf{a}_0 + \varepsilon \begin{cases} \log |\mathbf{k}_1 \cdot \mathbf{n}_2| \\ \log |\mathbf{k}_1 \cdot \mathbf{k}_2/2| \end{cases} \longrightarrow$  additivity of grav. action:  
 $\mathbf{a}_0 = 0$

3) constant rescaling:  $k^\mu \rightarrow \boldsymbol{\alpha} k^\mu$

# Complexity = Action:

- gravitational action:

$$\begin{aligned}
 I = & \frac{1}{16\pi G_N} \int_{\mathcal{M}} d^{d+1}x \sqrt{-g} \left( R + \frac{d(d-1)}{L^2} \right) \\
 & + \frac{1}{8\pi G_N} \int_{\mathcal{B}} d^d x \sqrt{|h|} K \quad + \quad \frac{1}{8\pi G_N} \int_{\Sigma} d^{d-1}x \sqrt{\sigma} \eta \\
 & - \frac{1}{8\pi G_N} \int_{\mathcal{B}'} d\lambda d^{d-1}\theta \sqrt{\gamma} \kappa \quad + \quad \frac{1}{8\pi G_N} \int_{\Sigma'} d^{d-1}x \sqrt{\sigma} a
 \end{aligned}$$

- ambiguities:

1)  $k^\nu \nabla_\nu k^\mu = \boldsymbol{\kappa} k^\mu \longrightarrow$  affinely parameterize:  $\boldsymbol{\kappa} = 0$

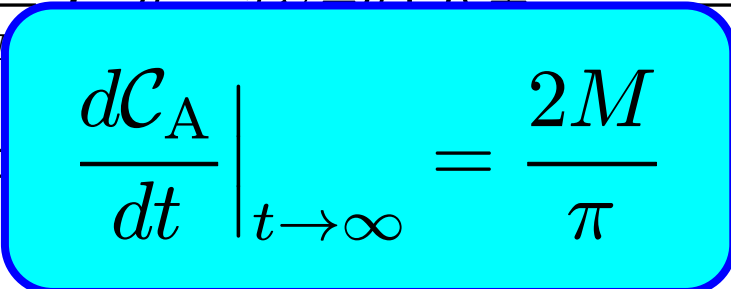
2)  $a = \mathbf{a}_0 + \varepsilon \begin{cases} \log |\mathbf{k}_1 \cdot \mathbf{n}_2| \\ \log |\mathbf{k}_1 \cdot \mathbf{k}_2 / 2| \end{cases} \longrightarrow$  additivity of grav. action:  
 $\mathbf{a}_0 = 0$

3) constant rescaling:  $k^\mu \rightarrow \boldsymbol{\alpha} k^\mu \longrightarrow$  fix normalization asymptotically:  
 $\mathbf{k} \cdot \hat{\mathbf{t}} = \pm 1$

# Complexity = Action:

- gravitational action:

$$I = \frac{1}{16\pi G_N} \int_{\mathcal{M}} d^{d+1}x \sqrt{-g} \left( R - \frac{d(d-1)}{2\ell^2} \right) + \frac{1}{8\pi G_N} \frac{d\mathcal{C}_A}{dt} \Big|_{t \rightarrow \infty} = \frac{2M}{\pi} + \frac{1}{8\pi G_N} \int_{\Sigma'} d^{d-1}x \sqrt{\sigma} a$$



- ambiguities:

1)  $k^\nu \nabla_\nu k^\mu = \boldsymbol{\kappa} k^\mu \longrightarrow$  affinely parameterize:  $\boldsymbol{\kappa} = 0$

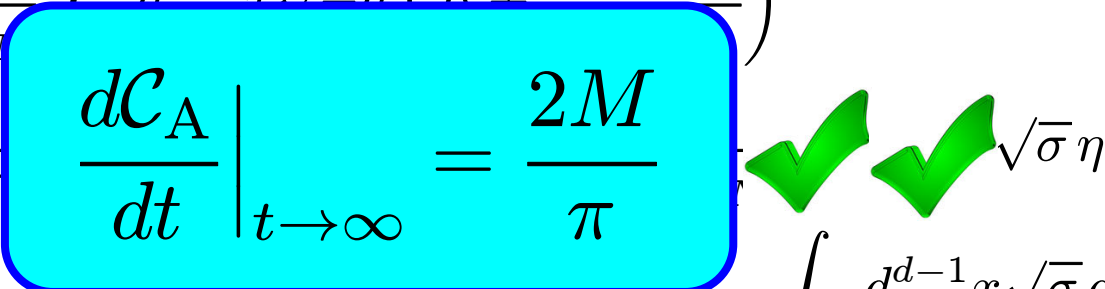
2)  $a = \mathbf{a}_0 + \varepsilon \begin{cases} \log |\mathbf{k}_1 \cdot \mathbf{n}_2| \\ \log |\mathbf{k}_1 \cdot \mathbf{k}_2/2| \end{cases} \longrightarrow$  additivity of grav. action:  
 $\mathbf{a}_0 = 0$

3) constant rescaling:  $k^\mu \rightarrow \boldsymbol{\alpha} k^\mu \longrightarrow$  fix normalization asymptotically:  
 $\mathbf{k} \cdot \hat{\mathbf{t}} = \pm 1$

# Complexity = Action:

- gravitational action:

$$I = \frac{1}{16\pi G_N} \int_{\mathcal{B}} d^{d+1}x \sqrt{-g} \left( R - \frac{d(d-1)}{2} \Lambda \right) + \frac{1}{8\pi G_N} \frac{d\mathcal{C}_A}{dt} \Big|_{t \rightarrow \infty} = \frac{2M}{\pi} - \frac{1}{8\pi G_N} \int_{\Sigma'} d^{d-1}x \sqrt{\sigma} a$$



- ambiguities:

1)  $k^\nu \nabla_\nu k^\mu = \boldsymbol{\kappa} k^\mu \longrightarrow$  affinely parameterize:  $\boldsymbol{\kappa} = 0$


2)  $a = \mathbf{a}_0 + \varepsilon \begin{cases} \log |\mathbf{k}_1 \cdot \mathbf{n}_2| \\ \log |\mathbf{k}_1 \cdot \mathbf{k}_2/2| \end{cases} \longrightarrow$  additivity of grav. action:  
 $\mathbf{a}_0 = 0$

3) constant rescaling:  $k^\mu \rightarrow \boldsymbol{\alpha} k^\mu \longrightarrow$  fix normalization asymptotically:  
 ~~$\mathbf{k} \cdot \hat{\mathbf{t}} = \pm 1$~~

# Complexity = Action:

- gravitational action:

$$I = \frac{1}{16\pi G_N} \int_{\mathcal{B}} d^{d+1}x \sqrt{-g} \left( R - \frac{d(d-1)}{2} \Lambda \right) + \frac{1}{8\pi G_N} \frac{d\mathcal{C}_A}{dt} \Big|_{t \rightarrow \infty} = \frac{2M}{\pi} - \frac{1}{8\pi G_N} \int_{\Sigma'} d^{d-1}x \sqrt{\sigma} a$$



- ambiguities:

1)  $k^\nu \nabla_\nu k^\mu = \boldsymbol{\kappa} k^\mu \longrightarrow$  affinely parameterize:  $\boldsymbol{\kappa} = 0$

2)  $a = \mathbf{a}_0 + \varepsilon \begin{cases} \log |\mathbf{k}_1 \cdot \mathbf{n}_2| \\ \log |\mathbf{k}_1 \cdot \mathbf{k}_2/2| \end{cases} \longrightarrow$  ~~additivity of grav. action.~~  
 ~~$\mathbf{a}_0 = 0$~~

3) constant rescaling:  $k^\mu \rightarrow \boldsymbol{\alpha} k^\mu \longrightarrow$  ~~fix normalization asymptotically:~~  
 ~~$\mathbf{k} \cdot \hat{\mathbf{t}} = \pm 1$~~

# Complexity = Action:

- gravitational action:

$$I = \frac{1}{16\pi G_N} \int_{\mathcal{B}} d^{d+1}x \sqrt{-g} \left( R - \frac{d(d-1)}{2\ell^2} \right) + \frac{1}{8\pi G_N} \frac{d\mathcal{C}_A}{dt} \Big|_{t \rightarrow \infty} = \frac{2M}{\pi} - \frac{1}{8\pi G_N} \int_{\Sigma'} d^{d-1}x \sqrt{\sigma} a$$

- ambiguities:

1)  $k^\nu \nabla_\nu k^\mu = \kappa k^\mu \longrightarrow$  ~~affinely parameterize.  $\kappa = 0$~~

2)  $a = \mathbf{a}_0 + \varepsilon \begin{cases} \log |\mathbf{k}_1 \cdot \mathbf{n}_2| \\ \log |\mathbf{k}_1 \cdot \mathbf{k}_2/2| \end{cases} \longrightarrow$  ~~additivity of grav. action.  $\mathbf{a}_0 = 0$~~

3) constant rescaling:  $k^\mu \rightarrow \alpha k^\mu \longrightarrow$  ~~fix normalization asymptotically:  $\mathbf{k} \cdot \hat{\mathbf{t}} = \pm 1$~~



# Complexity = Action:

- gravitational action:

$$I = \frac{1}{16\pi G_N} \int_{\mathcal{M}} d^{d+1}x \sqrt{-g} \left( R - \frac{d(d-1)}{2} \Lambda \right) + \frac{1}{8\pi G_N} \frac{d\mathcal{C}_A}{dt} \Big|_{t \rightarrow \infty} = \frac{2M}{\pi} - \frac{1}{8\pi G_N} \int_{\Sigma'} d^{d-1}x \sqrt{\sigma} a$$

- ambiguities:

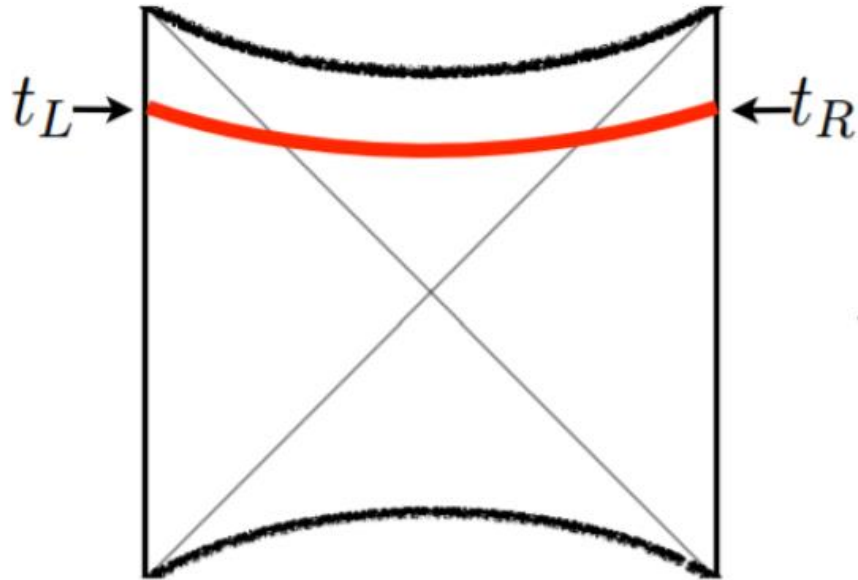
1)  $k^\nu \nabla_\nu k^\mu = \kappa k^\mu \longrightarrow$  ~~affinely parameterize.  $\kappa = 0$~~

**Aside:** make choices without referring to particular metric/coordinates, which allow for meaningful comparison of different states/geometries

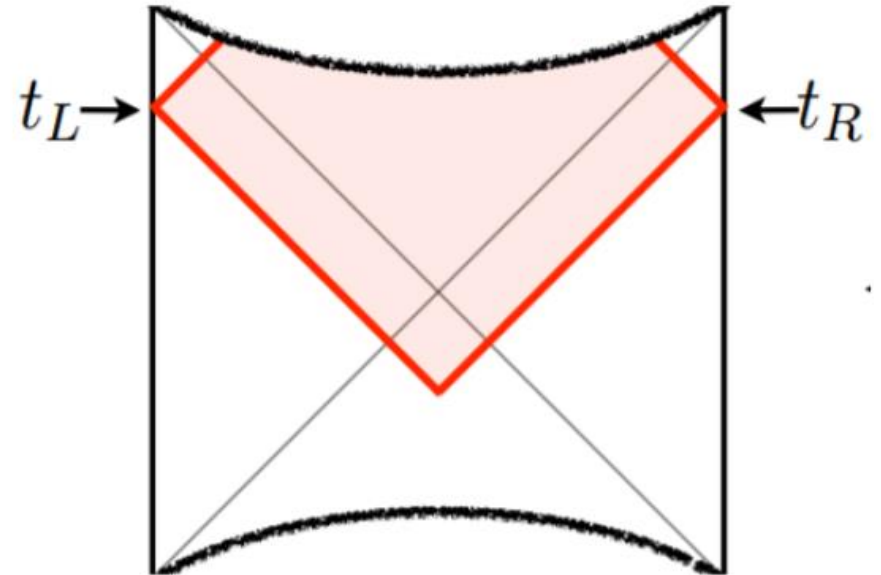
3) constant rescaling:  $k^\mu \rightarrow \alpha k^\mu \longrightarrow$  ~~fix normalization asymptotically:  $k \cdot \hat{t} = \pm 1$~~

# Holographic Complexity:

Complexity = Volume



Complexity = Action

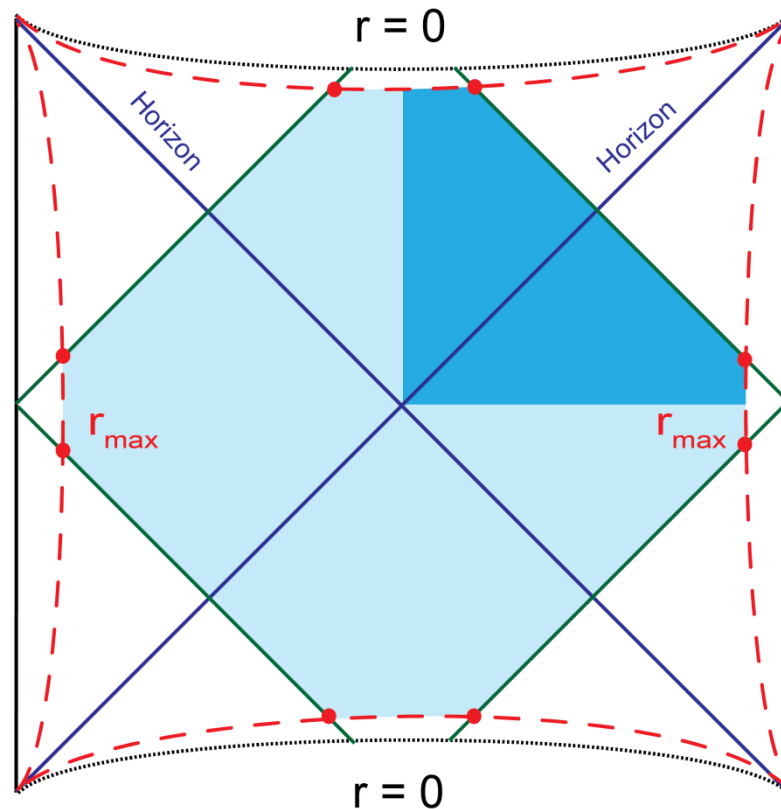


$$\left. \frac{d\mathcal{C}_V}{dt} \right|_{t \rightarrow \infty} = \frac{8\pi}{d-1} M \quad (\text{planar})$$

$$\left. \frac{d\mathcal{C}_A}{dt} \right|_{t \rightarrow \infty} = \frac{2M}{\pi}$$

# Complexity of Formation:

$$|\text{TFD}\rangle = Z^{-1/2} \sum_{\alpha} e^{-E_{\alpha}/(2T)} |E_{\alpha}\rangle_L |E_{\alpha}\rangle_R$$

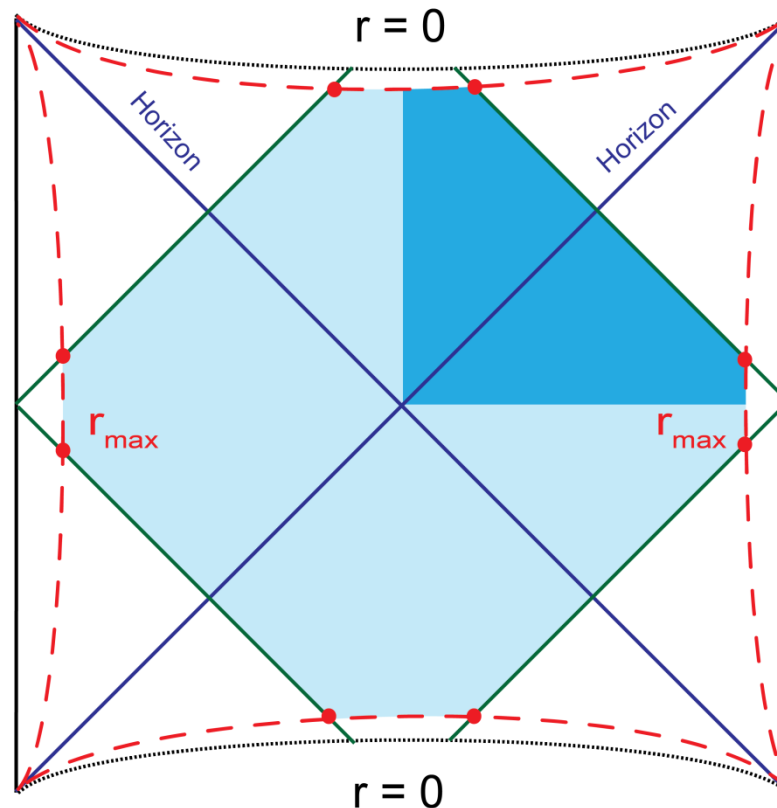


# Complexity of Formation:

$$|\text{TFD}\rangle = Z^{-1/2} \sum_{\alpha} e^{-E_{\alpha}/(2T)} |E_{\alpha}\rangle_L |E_{\alpha}\rangle_R$$

- additional complexity involved in forming thermofield double state compared to preparing two copies of vacuum state?

$$\Delta\mathcal{C} = \mathcal{C}(|\text{TFD}\rangle) - \mathcal{C}(|0\rangle \quad |0\rangle)$$





# Complexity of Formation:

- additional complexity involved in forming thermofield double state compared to preparing two copies of vacuum state?

$$\Delta\mathcal{C} = \mathcal{C}(|\text{TFD}\rangle) - \mathcal{C}(|0\rangle \quad |0\rangle)$$

$$\Delta\mathcal{C}_A = \frac{d-2}{d\pi} \cot\left(\frac{\pi}{d}\right) S + \dots$$

↑ thermal/ent. entropy      ↑ curvature corrections

( $d =$  boundary dimension)

# Complexity of Formation:

- additional complexity involved in forming thermofield double state compared to preparing two copies of vacuum state?

$$\Delta\mathcal{C} = \mathcal{C}(|\text{TFD}\rangle) - \mathcal{C}(|0\rangle \quad |0\rangle)$$

$$\Delta\mathcal{C}_A = \underbrace{\frac{d-2}{d\pi} \cot\left(\frac{\pi}{d}\right)}_{\frac{d-2}{\pi^2} + \mathcal{O}(1/d)} S + \dots$$

↑ thermal/ent. entropy      ↑ curvature corrections

( $d$  = boundary dimension)

# Complexity of Formation:

- additional complexity involved in forming thermofield double state compared to preparing two copies of vacuum state?

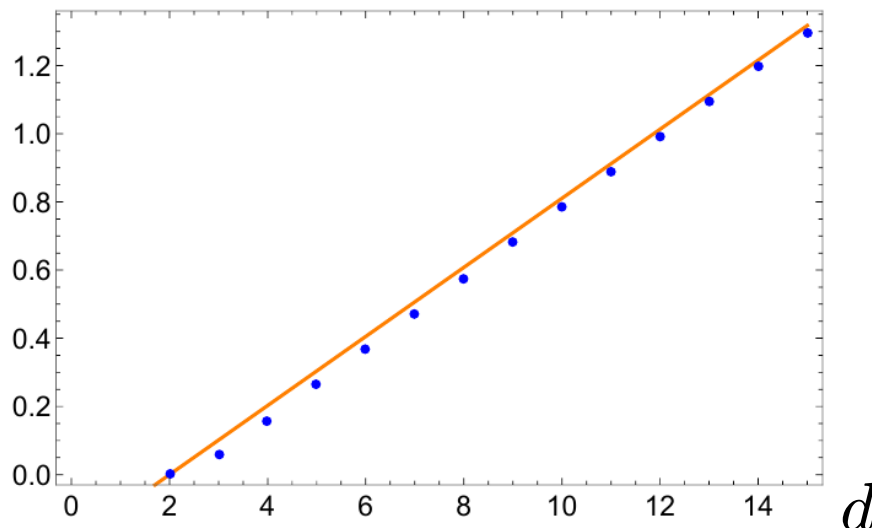
$$\Delta\mathcal{C} = \mathcal{C}(|\text{TFD}\rangle) - \mathcal{C}(|0\rangle |0\rangle)$$

$$\Delta\mathcal{C}_A = \underbrace{\frac{d-2}{d\pi} \cot\left(\frac{\pi}{d}\right)}_{\frac{d-2}{\pi^2} + \mathcal{O}(1/d)} S + \dots$$

↑ thermal/ent. entropy  
 ↑ curvature corrections

•  $\frac{d-2}{d\pi} \cot\left(\frac{\pi}{d}\right)$   
 —  $\frac{d-2}{\pi^2}$

( $d =$  boundary dimension)





# Complexity of Formation:

- additional complexity involved in forming thermofield double state compared to preparing two copies of vacuum state?

$$\Delta\mathcal{C} = \mathcal{C}(|\text{TFD}\rangle) - \mathcal{C}(|0\rangle \quad |0\rangle)$$

$$\Delta\mathcal{C}_A = \underbrace{\frac{d-2}{d\pi} \cot\left(\frac{\pi}{d}\right)}_{\frac{d-2}{\pi^2} + \mathcal{O}(1/d)} S + \dots$$

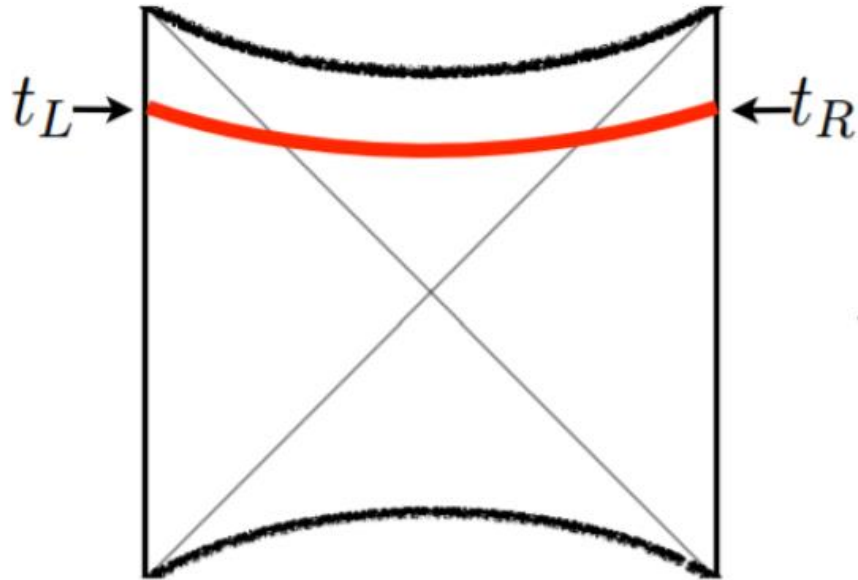
↑ thermal/ent. entropy  
↑ curvature corrections

$$\Delta\mathcal{C}_V = \underbrace{4\sqrt{\pi} \frac{(d-2) \Gamma(1 + \frac{1}{d})}{(d-1) \Gamma(\frac{1}{2} + \frac{1}{d})}}_{4 + \mathcal{O}(1/d)} S + \dots$$

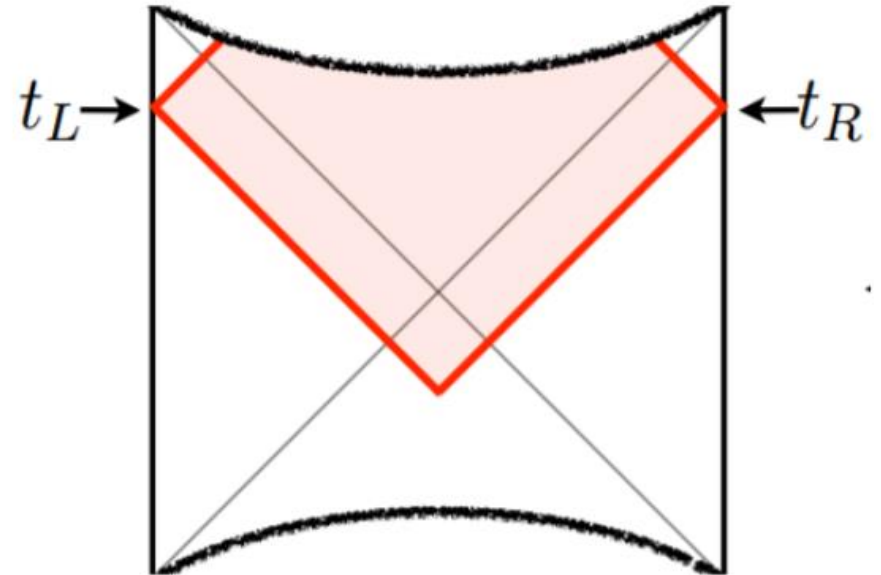
( $d$  = boundary dimension)

# Holographic Complexity:

Complexity = Volume



Complexity = Action



$$\left. \frac{d\mathcal{C}_V}{dt} \right|_{t \rightarrow \infty} = \frac{8\pi}{d-1} M \quad (\text{planar})$$

$$\left. \frac{d\mathcal{C}_A}{dt} \right|_{t \rightarrow \infty} = \frac{2M}{\pi}$$

## Compare?

$$R_{\text{form}} = \frac{\Delta \mathcal{C}_A}{\Delta \mathcal{C}_V} = \frac{d-1}{4\pi^{3/2}} \frac{\Gamma\left(1 - \frac{1}{d}\right)}{\Gamma\left(\frac{1}{2} - \frac{1}{d}\right)} \simeq \frac{d}{4\pi^2}$$

$$R_{\text{rate}} = \frac{d\mathcal{C}_A/dt}{d\mathcal{C}_V/dt} = \frac{d-1}{4\pi^2} \simeq \frac{d}{4\pi^2}$$

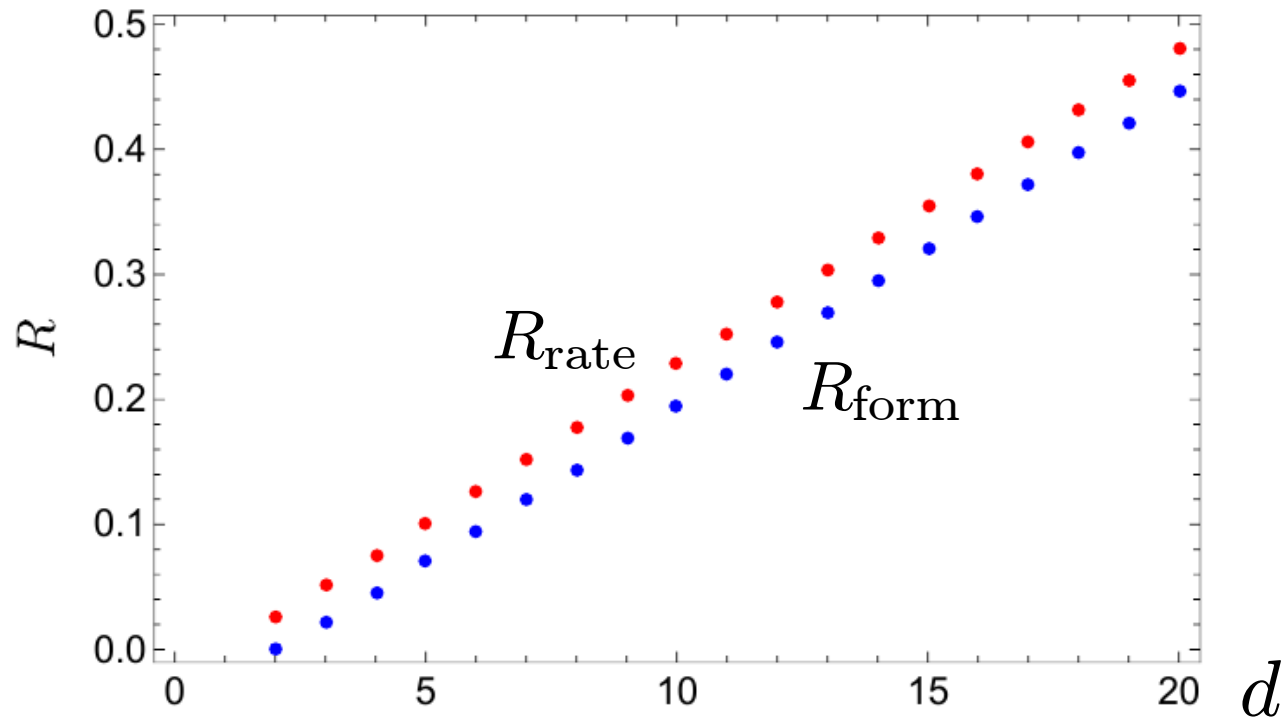
$$R_{\text{rate}} - R_{\text{form}} = \frac{\log 2}{2\pi^2} + \mathcal{O}(1/d)$$

# Compare?

$$R_{\text{form}} = \frac{\Delta \mathcal{C}_A}{\Delta \mathcal{C}_V} = \frac{d-1}{4\pi^{3/2}} \frac{\Gamma\left(1 - \frac{1}{d}\right)}{\Gamma\left(\frac{1}{2} - \frac{1}{d}\right)} \simeq \frac{d}{4\pi^2}$$

$$R_{\text{rate}} = \frac{d\mathcal{C}_A/dt}{d\mathcal{C}_V/dt} = \frac{d-1}{4\pi^2} \simeq \frac{d}{4\pi^2}$$

$$R_{\text{rate}} - R_{\text{form}} = \frac{\log 2}{2\pi^2} + \mathcal{O}(1/d)$$



( $d =$  boundary dimension)

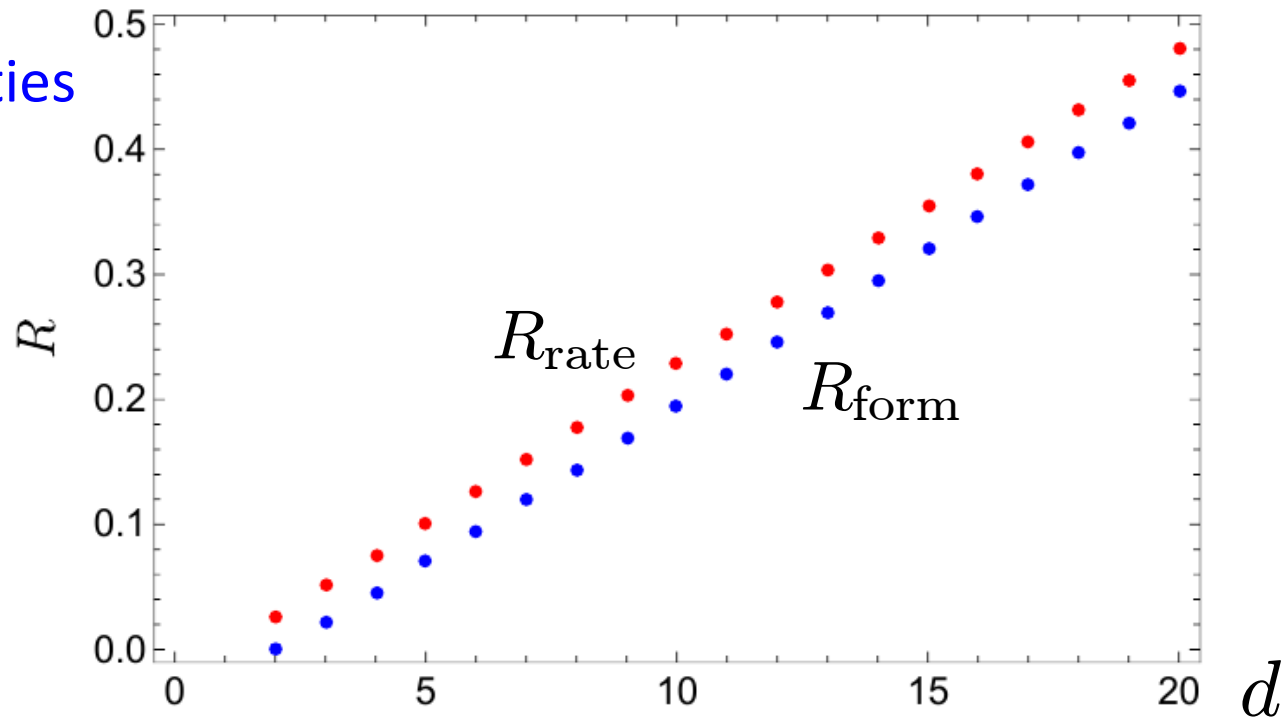
# Compare?

$$R_{\text{form}} = \frac{\Delta \mathcal{C}_A}{\Delta \mathcal{C}_V} = \frac{d-1}{4\pi^{3/2}} \frac{\Gamma(1 - \frac{1}{d})}{\Gamma(\frac{1}{2} - \frac{1}{d})} \simeq \frac{d}{4\pi^2}$$

$$R_{\text{rate}} = \frac{d\mathcal{C}_A/dt}{d\mathcal{C}_V/dt} = \frac{d-1}{4\pi^2} \simeq \frac{d}{4\pi^2}$$

$$R_{\text{rate}} - R_{\text{form}} = \frac{\log 2}{2\pi^2} + \mathcal{O}(1/d)$$

points to consistency  
of C=V and C=A dualities  
up to differences in  
microscopic rules,  
eg, gate set



( $d =$  boundary dimension)

# Complexity of Formation for $d=2$ :

- additional complexity involved in forming thermofield double state compared to preparing two copies of vacuum state?

$$\Delta\mathcal{C} = \mathcal{C}(|\text{TFD}\rangle) - \mathcal{C}(|0\rangle \quad |0\rangle)$$

$$\Delta\mathcal{C}_A = \frac{d-2}{d\pi} \cot\left(\frac{\pi}{d}\right) S + \dots \quad \simeq 0!$$

leading term vanishes

$$\Delta\mathcal{C}_V = 4\sqrt{\pi} \frac{(d-2) \Gamma(1 + \frac{1}{d})}{(d-1) \Gamma(\frac{1}{2} + \frac{1}{d})} S + \dots \quad \simeq 0!$$

- actually holographic calculations apply for  $d \geq 3$  (but still correct)

( $d$  = boundary dimension)

## Complexity of Formation for $d=2$ :

- additional complexity involved in forming thermofield double state compared to preparing two copies of vacuum state?

$$\Delta\mathcal{C} = \mathcal{C}(|\text{TFD}\rangle) - \mathcal{C}(|0\rangle \quad |0\rangle)$$

- reconsider holographic calculations for 3D BTZ black hole:

$$\Delta\mathcal{C}_A = -\frac{c}{3} \qquad \Delta\mathcal{C}_V = +\frac{8\pi}{3} c$$

$c$  = central charge of boundary CFT

- perhaps related to BTZ black hole being locally AdS geometry

( $d$  = boundary dimension)

## Complexity of Formation for $d=2$ :

- additional complexity involved in forming thermofield double state compared to preparing two copies of vacuum state?

$$\Delta\mathcal{C} = \mathcal{C}(|\text{TFD}\rangle) - \mathcal{C}(|0\rangle \quad |0\rangle)$$

- reconsider holographic calculations for 3D BTZ black hole:

$$\Delta\mathcal{C}_A = -\frac{c}{3} \qquad \Delta\mathcal{C}_V = +\frac{8\pi}{3} c \qquad (\text{NS vac})$$

$$= 0 \qquad \qquad \qquad = 0 \qquad \qquad \qquad (\text{R vac})$$

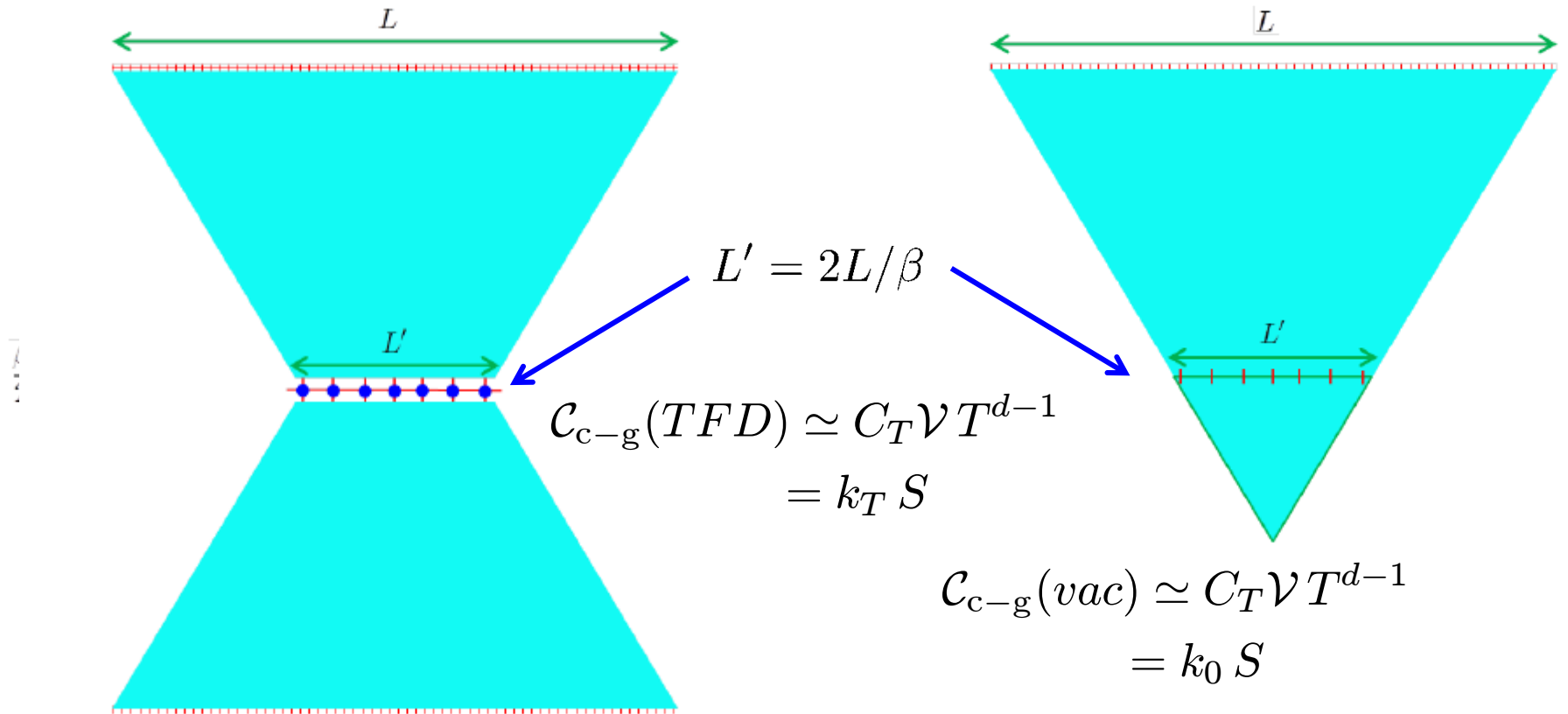
$c$  = central charge of boundary CFT

- perhaps related to BTZ black hole being locally AdS geometry

( $d$  = boundary dimension)



# Complexity of Formation from MERA?



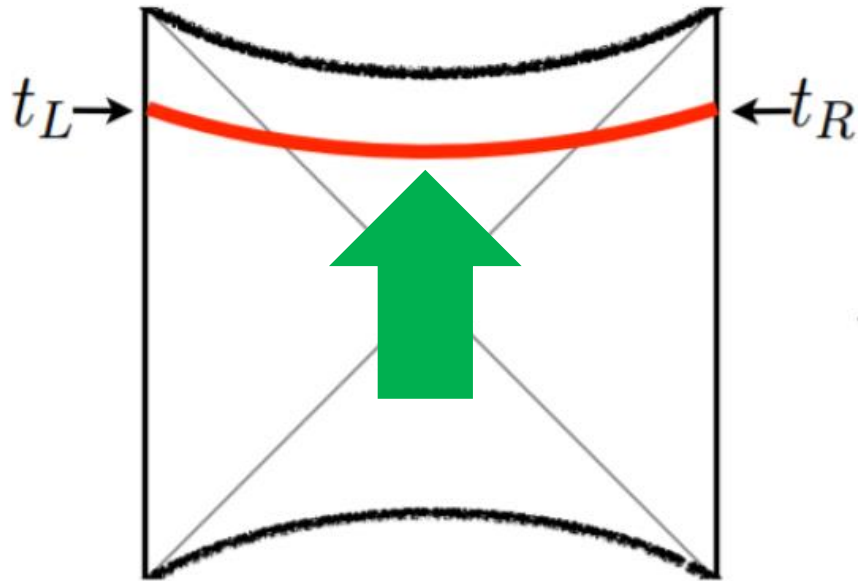
$$\Delta \mathcal{C} = \mathcal{C}(TFD) - 2\mathcal{C}(vac) = (k_T - 2k_0) S \stackrel{?}{\geq} 0 \quad ?$$

(a) Thermofield double state

(b) Vacuum state

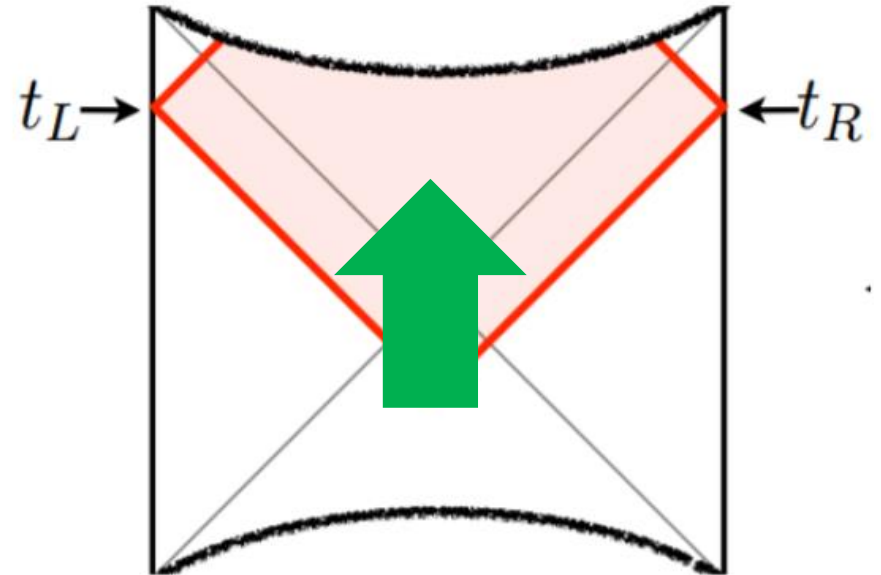
# Holographic Complexity:

Complexity = Volume



$$\mathcal{C}_V(\Sigma) = \max_{\Sigma=\partial\mathcal{B}} \left[ \frac{\mathcal{V}(\mathcal{B})}{G_N \ell} \right]$$

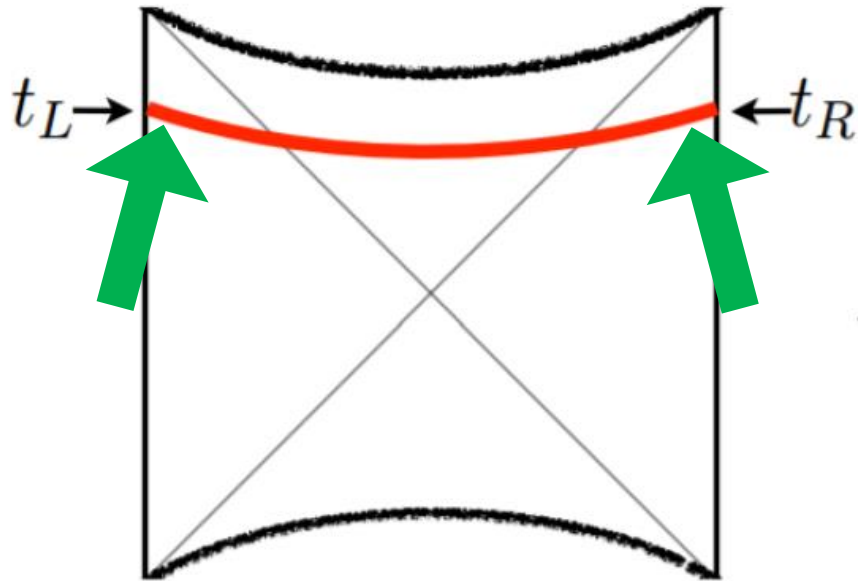
Complexity = Action



$$\mathcal{C}_A(\Sigma) = \frac{I_{\text{WDW}}}{\pi \hbar}$$

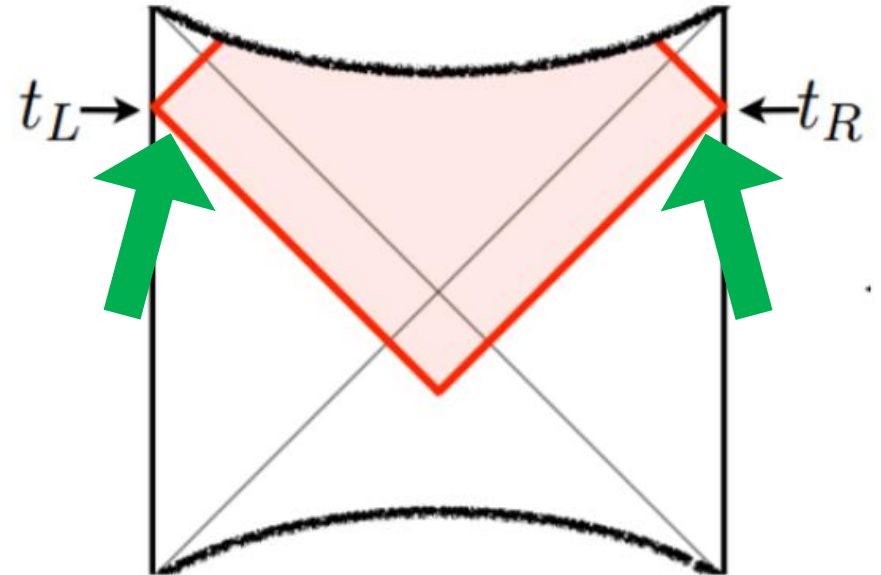
# Holographic Complexity:

Complexity = Volume



$$C_V(\Sigma) = \max_{\Sigma=\partial\mathcal{B}} \left[ \frac{\mathcal{V}(\mathcal{B})}{G_N \ell} \right]$$

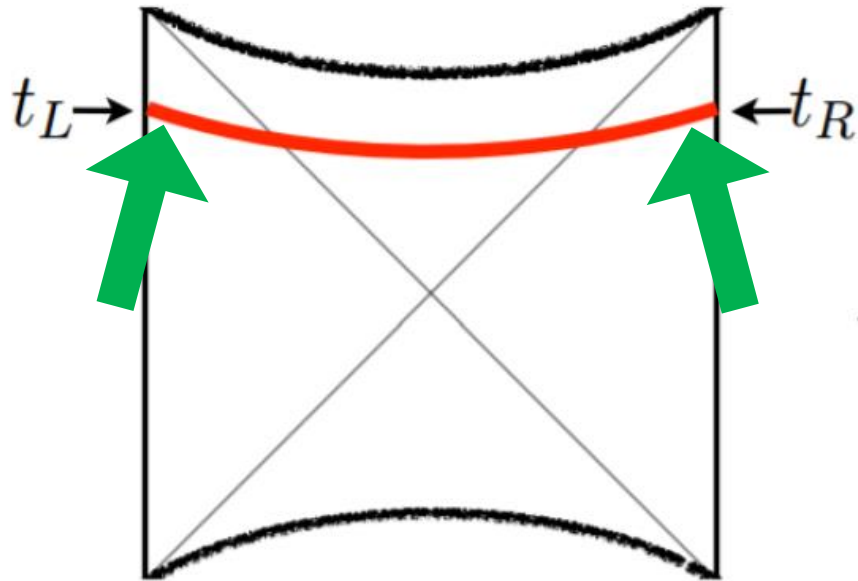
Complexity = Action



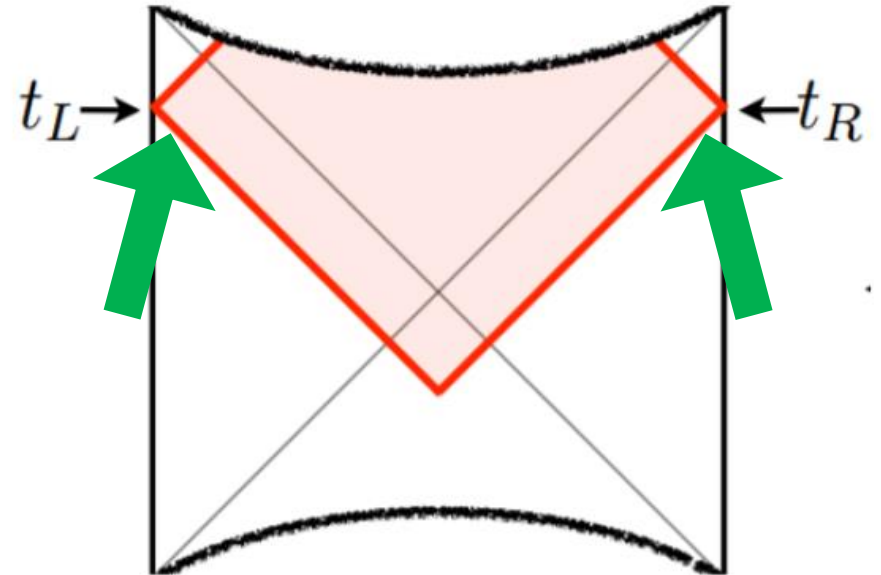
$$C_A(\Sigma) = \frac{I_{\text{WDW}}}{\pi \hbar}$$

# Holographic Complexity:

Complexity = Volume



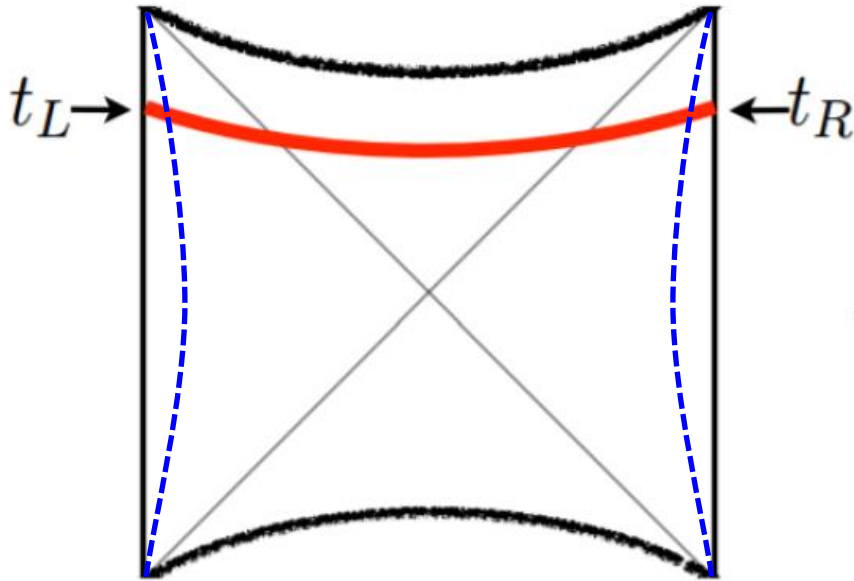
Complexity = Action



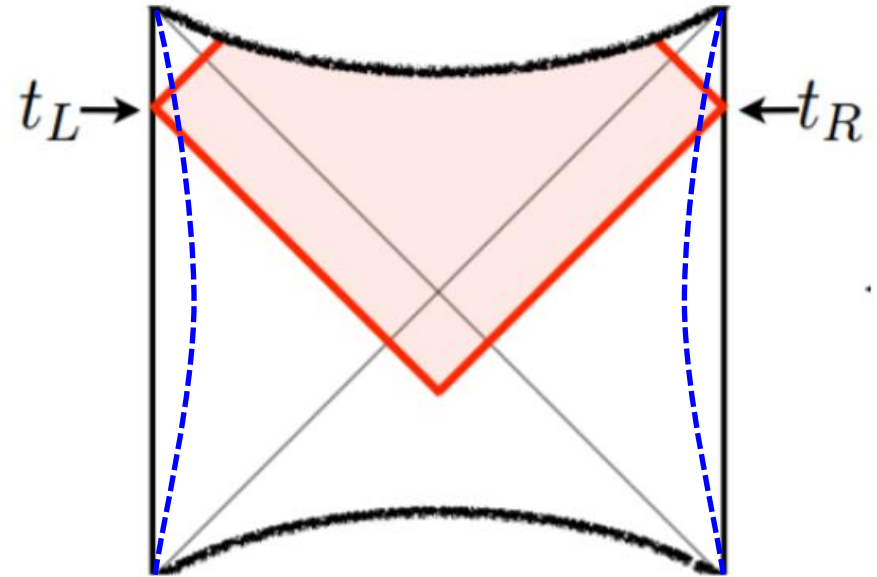
- UV divergences naturally associated with establishing correlations or entanglement down to arbitrarily small length scales

# Holographic Complexity:

Complexity = Volume



Complexity = Action



- UV divergences naturally associated with establishing correlations or entanglement down to arbitrarily small length scales
- regulate volume/action with the introduction of UV regulator surface at large radius ( $r_{max} = L^2/\delta$ ), as usual

# Holographic Complexity:

- UV divergences naturally associated with establishing correlations down to arbitrarily small length scales
- regulate volume/action with the introduction of UV regulator surface at large radius ( $r_{max} = L^2/\delta$ ), as usual

$$\mathcal{C}_V(\Sigma) = \frac{L^{d-1}}{(d-1)G_N} \int_{\Sigma} d^{d-1}\sigma \sqrt{\bar{h}} \left[ \frac{1}{\delta^{d-1}} - \frac{(d-1)}{2(d-2)(d-3)\delta^{d-3}} \left( \mathcal{R}_a^a - \frac{1}{2} \mathcal{R} - \frac{(d-2)^2}{(d-1)^2} K^2 \right) + \dots \right]$$

- UV divergences appear as local integrals of geometric invariants (as with holographic entanglement entropy)

# Holographic Complexity:

- UV divergences naturally associated with establishing correlations down to arbitrarily small length scales
- regulate volume/action with the introduction of UV regulator surface at large radius ( $r_{max} = L^2/\delta$ ), as usual

$$\mathcal{C}_V(\Sigma) = \frac{8\pi^{\frac{d+2}{2}}\Gamma(d/2)}{\Gamma(d+2)} C_T \int_{\Sigma} d^{d-1}\sigma \sqrt{h} \left[ \frac{1}{\delta^{d-1}} - \frac{(d-1)}{2(d-2)(d-3)\delta^{d-3}} \left( \mathcal{R}_a^a - \frac{1}{2}\mathcal{R} - \frac{(d-2)^2}{(d-1)^2}K^2 \right) + \dots \right]$$

- UV divergences appear as local integrals of geometric invariants (as with holographic entanglement entropy)

# Holographic Complexity:

- UV divergences naturally associated with establishing correlations down to arbitrarily small length scales
- regulate volume/action with the introduction of UV regulator surface at large radius ( $r_{max} = L^2/\delta$ ), as usual

$$\mathcal{C}_V(\Sigma) = \frac{1}{\delta^{d-1}} \int_{\Sigma} d^{d-1}\sigma \sqrt{h} v_0(\mathcal{R}, K)$$

$$\mathcal{C}_A(\Sigma) = \frac{1}{\delta^{d-1}} \int_{\Sigma} d^{d-1}\sigma \sqrt{h} \left[ v_1(\mathcal{R}, K) + \log \left( \frac{L}{\alpha \delta} \right) v_2(\mathcal{R}, K) \right]$$

with

$$v_k(\mathcal{R}, K) = \sum_{n=0}^{\lfloor \frac{d-1}{2} \rfloor} \sum_i c_{i,n}^{[k]}(d) \delta^{2n} [\mathcal{R}, K]_i^{2n}$$

- UV divergences appear as local integrals of geometric invariants (as with holographic entanglement entropy)



# Holographic Complexity:

- UV divergences naturally associated with establishing correlations down to arbitrarily small length scales
- regulate volume/action with the introduction of UV regulator surface at large radius ( $r_{max} = L^2/\delta$ ), as usual

$$\mathcal{C}_V(\Sigma) = \frac{1}{\delta^{d-1}} \int_{\Sigma} d^{d-1}\sigma \sqrt{h} v_0(\mathcal{R}, K)$$

$$\mathcal{C}_A(\Sigma) = \frac{1}{\delta^{d-1}} \int_{\Sigma} d^{d-1}\sigma \sqrt{h} \left[ v_1(\mathcal{R}, K) + \log \left( \frac{L}{\alpha \delta} \right) v_2(\mathcal{R}, K) \right]$$

normalization AdS scale

from asymptotic joint

$$\text{with } v_k(\mathcal{R}, K) = \sum_{n=0}^{\lfloor \frac{d-1}{2} \rfloor} \sum_i c_{i,n}^{[k]}(d) \delta^{2n} [\mathcal{R}, K]_i^{2n}$$

- UV divergences appear as local integrals of geometric invariants (as with holographic entanglement entropy)

# Holographic Complexity:

- UV divergences appear as local integrals of geometric invariants

$$\mathcal{C}(\Sigma) \simeq c_0 \mathcal{V}(\Sigma) / \delta^{d-1} + \dots$$

$t = \text{constant}$

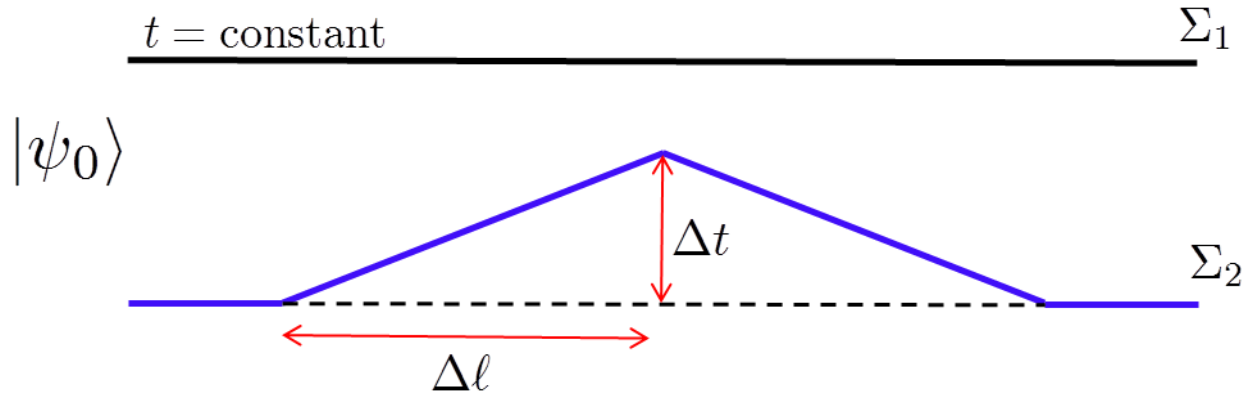
$\Sigma_1$

$|\psi_0\rangle$

# Holographic Complexity:

- UV divergences appear as local integrals of geometric invariants

$$\mathcal{C}(\Sigma) \simeq c_0 \mathcal{V}(\Sigma) / \delta^{d-1} + \dots$$

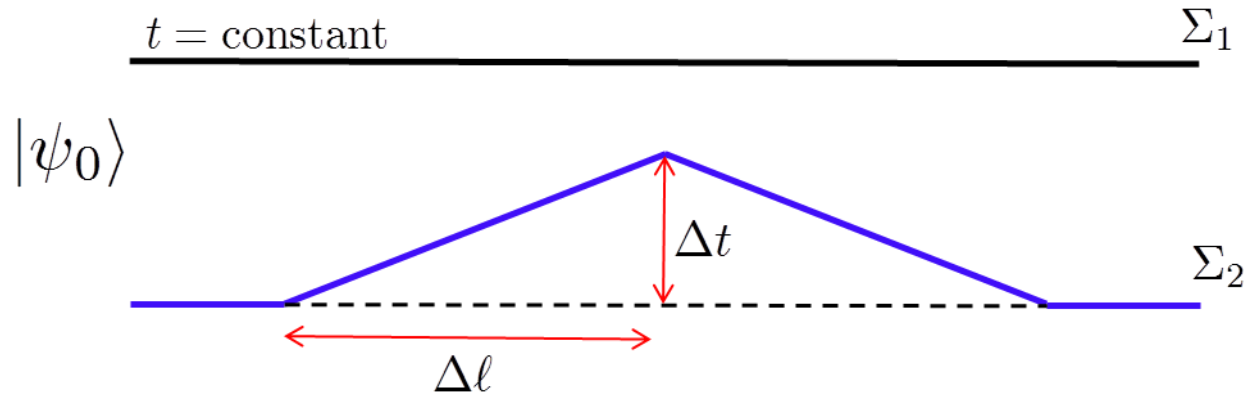


$$\Delta\mathcal{C} \simeq 2c_0 (\sqrt{\Delta\ell^2 - \Delta t^2} - \Delta\ell) \mathcal{V}_{\text{trans}} / \delta^{d-1} + \dots < 0$$

# Holographic Complexity:

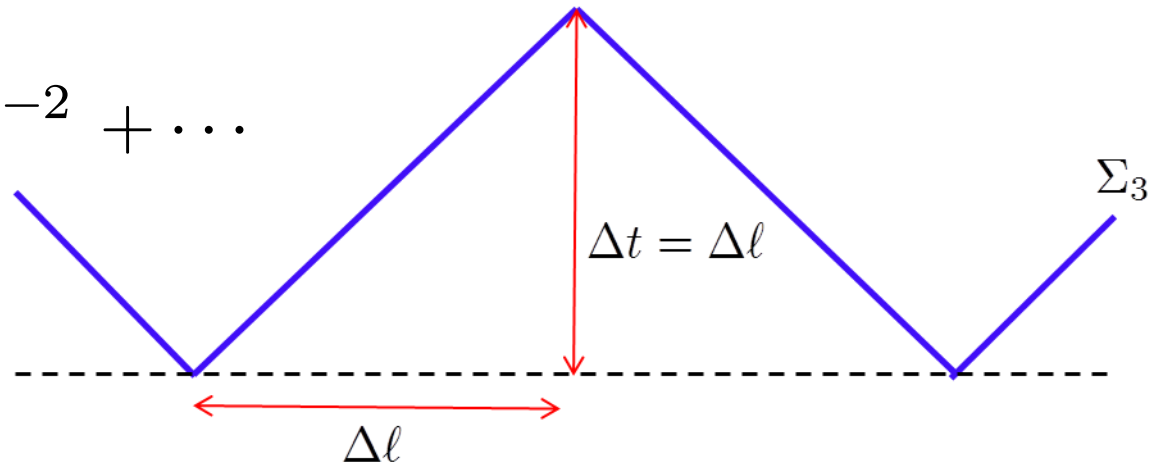
- UV divergences appear as local integrals of geometric invariants

$$\mathcal{C}(\Sigma) \simeq c_0 \mathcal{V}(\Sigma) / \delta^{d-1} + \dots$$



$$\Delta \mathcal{C} \simeq 2c_0 (\sqrt{\Delta \ell^2 - \Delta t^2} - \Delta \ell) \mathcal{V}_{\text{trans}} / \delta^{d-1} + \dots < 0$$

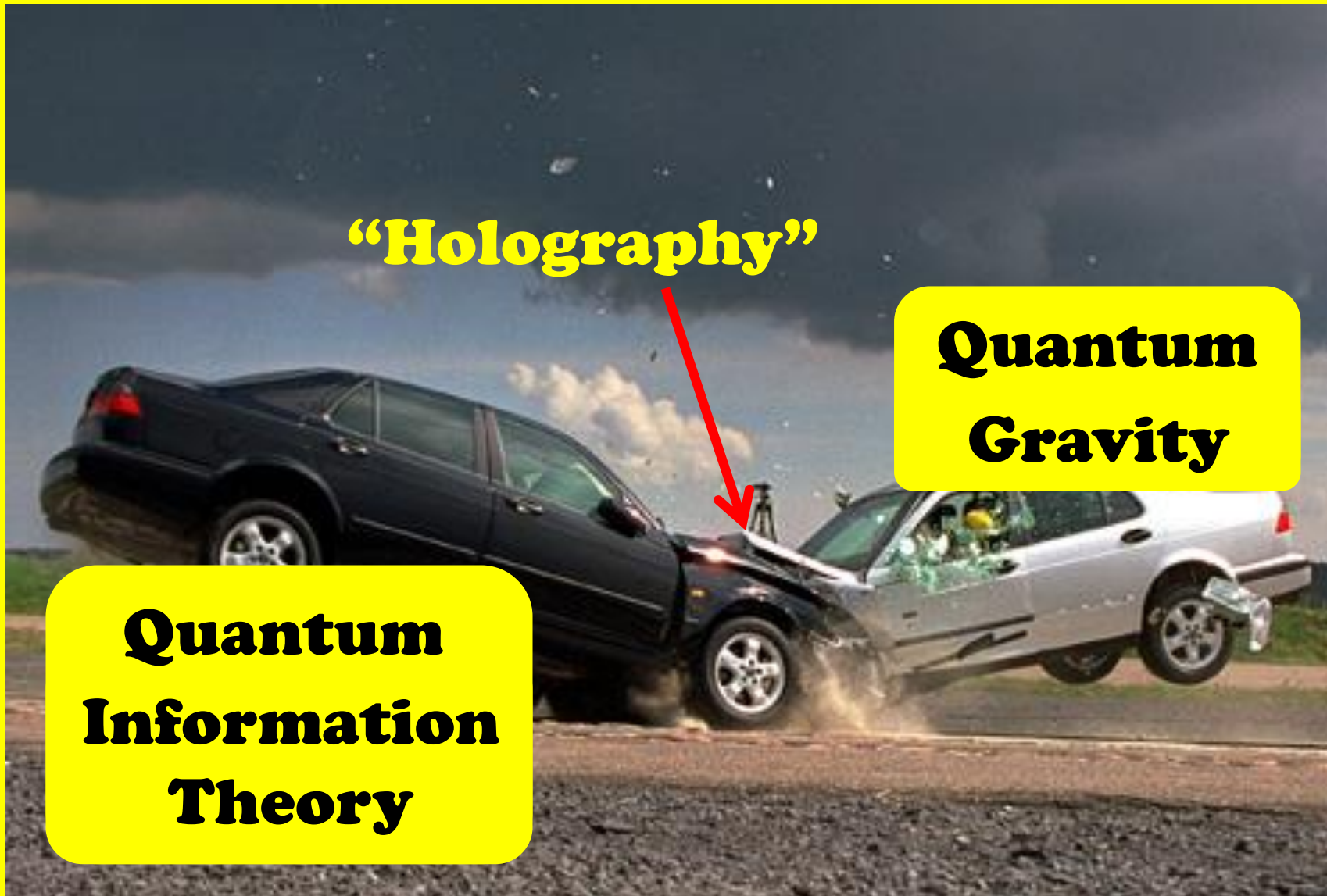
$$\mathcal{C}(\Sigma_3) \simeq c_1 \mathcal{V}_{\text{trans}} / \delta^{d-2} + \dots$$



# Questions?

- What is “holographic complexity”?
  - QFT/path integral description of “complexity” in boundary CFT?
  - what is boundary dual of these gravitational observables?
- is there a privileged role for (states on) null Cauchy surfaces?
  - provide distinguished reference states?
- is there a “renormalized holographic complexity”?
  - what’s it good for?; (EE vs mutual information versions of F)
- ambiguities? ambiguities? ambiguities?
  - connections between ambiguities in gravity and boundary?
- more boundary terms: higher codim. intersections; “complex” joint contributions; boundary “counterterms”
- why is complexity of formation positive?
- $\mathcal{C}_A$  contribution of spacetime singularity? • subregion complexity?

# **“It from Qubit”**: New Collision of Ideas



**“Holography”**

**Quantum  
Gravity**

**Quantum  
Information  
Theory**

# **“It from Qubit”: New Collision of Ideas**



<http://www.perimeterinstitute.ca/it-qubit-summer-school/it-qubit-summer-school-resources>

**Quantum  
Information  
Theory**

